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NuFact 2003

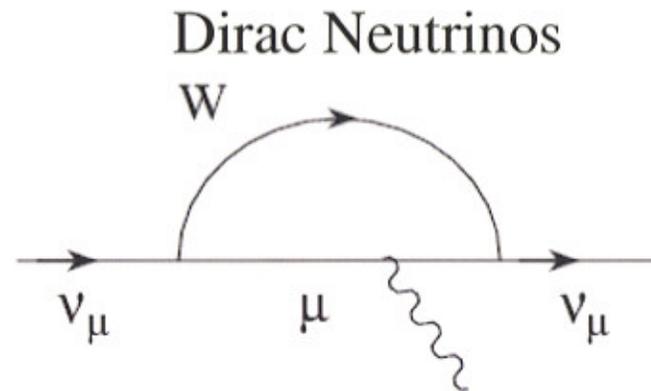
Neutrino Magnetic Moments:  
Measurements and predictions from the Standard  
Model and astrophysics

- Neutrino magnetic moments in beyond the Standard Model theories
- Predictions from Astrophysics
- Limits and upcoming measurements
- Next generation measurements

How does a neutrino magnetic moment (NMM) arise?

Minimally extended Standard Model:

Massive neutrinos can have a magnetic moment  
via one-loop radiative corrections



Neutrinos of mass  $m_\nu$ :

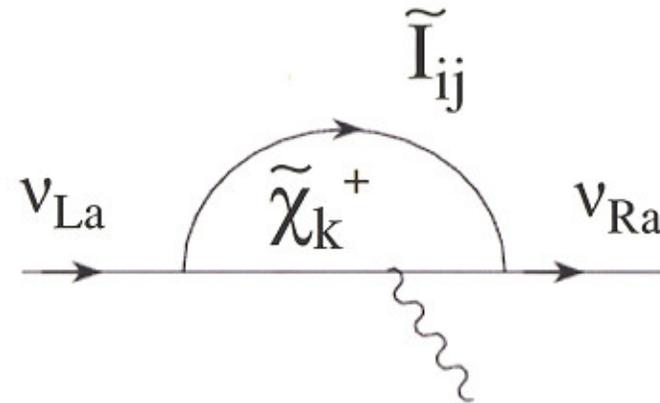
$$\mu_\nu = \frac{3eG_F}{8\sqrt{2}\pi^2} m_\nu \sim 3 \times 10^{-19} \mu_B \frac{m_\nu}{1\text{eV}}$$

