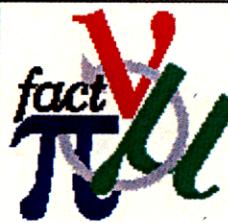


# 3

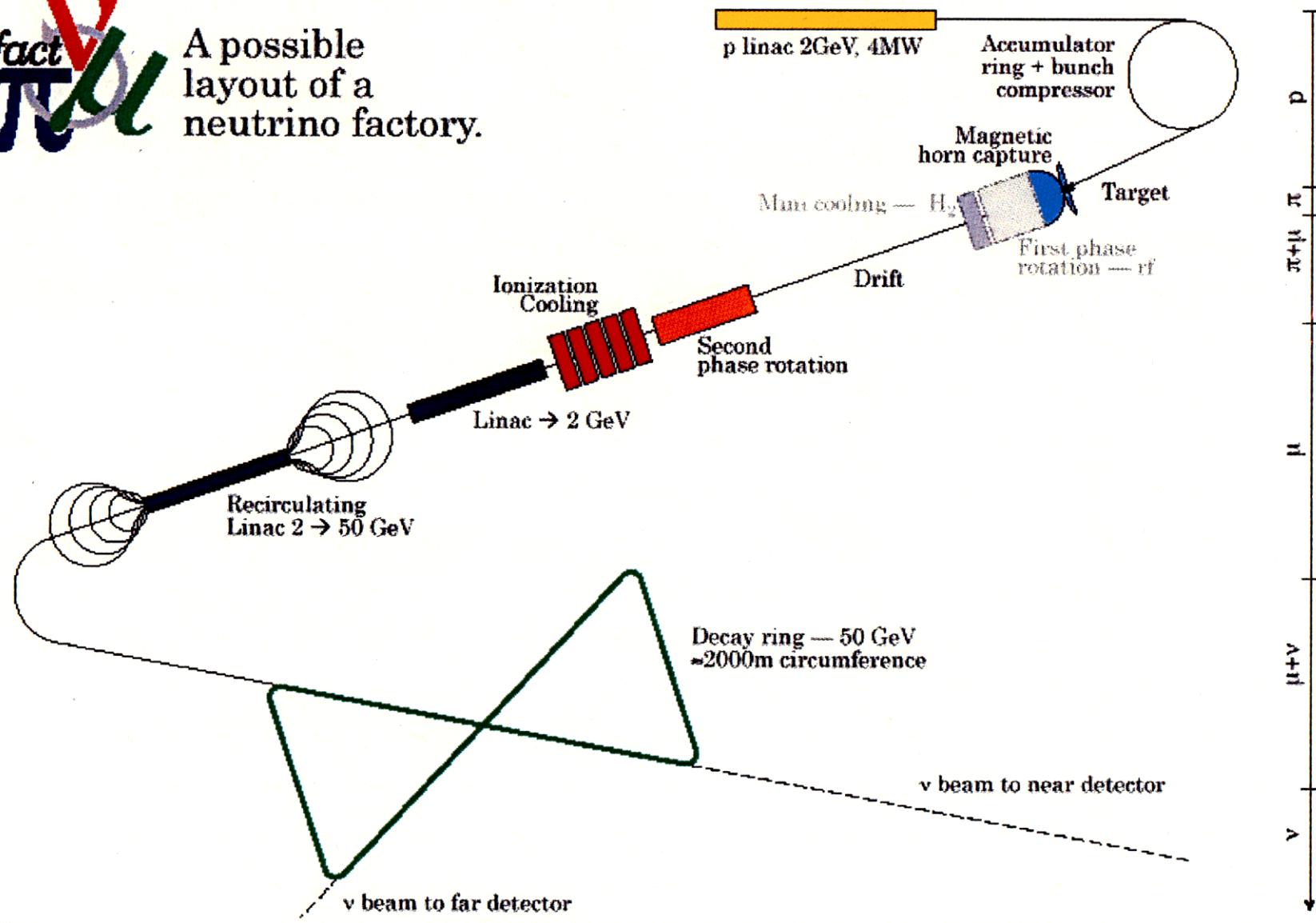
Monday  
morning

de Rijster



## A possible layout of a neutrino factory.

2000-05-01 V Peter Gruber CERN-PS



**A possible layout of a neutrino factory.**

$\mu$  COLLIDERS BUDKER 1969  
SKRINSKI, colls 1971  
NEUFFER 1979

$\nu$  FACTORY KOSHANEV 1974  
WOJCICKI COLLINS 1974

$\nu$  FACTORY BANG GEER 1998

A BOOM OF INTEREST  
AND ACTIVITY

SKEPTICS WHO HAVE  
LOOKED AT THE POTENTIAL  
OF A  $\nu$ -FACTORY CONVERTED  
INTO FANS BY THE FACTS

# EXAMPLES OF CONVERTED SKEPTICS

● BELÉN GAVELA and PILAR HERNANDEZ

GIGANTIC POTENTIAL OF THE  
“WRONG-SIGN” MUON SIGNALS  
IN A (REALISTIC) 3-FAMILY CONTEXT

→ THEMSELVES, ME

● SOME POOH POOHING  
EXPERIMENTALISTS

● EVEN SOHE DGs

$\Delta m^2$ , OBSERVED BY

ATMOSPHERIC, SOLAR

$\nu$ -DETECTORS

ARE AT THE SCALE

EXPECTED FOR  $m_\nu$ 's

