

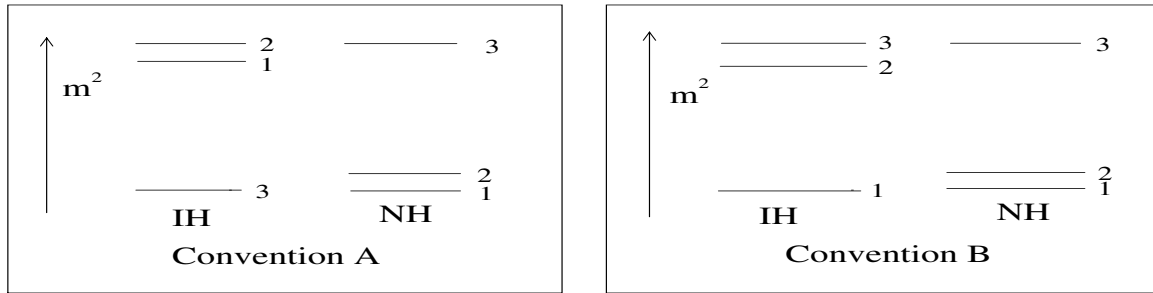
# Precision Measurement of Oscillation Parameters with Reactors

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# Three-Generation Survival Probability

- Two possible neutrino mass hierarchies



$$P_{NH}(\bar{\nu}_e \rightarrow \bar{\nu}_e)$$

$$= 1 - 2 \sin^2 \theta \cos^2 \theta \left( 1 - \cos \frac{\Delta m_{\text{atm}}^2 L}{2 E_\nu} \right)$$

$$- \frac{1}{2} \cos^4 \theta \sin^2 2\theta_\odot \left( 1 - \cos \frac{\Delta m_\odot^2 L}{2 E_\nu} \right)$$

$$+ 2 \sin^2 \theta \cos^2 \theta \sin^2 \theta_\odot \left( \cos \left( \frac{\Delta m_{\text{atm}}^2 L}{2 E_\nu} - \frac{\Delta m_\odot^2 L}{2 E_\nu} \right) - \cos \frac{\Delta m_{\text{atm}}^2 L}{2 E_\nu} \right)$$

$$P_{IH}(\bar{\nu}_e \rightarrow \bar{\nu}_e)$$

$$= 1 - 2 \sin^2 \theta \cos^2 \theta \left( 1 - \cos \frac{\Delta m_{\text{atm}}^2 L}{2 E_\nu} \right)$$

$$- \frac{1}{2} \cos^4 \theta \sin^2 2\theta_\odot \left( 1 - \cos \frac{\Delta m_\odot^2 L}{2 E_\nu} \right)$$

$$+ 2 \sin^2 \theta \cos^2 \theta \cos^2 \theta_\odot \left( \cos \left( \frac{\Delta m_{\text{atm}}^2 L}{2 E_\nu} - \frac{\Delta m_\odot^2 L}{2 E_\nu} \right) - \cos \frac{\Delta m_{\text{atm}}^2 L}{2 E_\nu} \right)$$

- Depends on **5 parameters**: 4 continuous oscillation parameters  $\Delta m_\odot^2$ ,  $\sin^2 \theta_\odot$ ,  $\Delta m_{\text{atm}}^2$ ,  $\sin^2 \theta$  and the **hierarchy**.
- Does **not** depend on  $\theta_{\text{atm}}$  and  $\delta$ .

# The Baseline

- For Reactors,  $E_\nu \sim 3.6$  MeV,  $L^* \equiv \frac{2\pi E_\nu}{\Delta m^2}$
- For  $\Delta m^2 \equiv \Delta m_{\text{atm}}^2 \sim 2.5 \times 10^{-3}$  eV<sup>2</sup>,  $L^* \equiv 1.1 - 1.7$  km

$$P_{ee} \approx 1 - \sin^2 2\theta \sin^2 \frac{\Delta m_{\text{atm}}^2 L}{4E_\nu} - \cos^4 \theta \sin^2 2\theta_\odot \sin^2 \frac{\Delta m_\odot^2 L}{4E_\nu}$$

## Short Baseline Experiment

CHOOZ, Palo Verde, Krasnoyarsk....  
Minakata, et al., hep-ph/0211111  
Huber, et al., hep-ph/0303232

- For  $\Delta m^2 \equiv \Delta m_\odot^2 \sim 7.2 \times 10^{-5}$  eV<sup>2</sup>,  $L^* \equiv 50 - 70$  km

$$P_{ee} \approx 1 - \cos^4 \theta \sin^2 2\theta_\odot \sin^2 \frac{\Delta m_\odot^2 L}{4E_\nu} - \frac{1}{2} \sin^2 2\theta$$

## Long Baseline Experiment

KamLAND, Eguchi et al., PRL 90,021802,2003  
Bandyopadhyay, S.C., Goswami, hep-ph/0302243 (PRD)

- For  $\Delta m^2 \equiv \Delta m_\odot^2 \sim 1.5 \times 10^{-4}$  eV<sup>2</sup>,  $L^* \equiv 20 - 30$  km

$$P_{ee} \approx 1 - \cos^4 \theta \sin^2 2\theta_\odot \sin^2 \frac{\Delta m_\odot^2 L}{4E_\nu} - \frac{1}{2} \sin^2 2\theta \sin^2 \frac{\Delta m_{\text{atm}}^2 L}{4E_\nu} \\ + 2 \sin^2 \theta \cos^2 \theta \sin^2 \theta_\odot \left( \cos \left( \frac{\Delta m_{\text{atm}}^2 L}{2E_\nu} - \frac{\Delta m_\odot^2 L}{2E_\nu} \right) - \cos \frac{\Delta m_{\text{atm}}^2 L}{2E_\nu} \right)$$

## Intermediate Baseline Experiment

Heilbronn, Schonert et al., hep-ex/0203013  
Petcov and Piai, PLB 533,94,2002  
S.C., Petcov and Piai, hep-ph/0306017

# Long Baseline Reactor Experiments

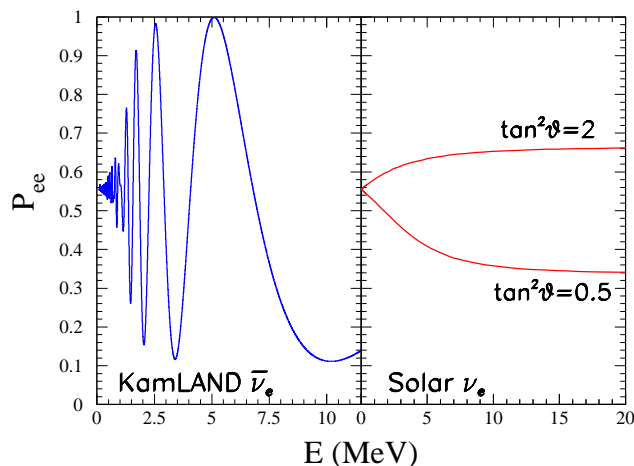
## KamLAND

- 1 kTon Liquid Scintillator detector situated at the old Kamioka site.
- Looks for  $\bar{\nu}_e$  oscillation coming from 16 reactors at distances 81 - 824 km.
- Most powerful reactors are at a distance  $\sim 160$  km.
- Detection process :  $\bar{\nu}_e p \rightarrow e^+ n$

- Survival Probability

$$P_i(\bar{\nu}_e \leftrightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 d_i}{E_\nu} \right)$$

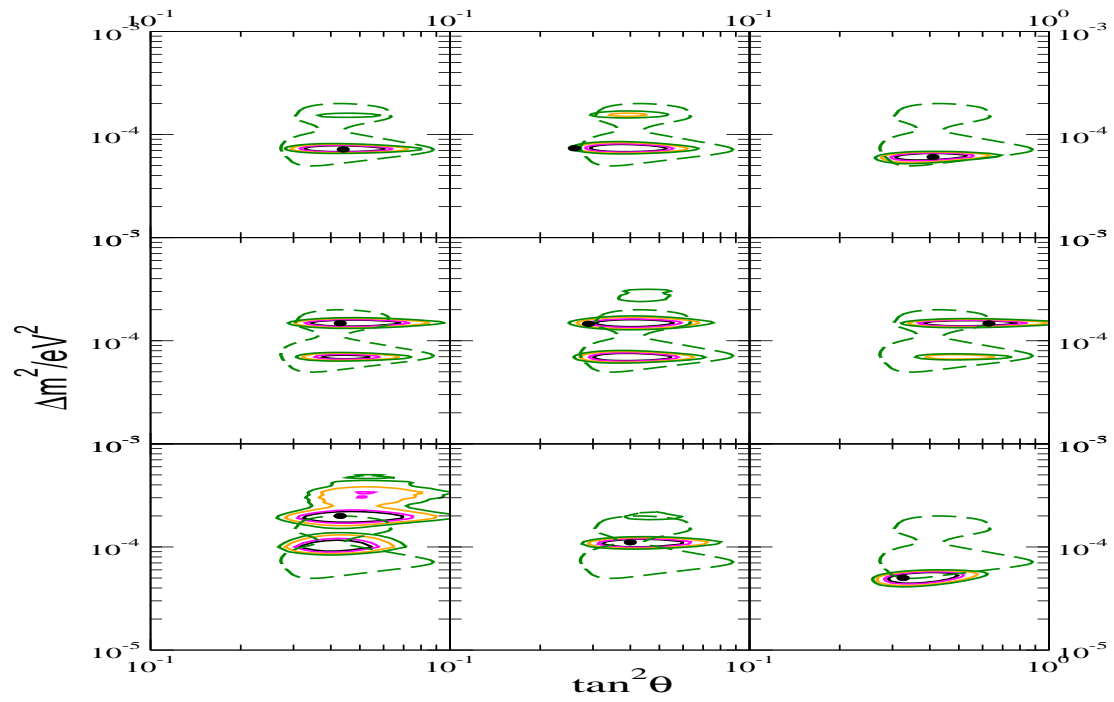
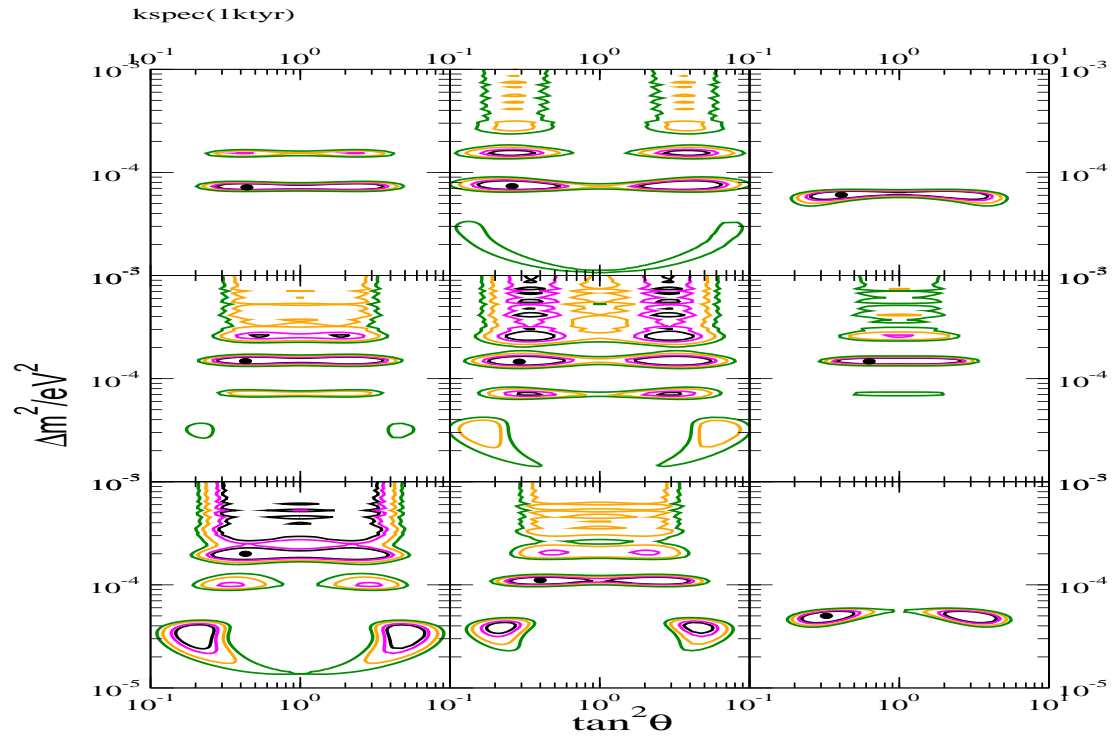
- $E_\nu \sim 3$  MeV,  $L \sim 1.8 \times 10^5$  m ,  $\Delta m^2 \sim 1.6 \times 10^{-5}$  eV<sup>2</sup>  
– sensitive to LMA region
- No matter effects:  $\theta \equiv \frac{\pi}{2} - \theta$