## SUSY models

$$\mu_{\nu_e} \cong 10^{-15} - 10^{-16} \mu_B$$

$$\mu_{\nu_\mu} \cong 10^{-12} - 10^{-13} \mu_B$$

$$\mu_{\nu_\tau} \cong 10^{-12} \mu_B$$



## Large Extra Dimensions

$$\mu_{\nu} \cong 10^{-11} \mu_B$$

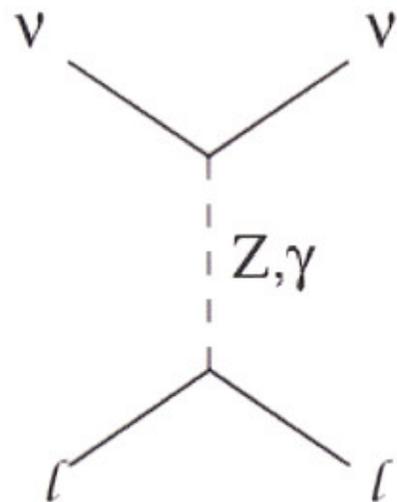
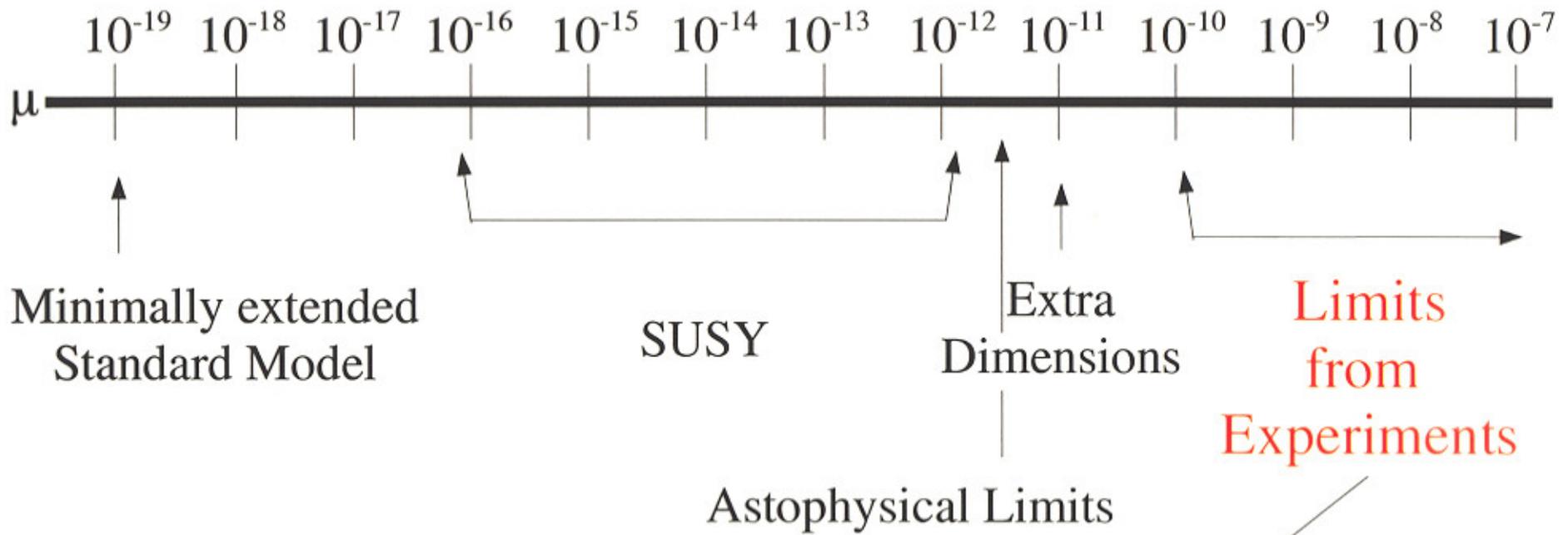
Neutrino magnetic moments  $\rightarrow$  indicator of which  
beyond the Standard model physics

→ *Astrophysical limits*

- slow rate of plasmon decay in horizontal branching stars  
→  $\mu_\nu < 10^{-11} \mu_B$
- neutrino energy loss rate from supernova 1987a  
→  $\mu_\nu < 10^{-12} \mu_B$

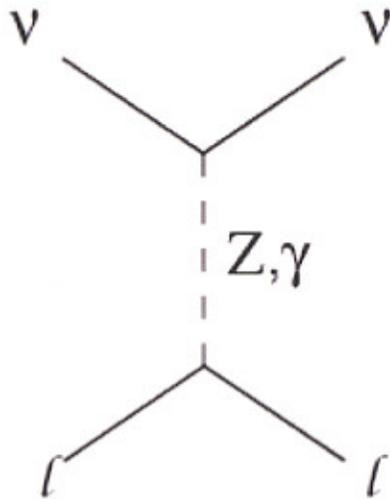
★ model assumptions made in these cases

Neutrino magnetic moments → indicator of what  
astrophysics



Experimental limits set by  
 measuring  
 $\nu e$  elastic scattering

Can look for excess of  $\nu e$  elastic cross section.....



$$\sigma_{EM} = \frac{\pi\alpha^2\mu_\nu^2}{m_e^2} \left[ \frac{E_e}{E_\nu} - \ln \frac{E_e}{E_\nu} - 1 \right]$$

sensitivity depends on

$$\left(\frac{\Delta\sigma}{\sigma}\right)^2 = \left(\frac{\Delta N}{N}\right)^2 + \left(\frac{\Delta F}{F}\right)^2$$

flux error a big issue!