IN (SUPERSYMMETRIC)

GRAND  
UNIFIED  
THEORIES

SM

$$L = \begin{pmatrix} \nu_L \\ e_L \end{pmatrix}$$

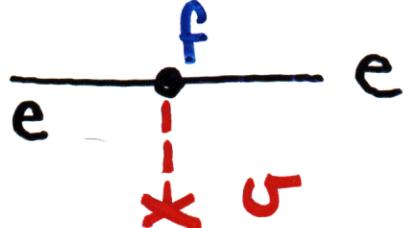
DOUB.  $e_R$  SINGLET  
 $\nu_R$  USELESS

$$\phi = \begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix} \rightarrow \begin{pmatrix} 0 \\ \sigma \end{pmatrix}$$

SSB

$$\Delta \mathcal{L} = f \bar{L} \phi e_R$$

$$\rightarrow f \nu \bar{e} e = m_e \bar{e} e$$



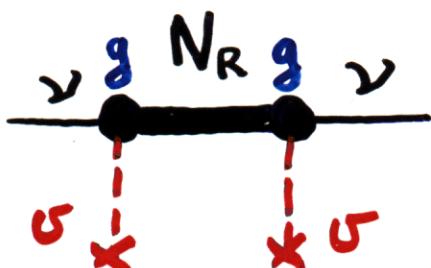
GUTs

$$SO(10) \supset SU(3) \otimes SU(2)_L \otimes U(1)$$

COMPLETE  $q, l$  FAMILY ( $+N_R$ )  $\in 16$

Inevitably

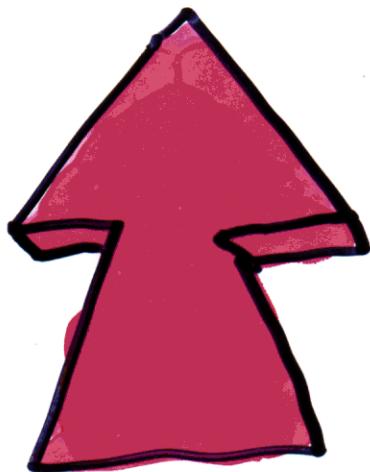
$$\Delta \mathcal{L}' = g \bar{L} \phi N_R$$



$$M_\nu = \frac{(g\sigma)^2}{M_N}$$

$$M_N = \frac{(g\sigma)^2}{m_\nu} \sim \frac{(250 \text{ GeV})^2}{0.01 \text{ eV}}$$

$M_N \sim 6 \cdot 10^{15} \text{ GeV}$  IS THE SCALE OF  
 (ss)  $SO(10)$  GRAND UNIF. !!