J.N.Bahcall et al., hep-ph/0212247

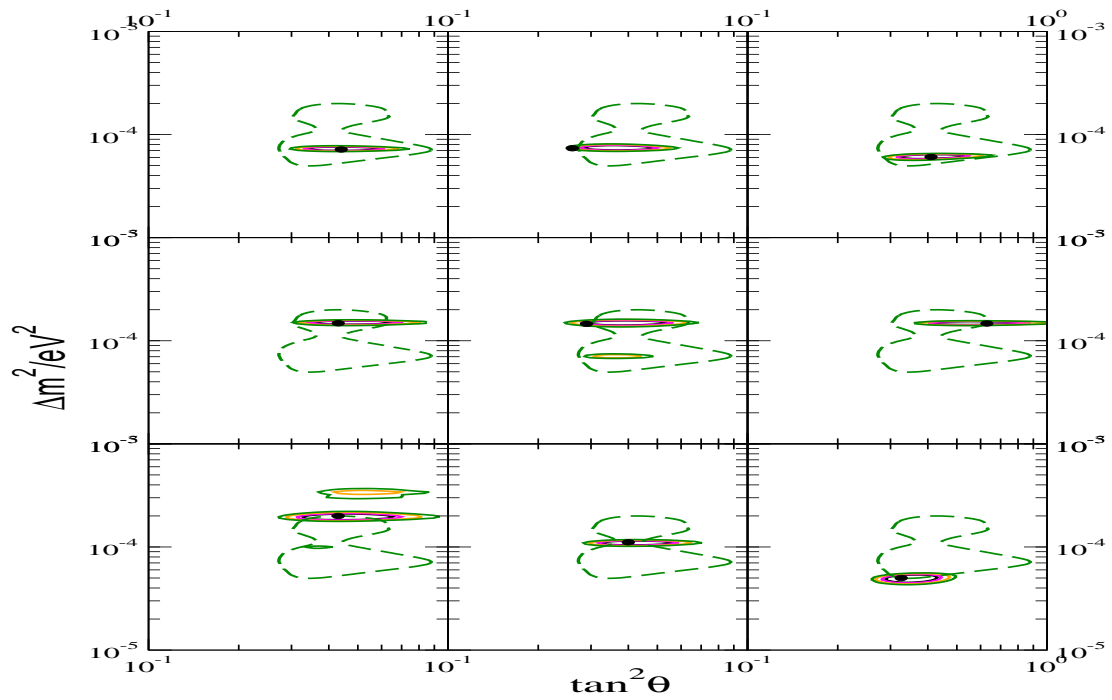
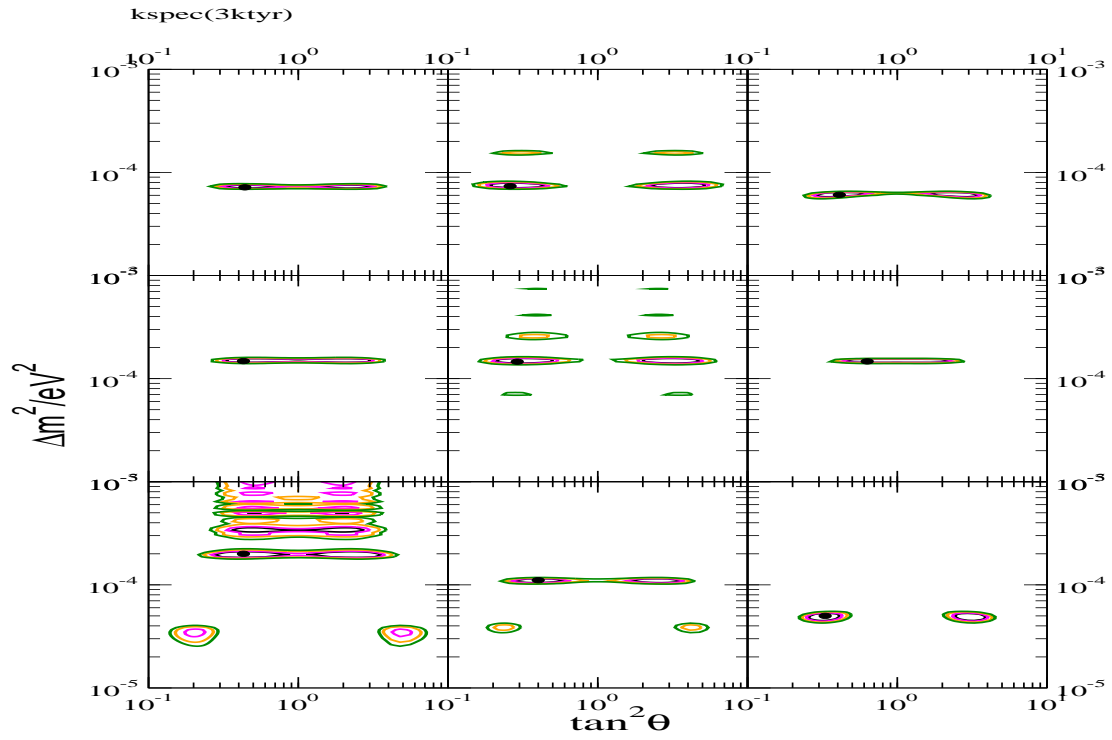
Can probe the L/E dependence of the oscillations in the LMA region —unprecedented sensitivity to  $\Delta m^2$

• KamLAND 1 kTy



A.Bandyopadhyay et al, hep-ph/0212146

• KamLAND 3 kTy



A.Bandyopadhyay et al, hep-ph/0212146

## Closer look at KamLAND sensitivity

- 5% systematic uncertainty for 1 kTy KamLAND
- 3% systematic uncertainty for 3 kTy KamLAND
- 7% total error for future SNO NC data

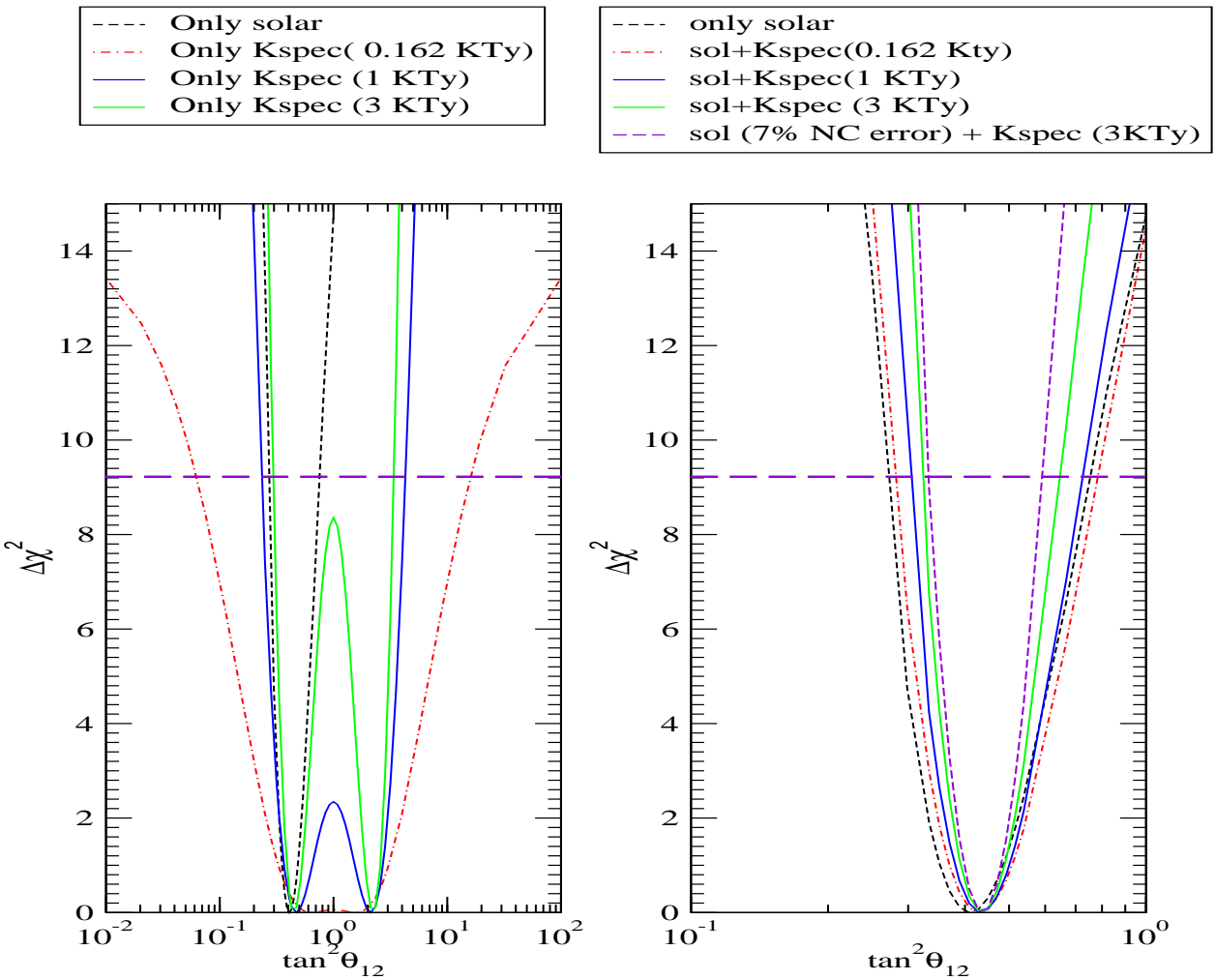
$$\text{spread} = \frac{a_{max} - a_{min}}{a_{max} + a_{min}} \times 100 \%$$