increase in  
overall cross section

$$\sigma_{tot} = \sigma_{weak} + \sigma_{EM}$$

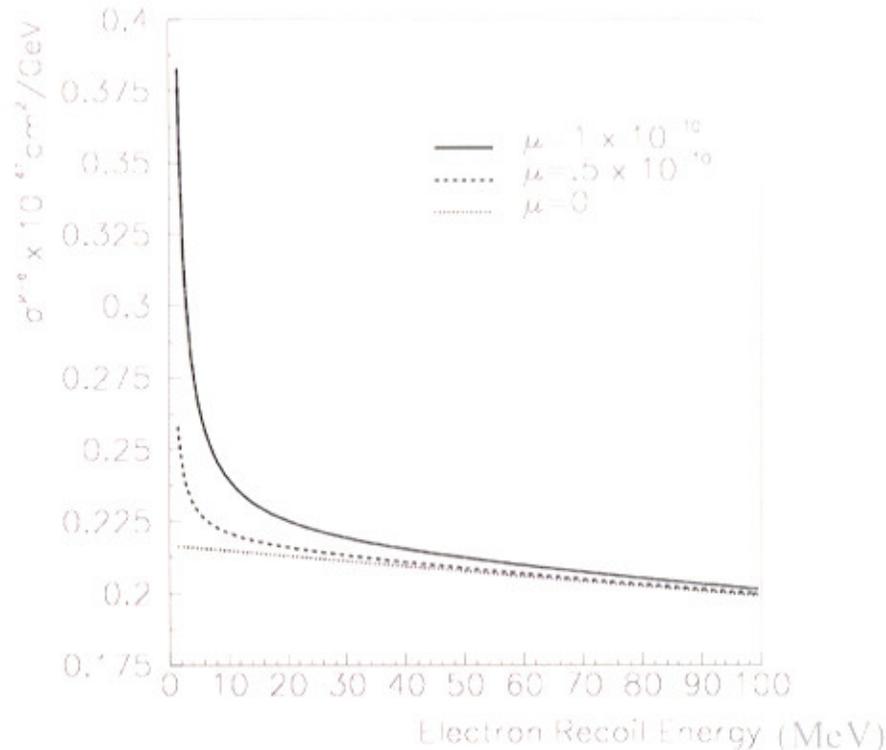
## Avoid flux error by measuring differential cross section!

$$\frac{d\sigma^{weak}}{dT} = \frac{2m_e G_F^2}{\pi} \left[ g_L^2 + g_R^2 \left( 1 - \frac{T}{E_\nu} \right)^2 - \frac{m_e}{E_\nu} g_R g_L \frac{T}{E_\nu} \right]$$

where  $T$  = electron recoil energy  
 $E_\nu$  = Neutrino energy

$$\frac{d\sigma^{EM}}{dT} = \frac{\pi \alpha^2 \mu_\nu^2}{m_e^2} \left( \frac{1}{T} - \frac{1}{E_\nu} \right)$$

Weak and EM Contributions to the  $\nu$ -e Cross Section



shape change in the  
differential cross section  
no flux error!

## Limits set from previous experiments

### → Electron neutrino magnetic moment

$$\mu_{\nu_e} \leq 1.5-1.8 \times 10^{-10} \mu_B$$

- Super K data: shape fit to recoil electron spectrum
- reactor experiments: combined measurements

Future experiments plan to measure  $\mu_{\nu_e}$  to  $\sim 10^{-12}$

tritium  $\nu$  experiments well understood energy spectrum endpoint

### → Muon n magnetic moment

$$\mu_{\nu_\mu} \leq 6.8 \times 10^{-10} \mu_B$$

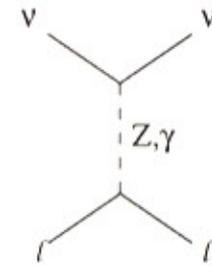
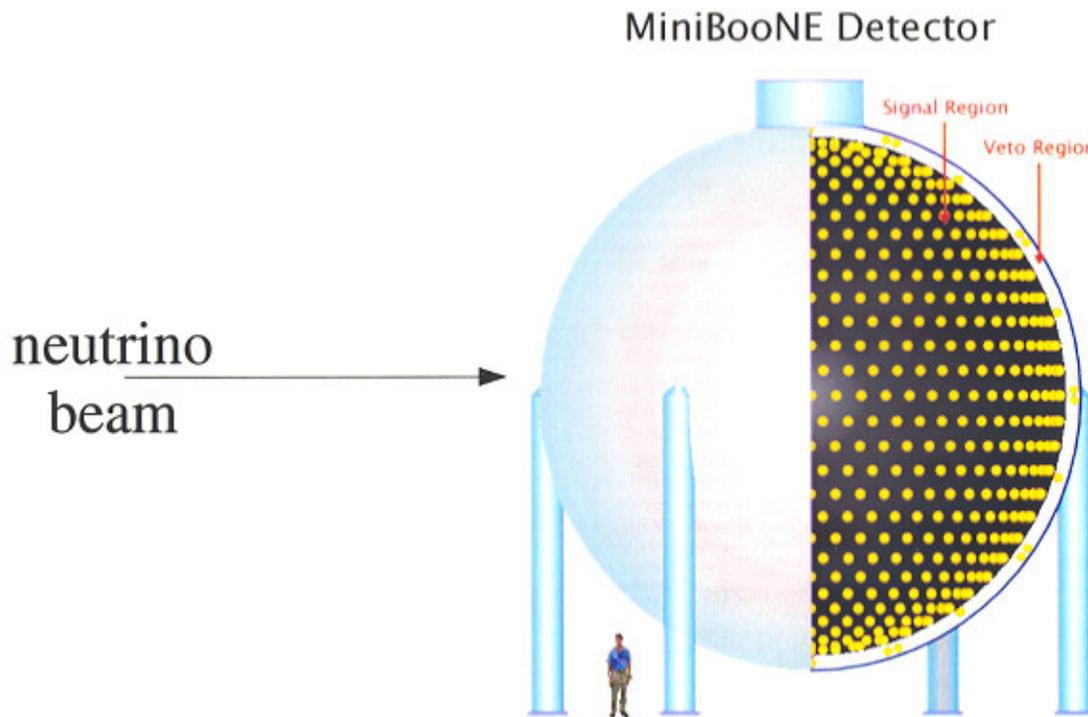
- LSND experiment: combine measurement for electron and muon neutrino magnetic moments using total  $\nu_{e,\mu} - e$  elastic cross section

### → Tau neutrino magnetic moment

$$\mu_{\nu_\tau} \leq 5.4 \times 10^{-7} \mu_B$$

## Experimental challenges and signatures:

- measuring electron recoil  
the lower the better!
- high statistics  
ve elastic cross section is small



*clean signature!*



ve elastic scatters  
are very forward

easy to separate  
from most  
backgrounds



- beam timing cuts
- forward cone cut
- $^{12}\text{B}$  -- energy cut

MiniBooNE's sensitivity depends on threshold for detection of recoil electrons and number of  $\nu_e$  elastic scatters

for 100  $\nu_e$  elastic scatters with electron recoils from 10-1000 MeV

Determining  
Sensitivity



throw fake data against  
the weak cross section

fit to weak + EM

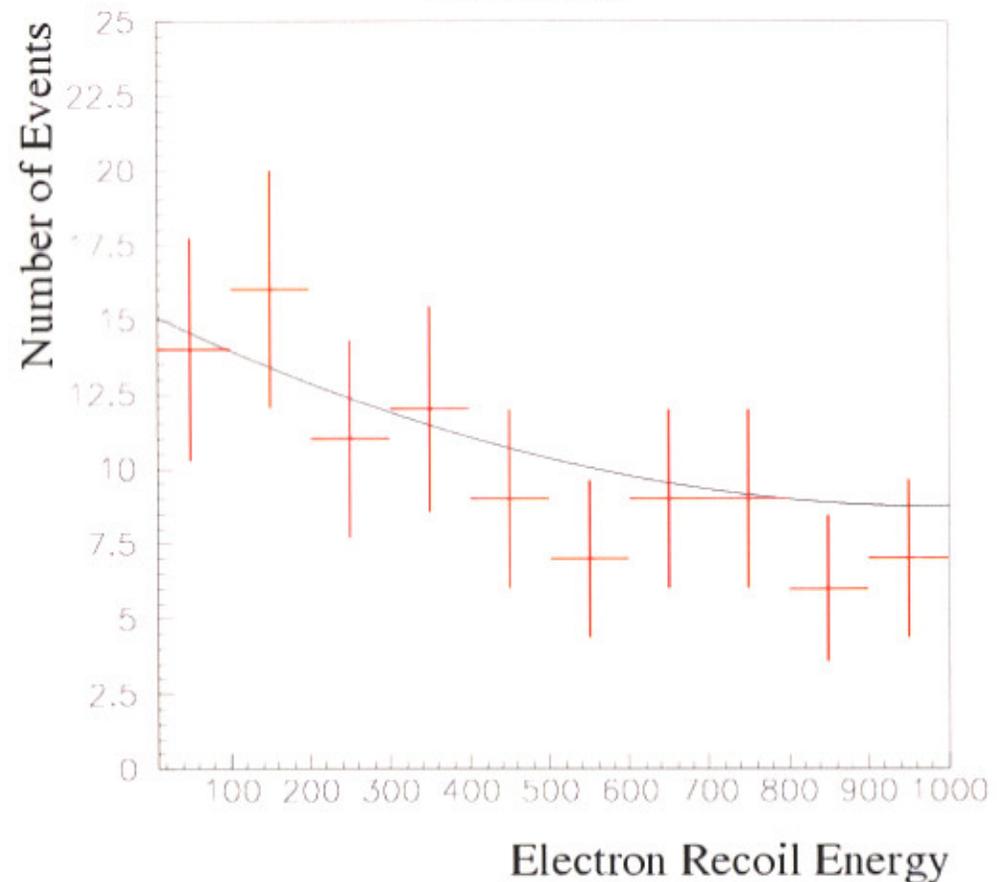
fit parameters:

magnetic moment,  $\mu$

Normalization -- flux

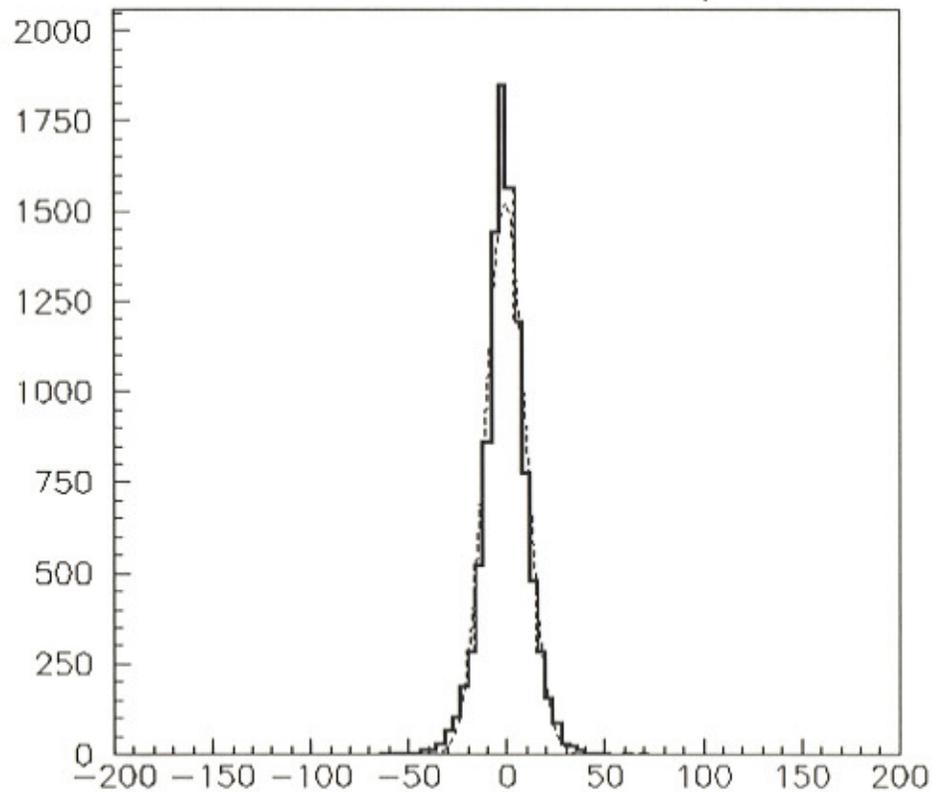
repeat many times!