Data set used	99% CL range of $\Delta m_{21}^2 \times 10^{-5} \text{eV}^2$	99% CL spread of $\Delta m_{21}^2$	$1 \sigma$ range of $\tan^2 \theta_{12}$	$2\sigma$ range of $\tan^2 \theta_{12}$	99% CL range of $\tan^2 \theta_{12}$	$1 \sigma$ spread in $\tan^2 \theta_{12}$	$2 \sigma$ spread in $\tan^2 \theta_{12}$	99% CL spread in $\tan^2 \theta_{12}$
only sol	3.2 - 24.0	76%	.33 - .53	.29 - .66	.27 - .75	23%	39%	47%
sol+162 Ty	5.3 - 9.9	30%	.34 - .55	.30 - .68	.28 - .78	23%	39%	47%
sol+1 kTy	6.7 - 8.0	9%	.36 - .54	.33 - .65	.30 - .72	20%	33%	41%
sol+3 kTy	6.8 - 7.7	6%	.37 - .52	.34 - .59	.33 - .65	17%	27%	33%
sol(7%)+3 kTy	6.8 - 7.7	6%	.38 - .50	.35 - .56	.33 - .60	14%	23%	29%

Table 1: The range of parameter values allowed and the corresponding spread. For the current observed solar+KamLAND analysis we show the ranges and the spread only in the low-LMA region. For the 1 kTy and 3 kTy ranges we have simulated the spectrum at the current low-LMA best-fit. We assume 5% systematic error for 1 kTy KamLAND spectrum and 3% systematic error for 3 kTy KamLAND spectrum. The last row of the Table corresponds to a combination of the 3 kTy KamLAND data and the global solar neutrino data where the SNO NC error has been reduced to only 7%.

- Tremendous sensitivity to  $\Delta m_{21}^2$
- However sensitivity to  $\theta_{12}$  does not look good

A.Bandyopadhyay, S.C., S.Goswami, hep-ph/0302243



- KamLAND has a  $\theta_{12}$  and  $\pi/2 - \theta_{12}$  ambiguity
- Even with 3 kTy statistics, KamLAND fails to constrain  $\theta_{12}$  better than the current solar neutrino experiments
- Even with 3 kTy KamLAND can rule out maximal mixing only at the  $2\sigma$  level

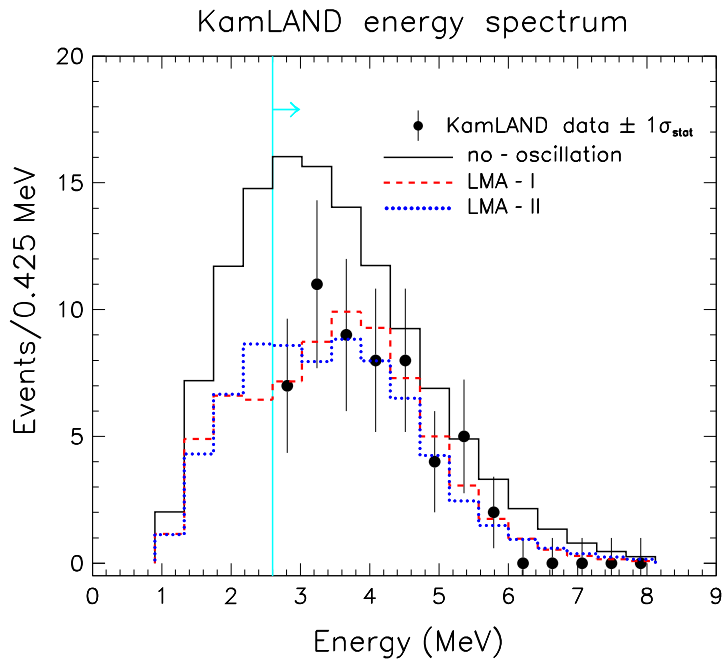
A.Bandyopadhyay, S.C., S.Goswami, hep-ph/0302243



## Still closer look at KamLAND

- The survival probability at KamLAND

$$P_{ee} = 1 - \sum_i \sin^2 2\theta_{12} \sin^2 \left( \frac{\Delta m_{21}^2 L_i}{4E} \right)$$



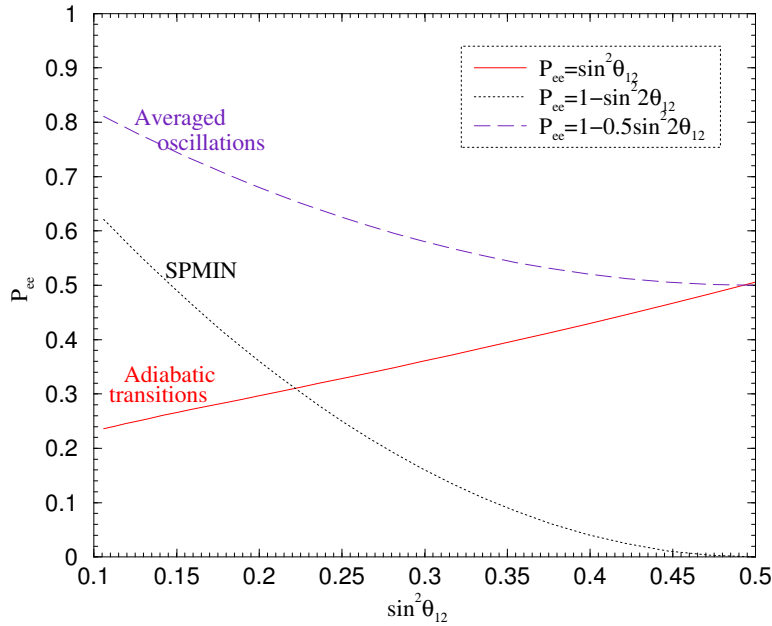
G.L. Fogli et al. hep-ph/0212127

- KamLAND has a **peak** in its survival probability

$$\Downarrow$$
$$\sin^2 \left( \frac{\Delta m_{21}^2 L_i}{4E} \right) \sim 0$$

- The  $\theta_{12}$  term gets smothered!

## Measurement of $\sin^2 \theta_{\odot}$



For  ${}^8B$  neutrinos undergoing matter enhanced resonance

$$P_{ee}^{AD} = \frac{1}{2} + \frac{1}{2}(1 - P_J) \cos 2\theta_m \cos 2\theta_{\odot}$$

$$\approx \sin^2 \theta_{\odot} \quad \text{LMA(AD)}$$

For  $pp$  neutrinos do not encounter resonance,

$$P_{ee}^{AV} = 1 - \frac{1}{2} \sin^2 2\theta_{\odot} \quad \text{Averaged Oscillations(AV)}$$

For  $\nu O$  such that  $\sin^2(\Delta m_{\odot}^2 L/E) = 1$

$$P_{ee}^{SPMIN} = 1 - \sin^2 2\theta_{\odot} \quad \text{Survival Probability MINima(SPMIN)}$$

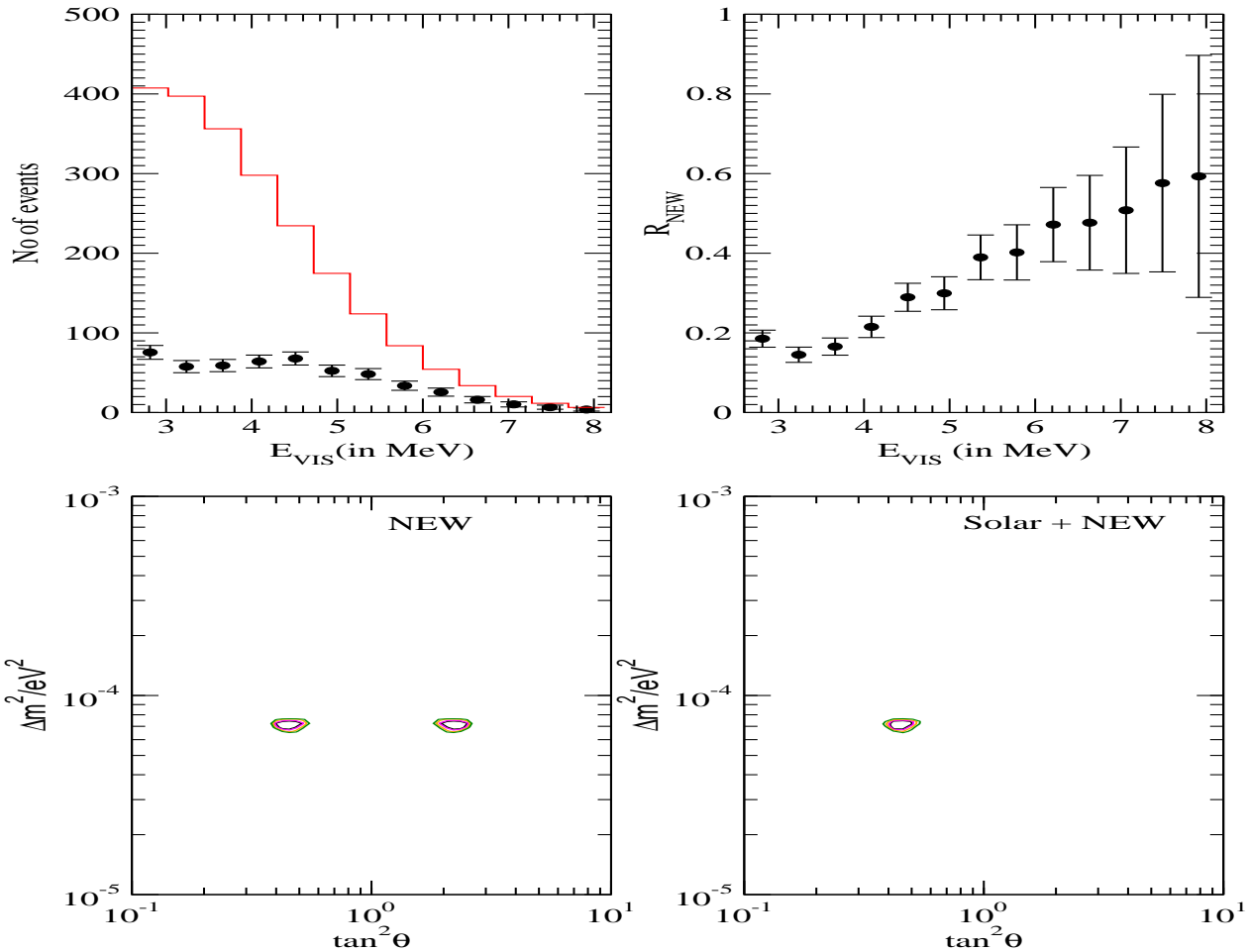
$$(\Delta \sin^2 \theta_{\odot})_{AD} \sim \Delta P_{ee}; \text{ good if } \sin^2 \theta_{\odot} \sim 0.5$$

$$(\Delta \sin^2 \theta_{\odot})_{AV} \sim \frac{\Delta P_{ee}}{-2 \cos 2\theta_{\odot}}; \text{ better if } \cos 2\theta_{\odot} \gtrsim 0.5 (\sin^2 \theta_{\odot} \lesssim 0.25)$$

$$(\Delta \sin^2 \theta_{\odot})_{spmin} \sim \frac{\Delta P_{ee}}{-4 \cos 2\theta_{\odot}}; \text{ best if } \cos 2\theta_{\odot} \gtrsim 0.25 (\sin^2 \theta_{\odot} \lesssim 0.375)$$

- Best-Fit:  $\sin^2 \theta_{\odot} = 0.3$ , Range:  $0.21 < \sin^2 \theta_{\odot} < 0.47$

# New Reactor Experiment for $\theta_{\odot}$



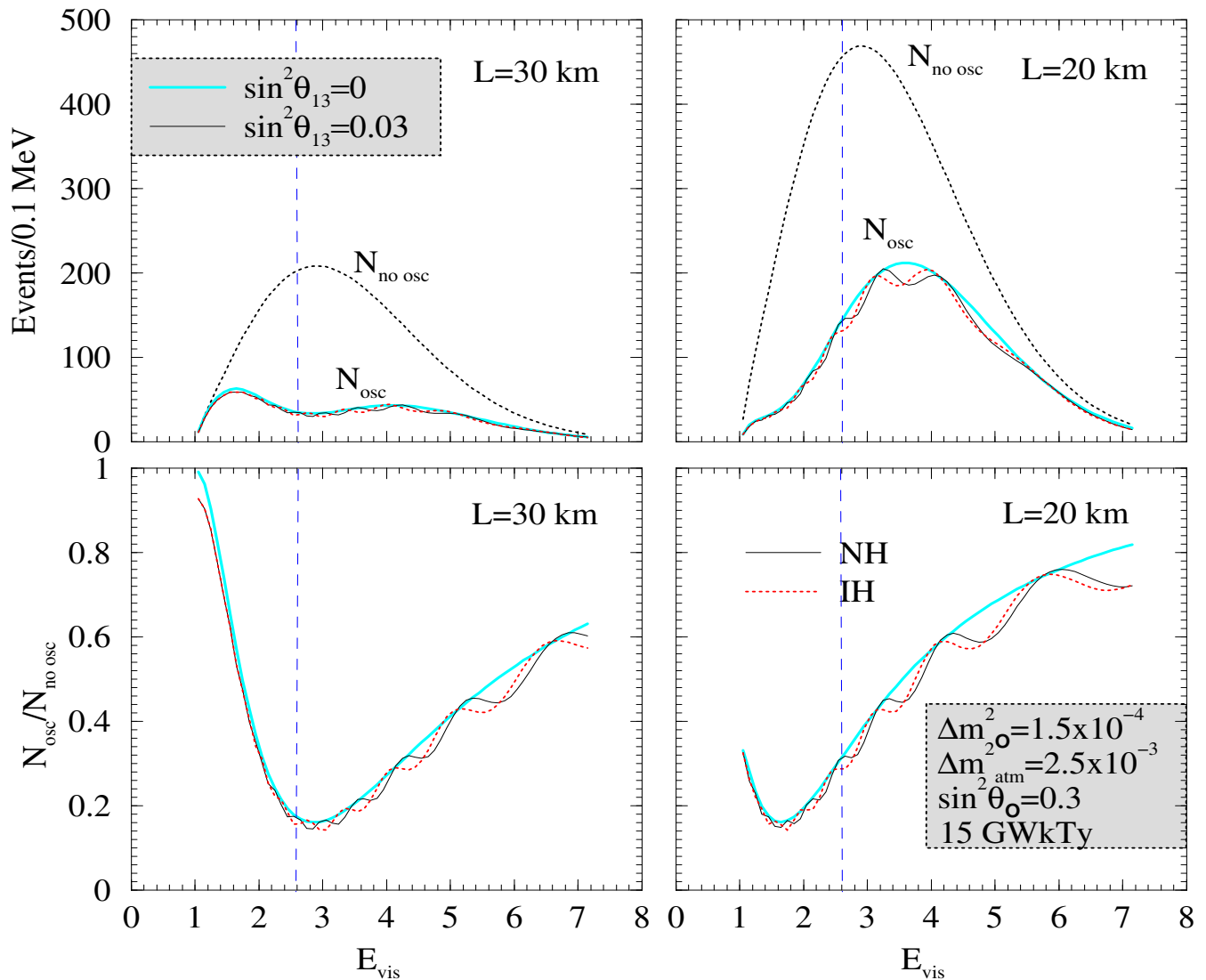
- Take a powerful reactor source: **Kashiwazaki**
- Take a baseline of  $L \sim 70$  km
- This corresponds to an **SPMIN** at  $\sim 3 - 4$  MeV
- Reduce systematic uncertainty to **2%**
- Spread in  $\tan^2 \theta_{\odot} (\sin^2 \theta_{\odot})$  at 99% **14% (9.6%)!!**

A. Bandyopadhyay, S.C., S. Goswami, hep-ph/0302243

- Check out 70 km baseline sites along the Rhone Valley

C. Bouchiat, hep-ph/0304253

# Intermediate Baseline Reactor Experiments



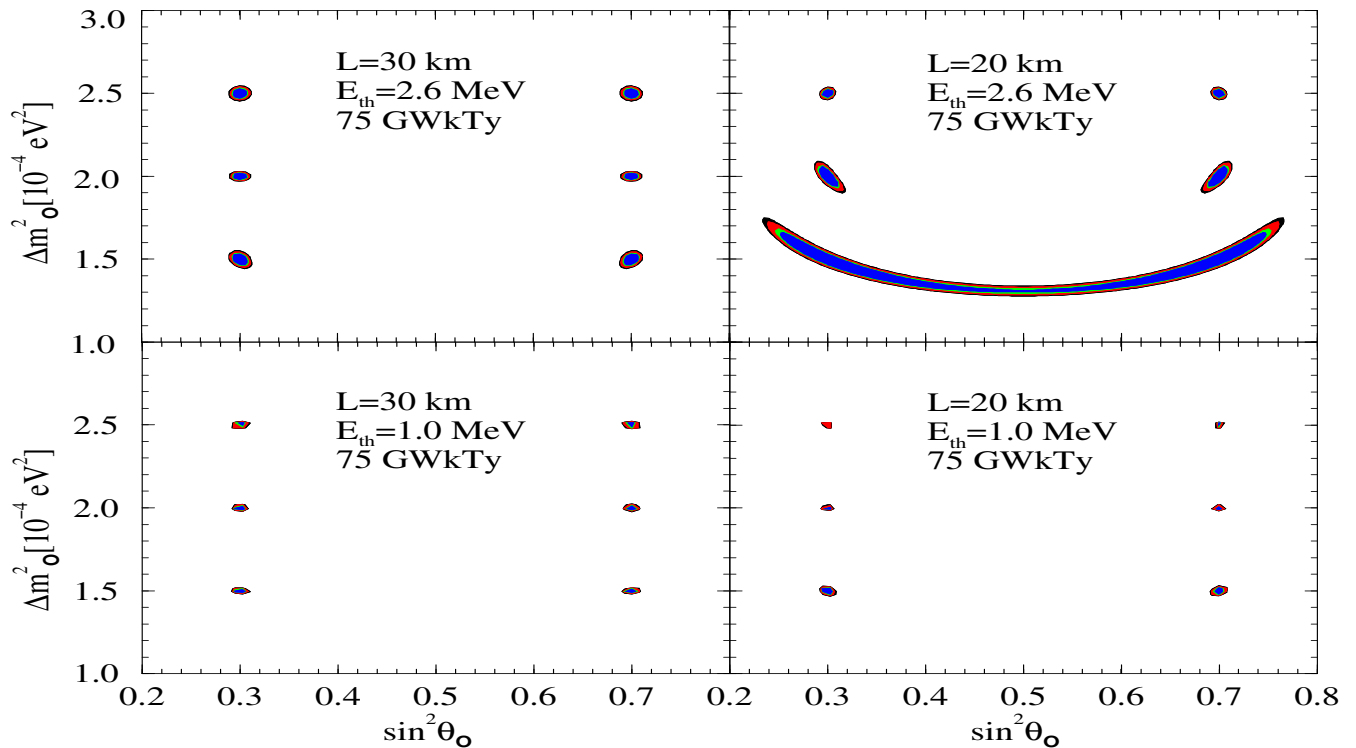
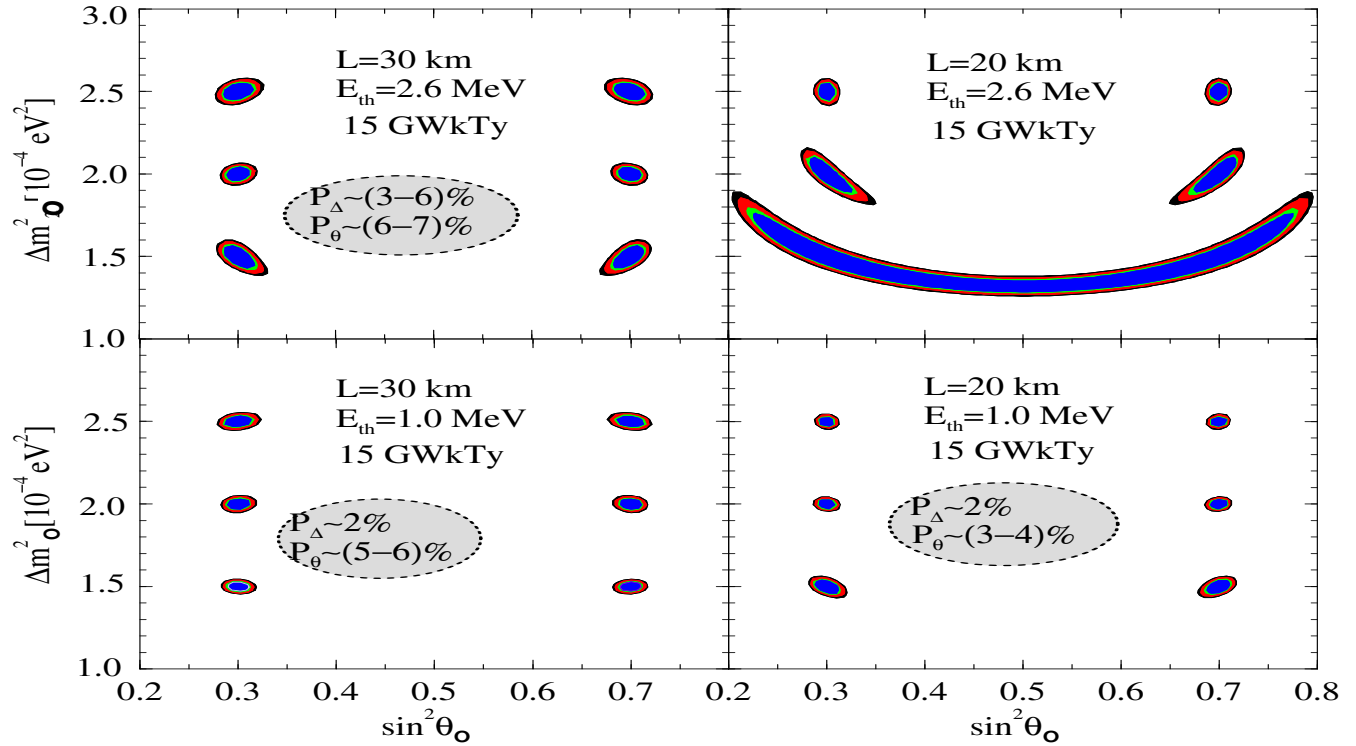
- Has the potential to measure **all the 5 parameters** if:

- ★ ★  $\Delta m_{\odot}^2$  is in **high-LMA**
- ★ ★  $E_{\text{th}}$  and  $L$  allow  $\Delta m_{\odot}^2$  driven **SPMIN**
- ★ ★ Energy **resolution** of the detector is good enough
- ★ ★ ★ **Statistics** are high enough
- ★ ★ ★ **Systematics** are low enough

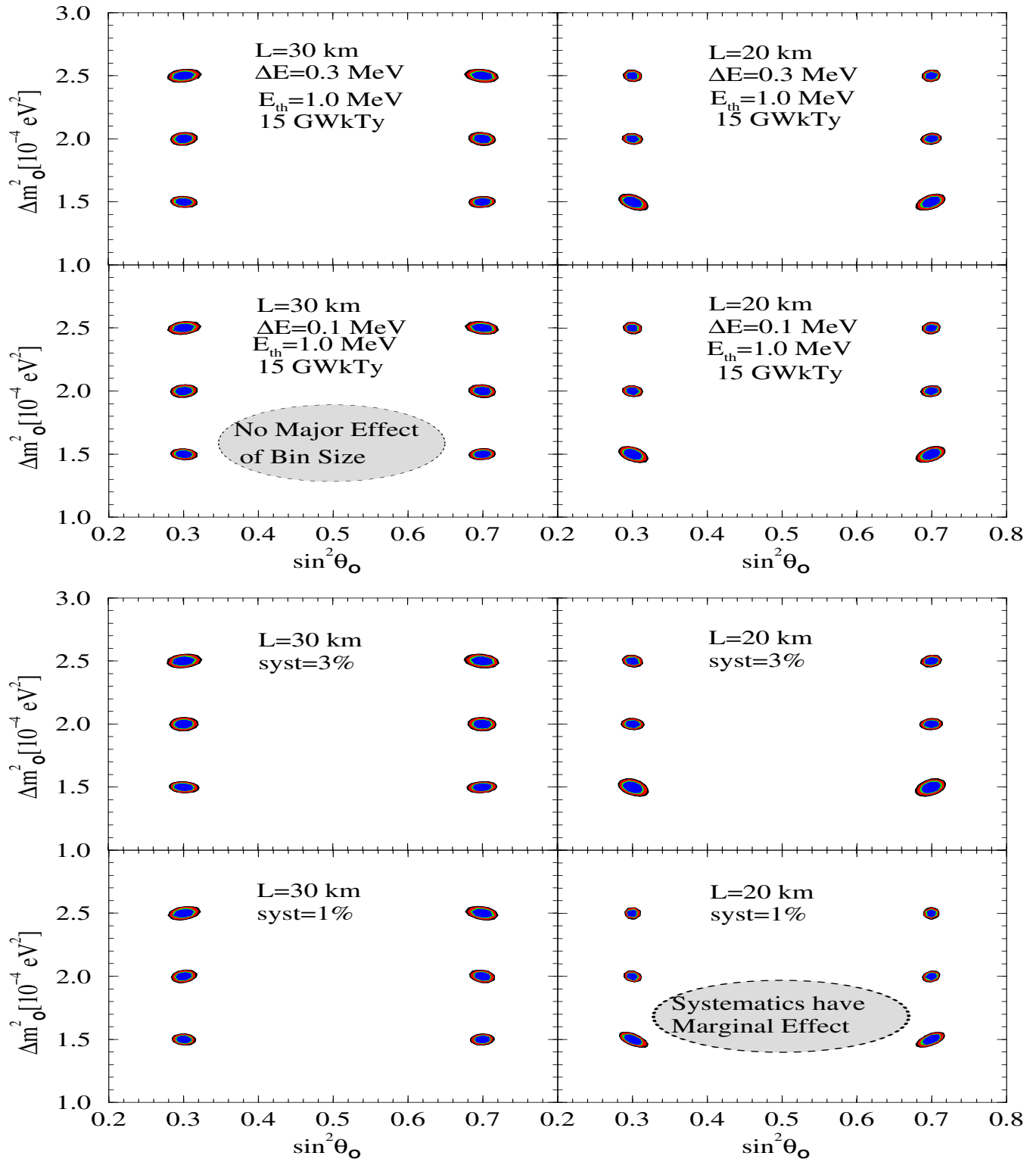
( $\Delta m_{\odot}^2$ ,  $\sin^2 \theta_{\odot}$ ,  $\sin^2 \theta$ ,  $\Delta m_{\text{atm}}^2$ , **Hierarchy**)

S.C., S.T.Petcov, M.Piai, hep-ph/0306017

# Precision Measurement of $\Delta m_{21}^2$ and $\sin^2 \theta_{12}$



# Precision Measurement of $\Delta m_{21}^2$ and $\sin^2 \theta_{12}$



# Effect of $\sin^2 \theta$ on Solar Parameters

$$P_{ee} \approx \cos^4 \theta \left( 1 - \sin^2 2\theta_{\odot} \sin^2 \frac{\Delta m_{\odot}^2 L}{4 E_{\nu}} \right)$$

- The factor  $\cos^4 \theta$  brings a  $\sim 10\%$  effect in  $P_{ee}$

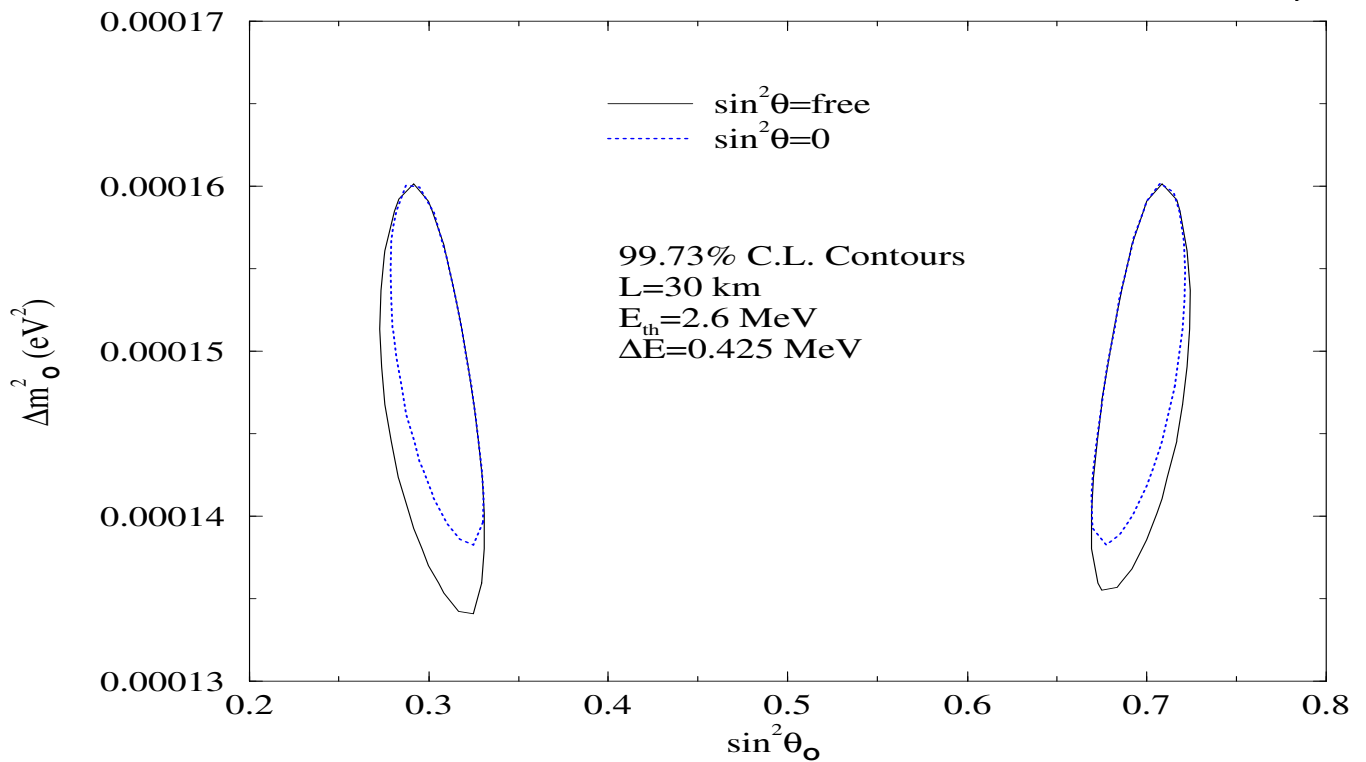
$$\delta(\sin^2 2\theta_{\odot}) \approx \frac{2\Delta P_{ee} \sin^2 \theta}{\sin^2 \frac{\Delta m_{\odot}^2 L}{4 E_{\nu}}} + 2 \frac{(1 - \sin^2 2\theta_{\odot} \sin^2 \frac{\Delta m_{\odot}^2 L}{4 E_{\nu}}) \Delta(\sin^2 \theta)}{\sin^2 \frac{\Delta m_{\odot}^2 L}{4 E_{\nu}}}$$



$\lesssim 0.01$



$\sim 0.017$  (for SPMIN)