Fake Data



# MiniBooNE sensitivity for $5 \times 10^{20}$ pot

Distribution of  $\mu^2 s$



fitted values of  $\mu_{\nu\mu}$   
for 10,000 fake data  
experiments



sensitivity:  
 $\mu_{\nu\mu} = 3.2 \times 10^{-10} \mu_B$

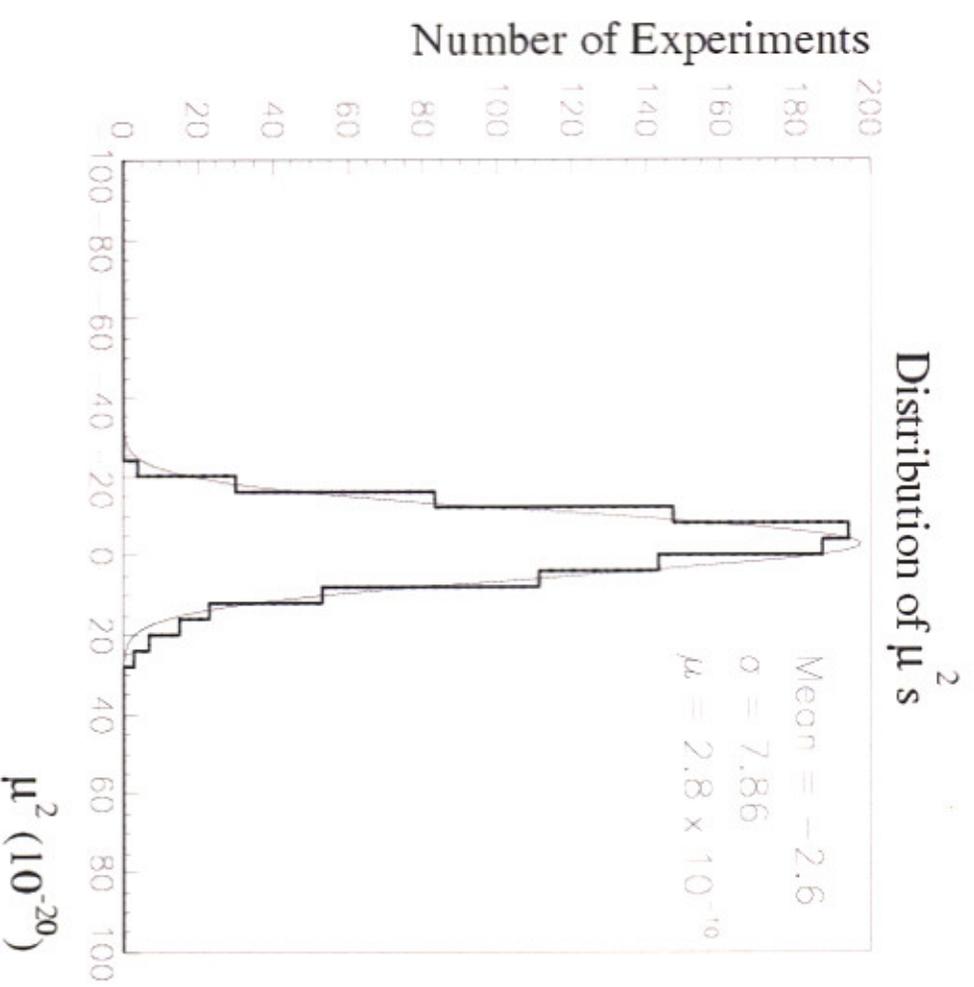
## Can we improve on this?

- lower electron recoil  
finely segmented detectors  
liquid argon TPCs  
→ few GeV
- higher statistics  
neutrino superbeams  
neutrino factories