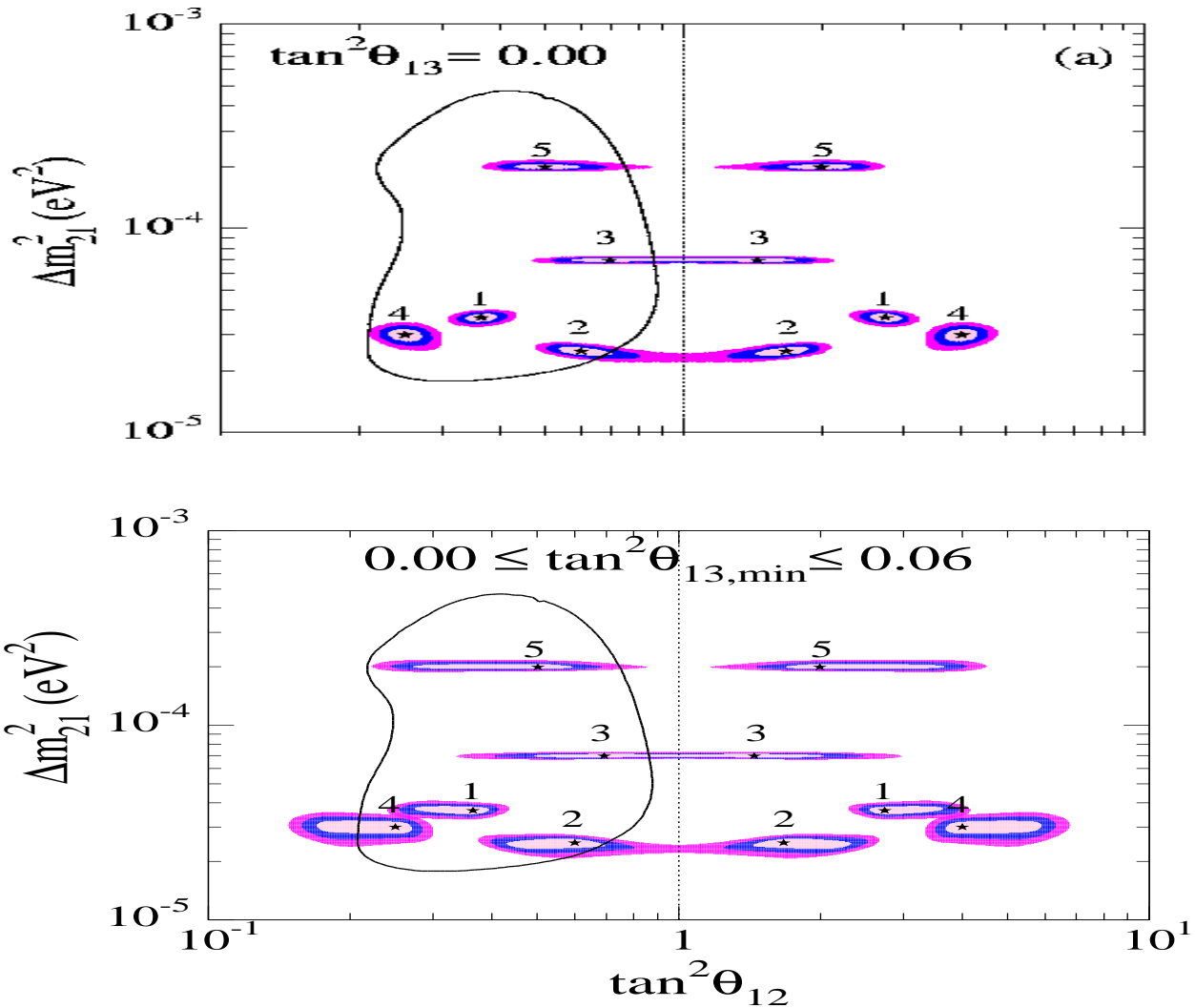


- Effect is very small on the  $\sin^2 2\theta_{\odot}$  precision:  $\lesssim 2 - 3\%$

# Effect of $\sin^2 \theta$ on Solar Parameters

## KamLAND

$$\delta(\sin^2 2\theta_{\odot}) \approx \frac{2\Delta P_{ee} \sin^2 \theta}{\sin^2 \frac{\Delta m_{\odot}^2 L}{4 E_{\nu}}} + 2 \frac{(1 - \sin^2 2\theta_{\odot} \sin^2 \frac{\Delta m_{\odot}^2 L}{4 E_{\nu}}) \Delta(\sin^2 \theta)}{\sin^2 \frac{\Delta m_{\odot}^2 L}{4 E_{\nu}}}$$

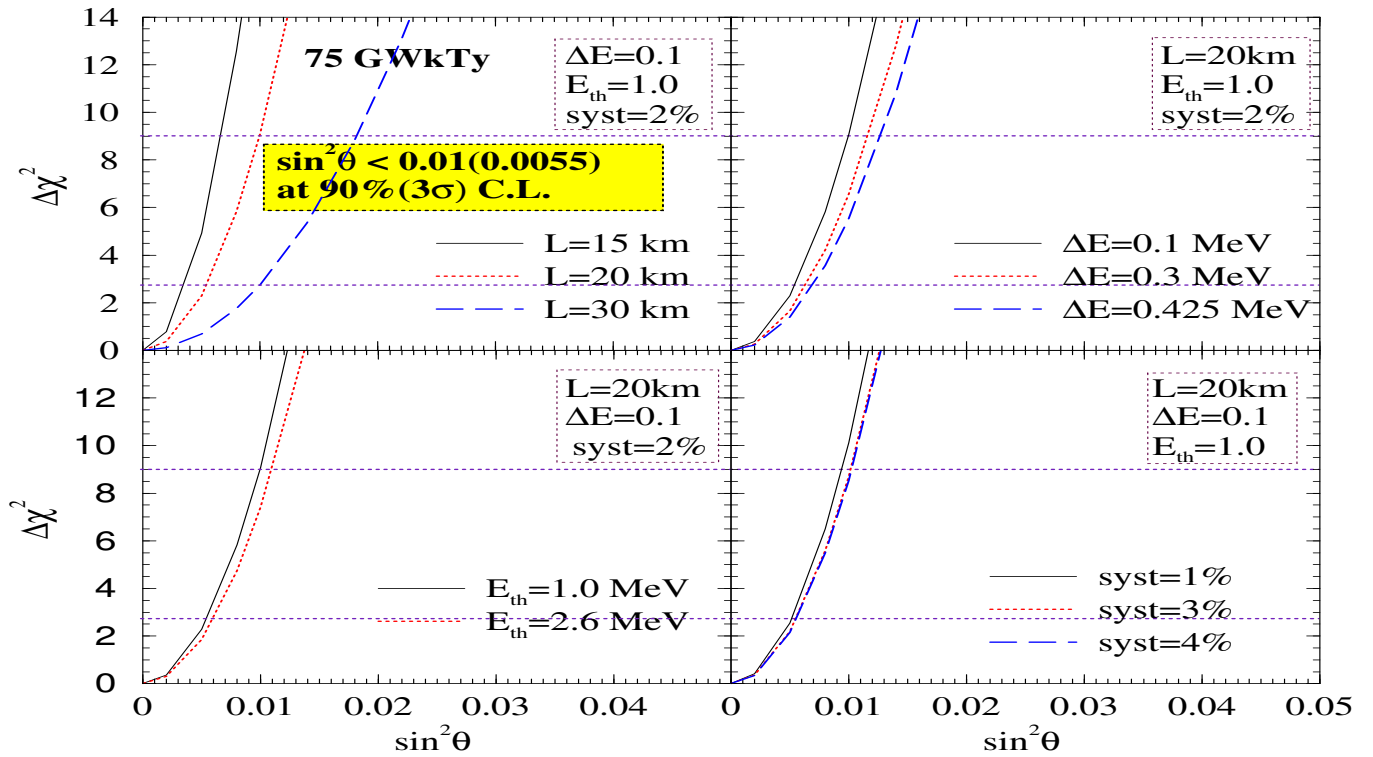
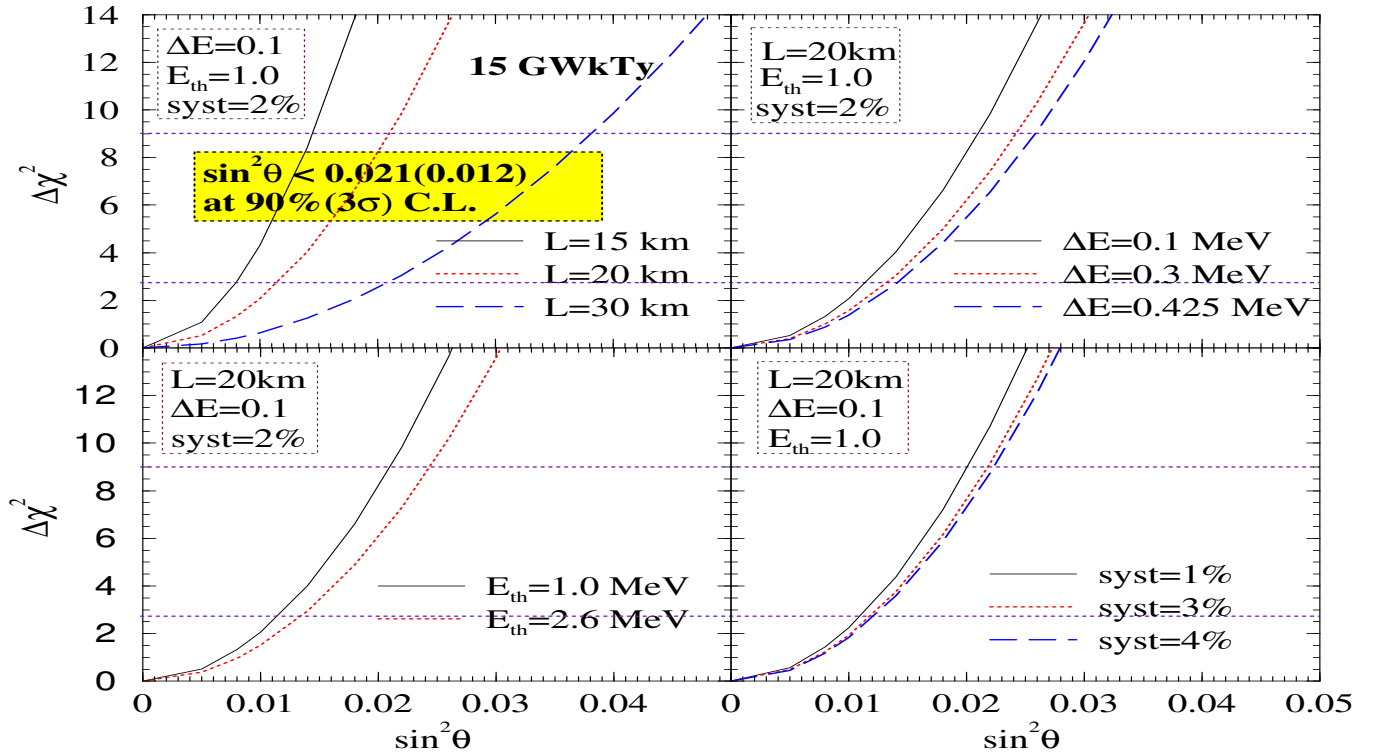


M.C.Gonzalez-Garcia, C.Pena-Garay, hep-ph/0111432

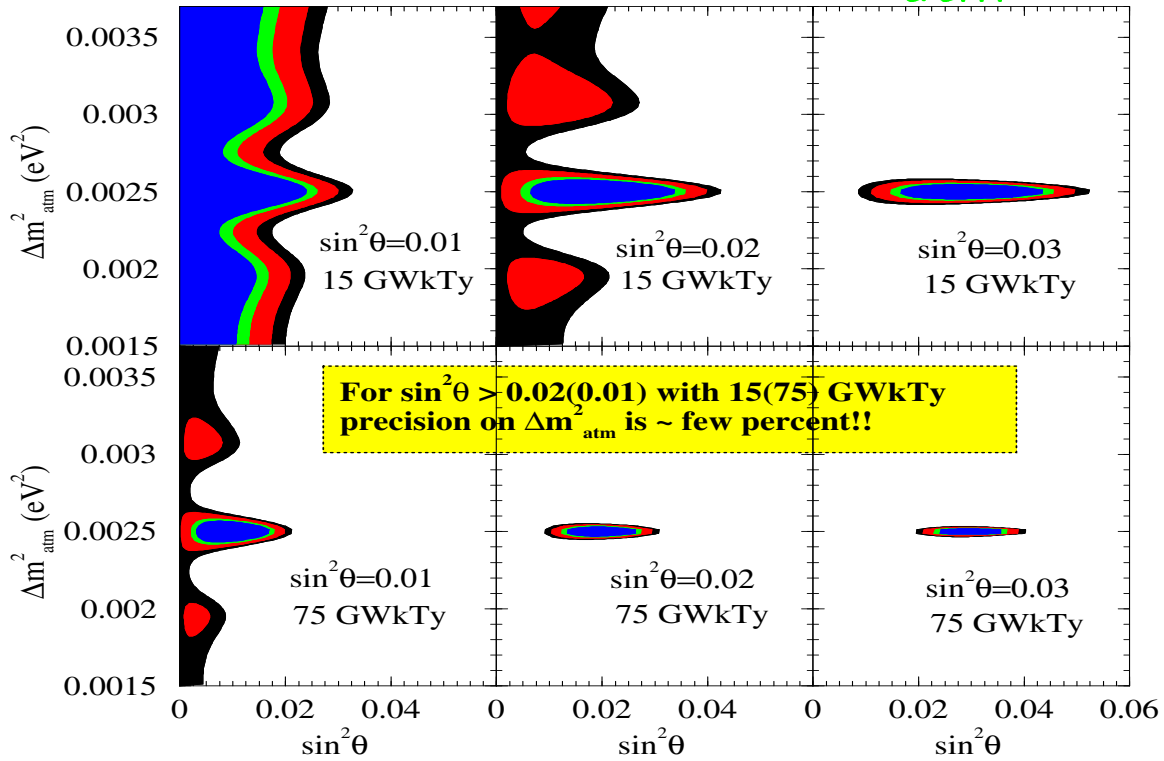
- KamLAND has a problem with  $\theta_{\odot}$  here as well
- Experiments sensitive to SPMIN are required



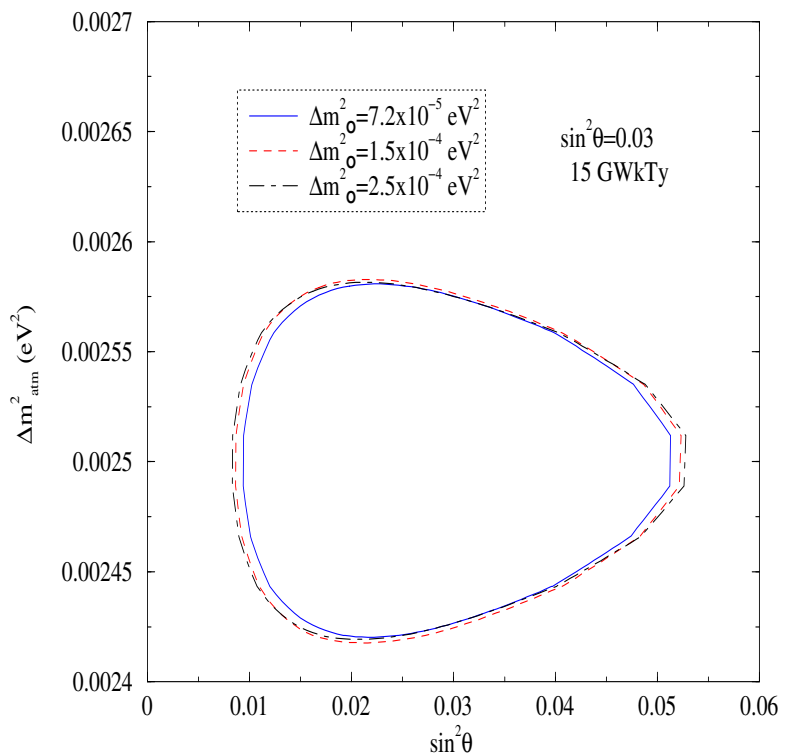
# Limit on $\sin^2 \theta$



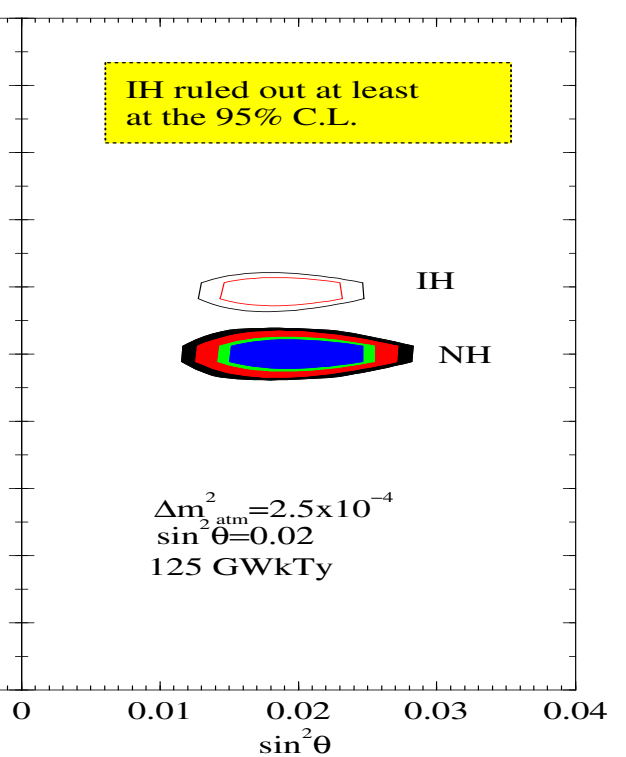
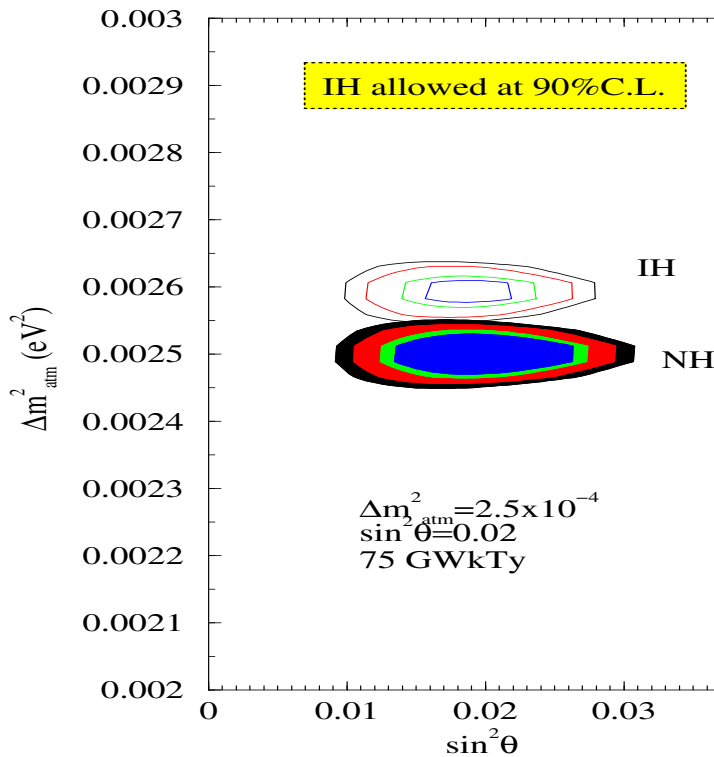
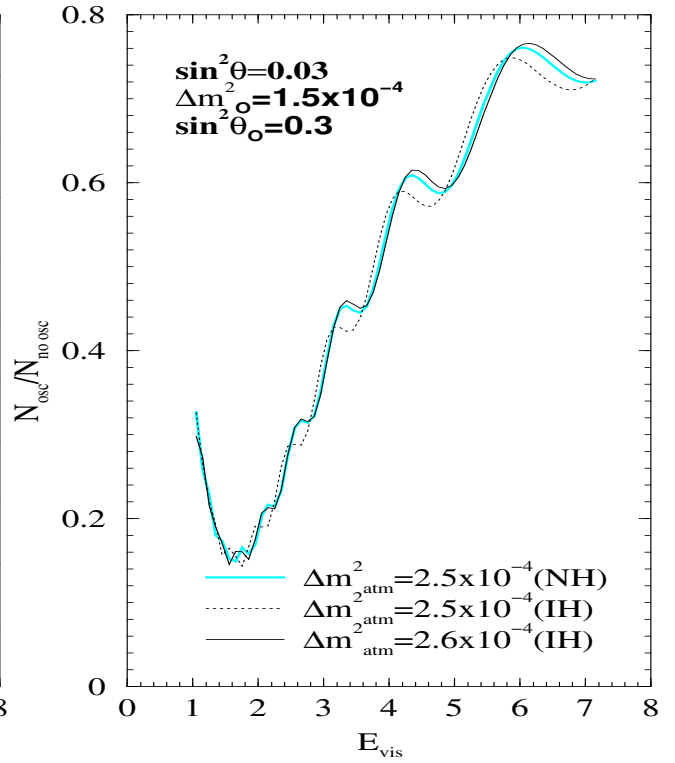
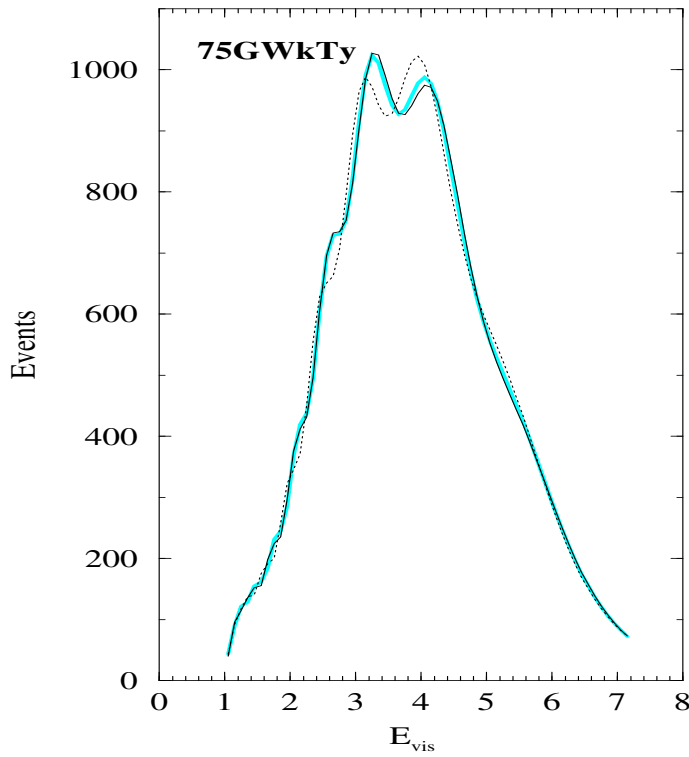
# Precision Measurement of $\Delta m_{\text{atm}}^2$ & $\sin^2 \theta$