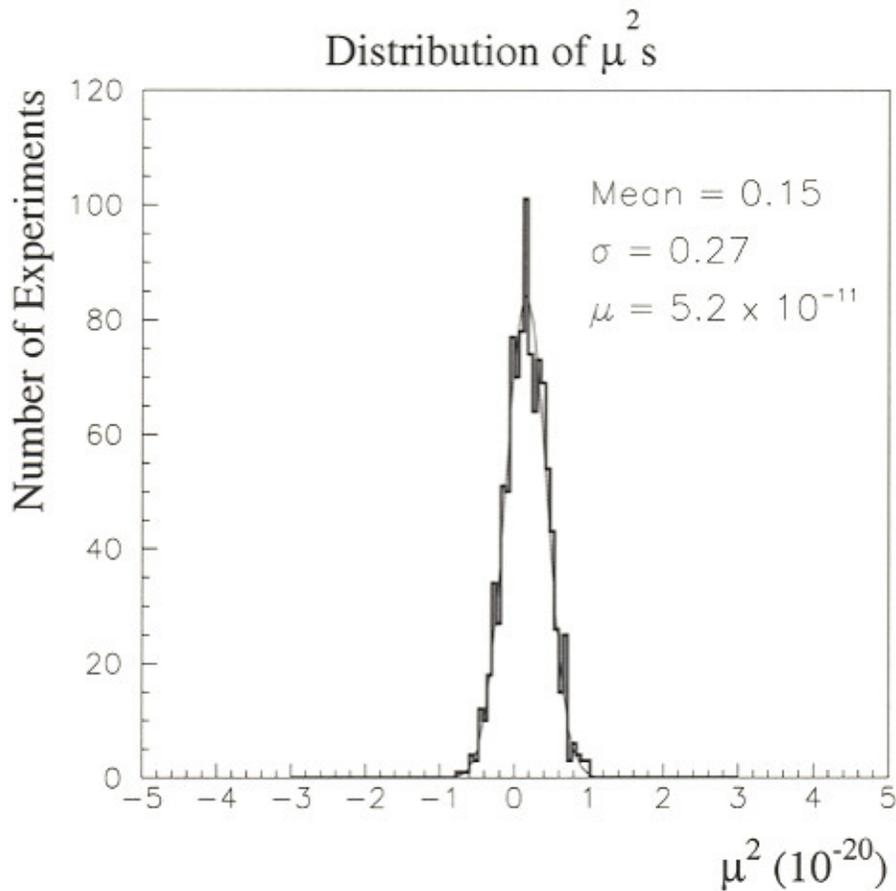
$\nu_{\mu}s$  and  $\nu_e s$

# MiniBooNE upgraded to reduce electron recoil to 1 MeV with 300 events



## at a superbeam facility<sup>1</sup>

15000  $\nu - e$  events with  $E_\nu = 1$  GeV  
 $1 \text{ MeV} < T_e < 1000 \text{ MeV}$

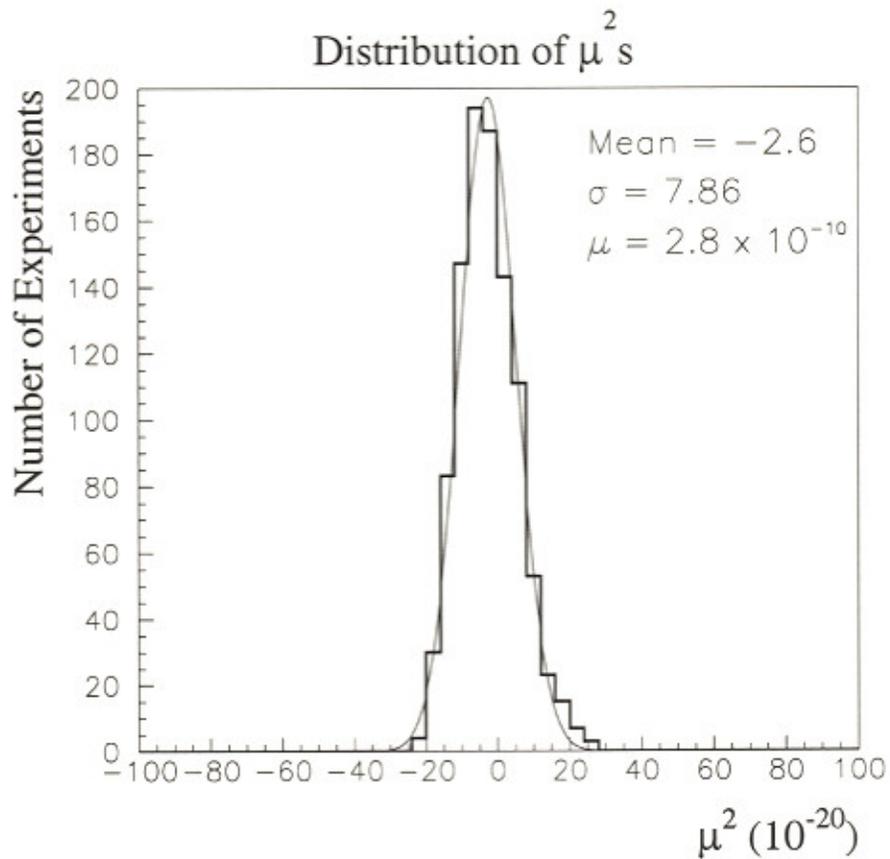


Sensitivity:  $\mu_{\nu\mu} < 5.2 \times 10^{-11} \mu_B$

<sup>1</sup>C. Ankenbrandt et. al., "Physics study group report on physics potential at FNAL with stronger proton sources," in preparation, <http://projects.fnal.gov/protondriver/>

# Neutrino magnetic moment measurements at an upgrade to MiniBooNE

300  $\nu - e$  events with  $E_\nu = 1$  GeV  
 $1 \text{ MeV} < T_e < 1000 \text{ MeV}$



Sensitivity:  $\mu_{\nu\mu} < 2.8 \times 10^{-10} \mu_B$

## Conclusions

- Neutrino magnetic moments can point us toward
  - which beyond the SM physics
  - what astrophysics
- With neutrino superbeams and high precision detectors -- into interesting region!