- $\sin^2 \theta \Rightarrow$  large
- Statistics  $\Rightarrow$  large
- $\Delta m_{\odot}^2 \Rightarrow$  any
- $\Delta E \Rightarrow$  small
- Possible for low-LMA
- high-LMA  $\Rightarrow$  hierarchy



# Determination of the Mass Hierarchy



## Closer look at hierarchy dependence

- NH “data”  $\Rightarrow \Delta m_{\text{atm}}^2$
- IH “fits” this “data” with  $\Rightarrow \Delta m_{\text{atm}}^2 + \Delta m^2$

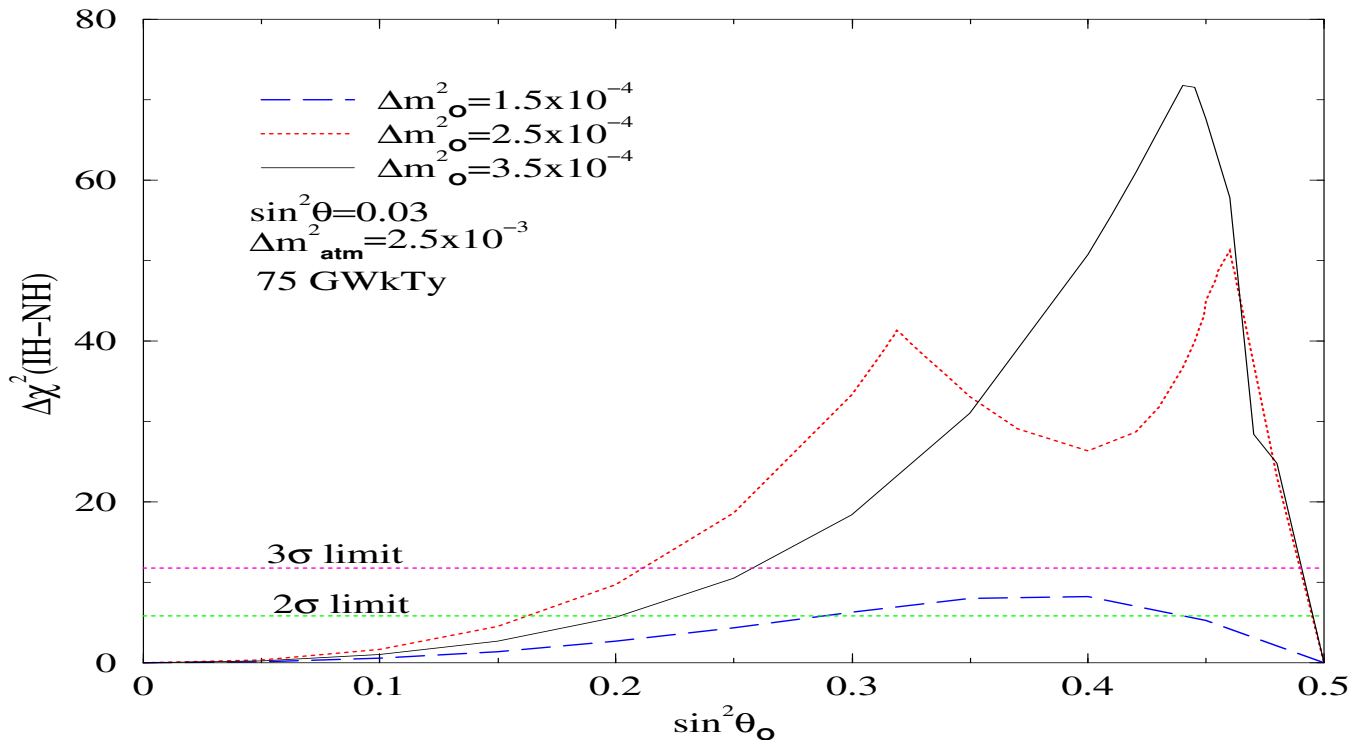
$$P_{IH} - P_{NH} \approx 4 \sin^2 \theta \cos^2 \theta \left[ -\sin \frac{\Delta m_{\text{atm}}^2 L}{2E} \sin \frac{\Delta m^2 L}{4E} + \cos 2\theta_{\odot} \sin \frac{\Delta m_{\text{atm}}^2 L}{2E} \sin \frac{\Delta m_{\odot}^2 L}{4E} \right]$$

↓

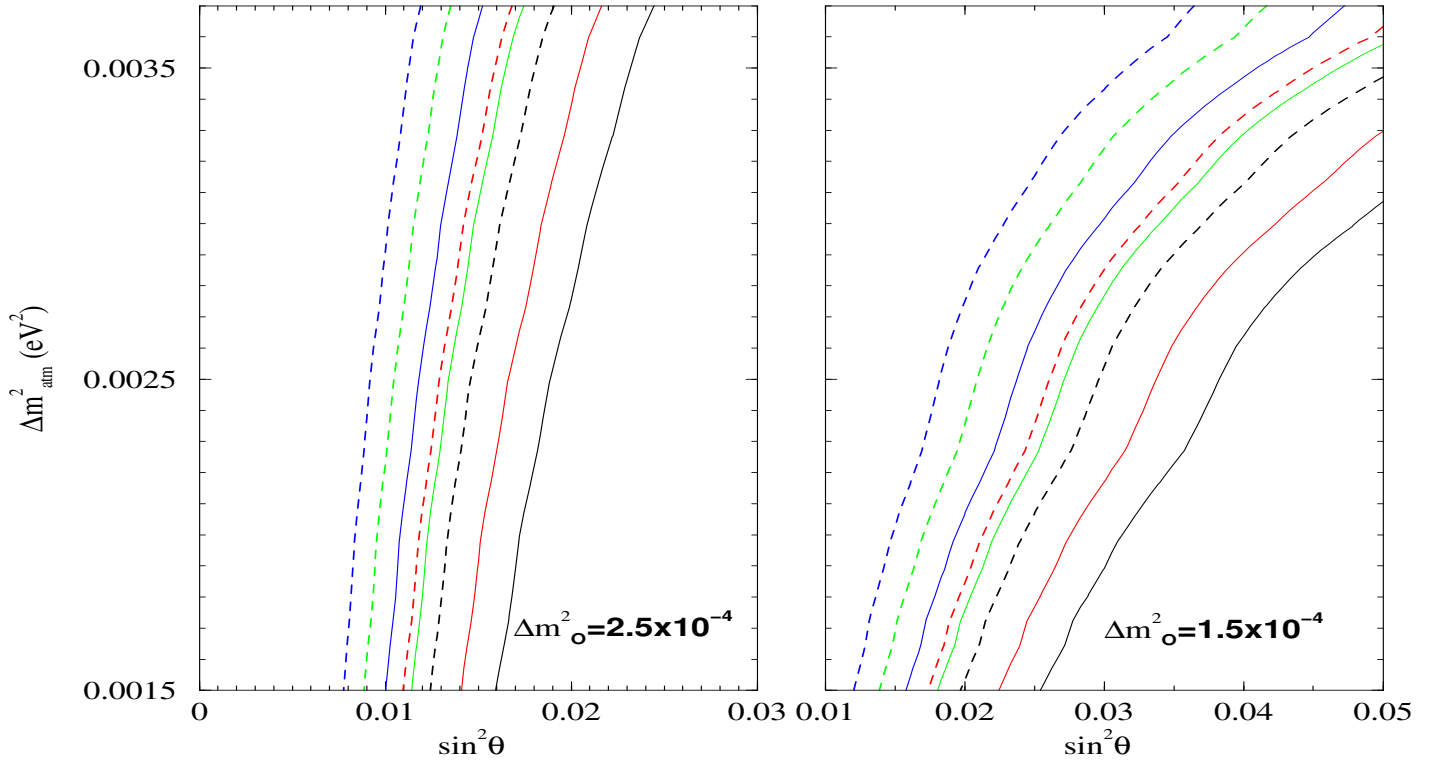
$$P_{IH} - P_{NH} = 0, \text{ when } \sin \frac{\Delta m^2 L}{4E} = \cos 2\theta_{\odot} \sin \frac{\Delta m_{\odot}^2 L}{4E}$$

↓

- $\Delta m^2 \leq \Delta m_{\odot}^2$
- Satisfied identically only for  $\cos 2\theta_{\odot} = 0, 1$



# Determination of the Mass Hierarchy



Hierarchy determination certainly possible



With a single intermediate baseline detector? YES

But more easily with the intermediate baseline reactor experiment + precise information on  $\Delta m^2_{\text{atm}}$  from Superbeam experiments

# Conclusions

- Reactor neutrino experiments have huge potential for precision measurement of oscillation parameters
- Baseline is crucial to identify which parameters would be best determined  $\Rightarrow$  SPMIN is important
- The long baseline KamLAND experiment can measure  $\Delta m_{\odot}^2$  with very high accuracy  $\Rightarrow$  SPMAX
- However the sensitivity of KamLAND to  $\sin^2 \theta_{\odot}$  is not good
- If low-LMA is true, then a 70 km baseline reactor experiment can measure  $\sin^2 \theta_{\odot}$  down to  $\sim 10\%$  accuracy
- Short baseline experiments like CHOOZ/Palo Verde but with one reactor two detector technique can measure  $\sin^2 \theta$  to a very high accuracy (talks by O.Yasuda and P.Huber)
- The Intermediate baseline reactor experiment with  $L \sim 20 - 30$  km can measure  $\Delta m_{\odot}^2$  and  $\sin^2 \theta_{\odot}$  to a few percent accuracy if the high-LMA is the true solution.
- It can also improve the limit on  $\sin^2 \theta$
- If  $\sin^2 \theta$  is large, under favorable experimental conditions it can put very precise limits on the value of  $\Delta m_{\text{atm}}^2$
- Finally under ideal conditions and/or input on  $\Delta m_{\text{atm}}^2$  from the superbeam experiments, it can lead to some insight on the type of Neutrino Mass Hierarchy