MECHANISM 1979  
GELLMANN, RAMOND,  
SLANSKI ; YANAGIDA

MORE RECENT IDEAS ON  
 $m_\nu$ 'S OR LEPTONIC WEAK  
MIXING ANGLES ARE ALL

"HOW TO..."

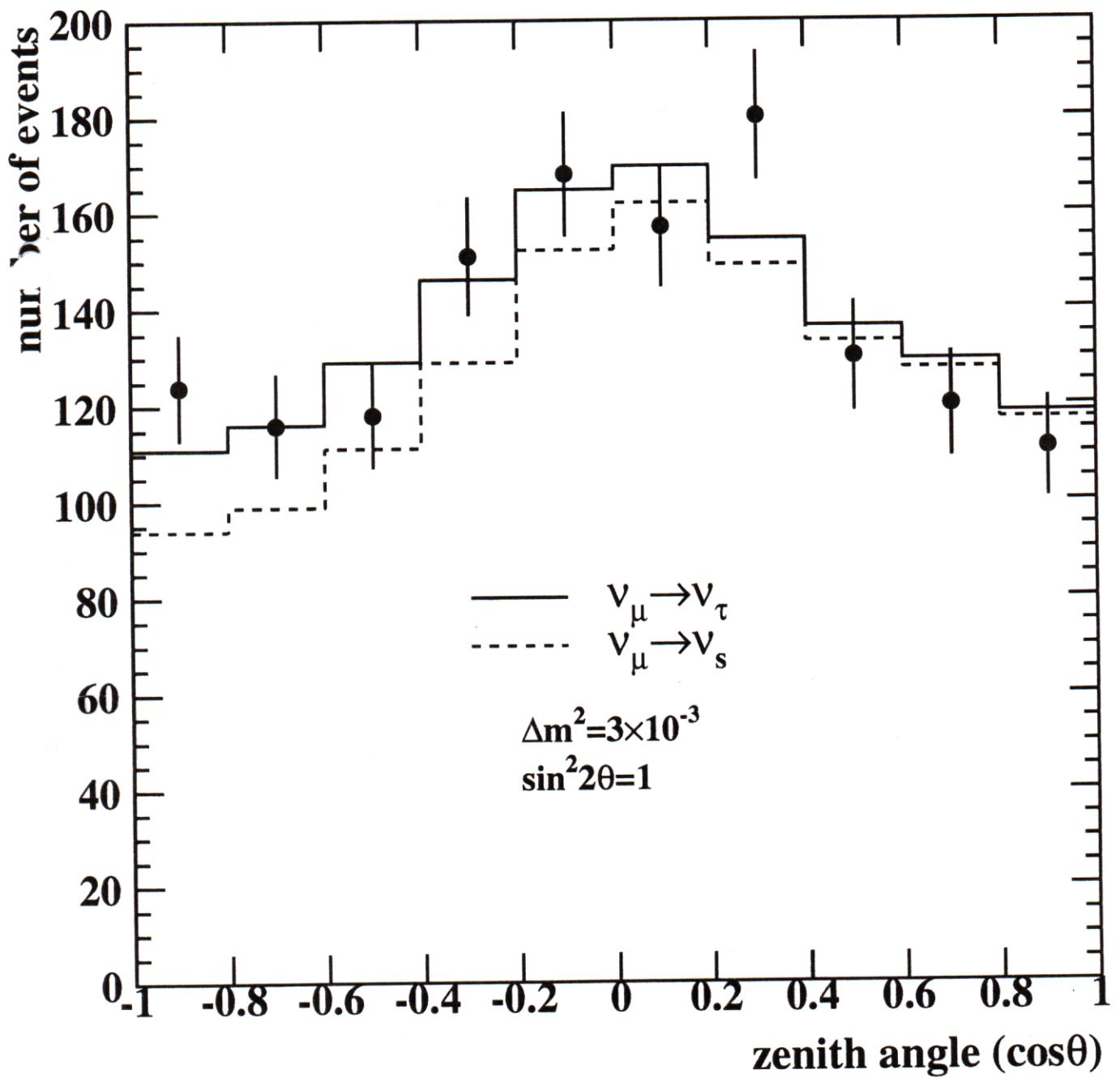
or

"WHY NOT..."

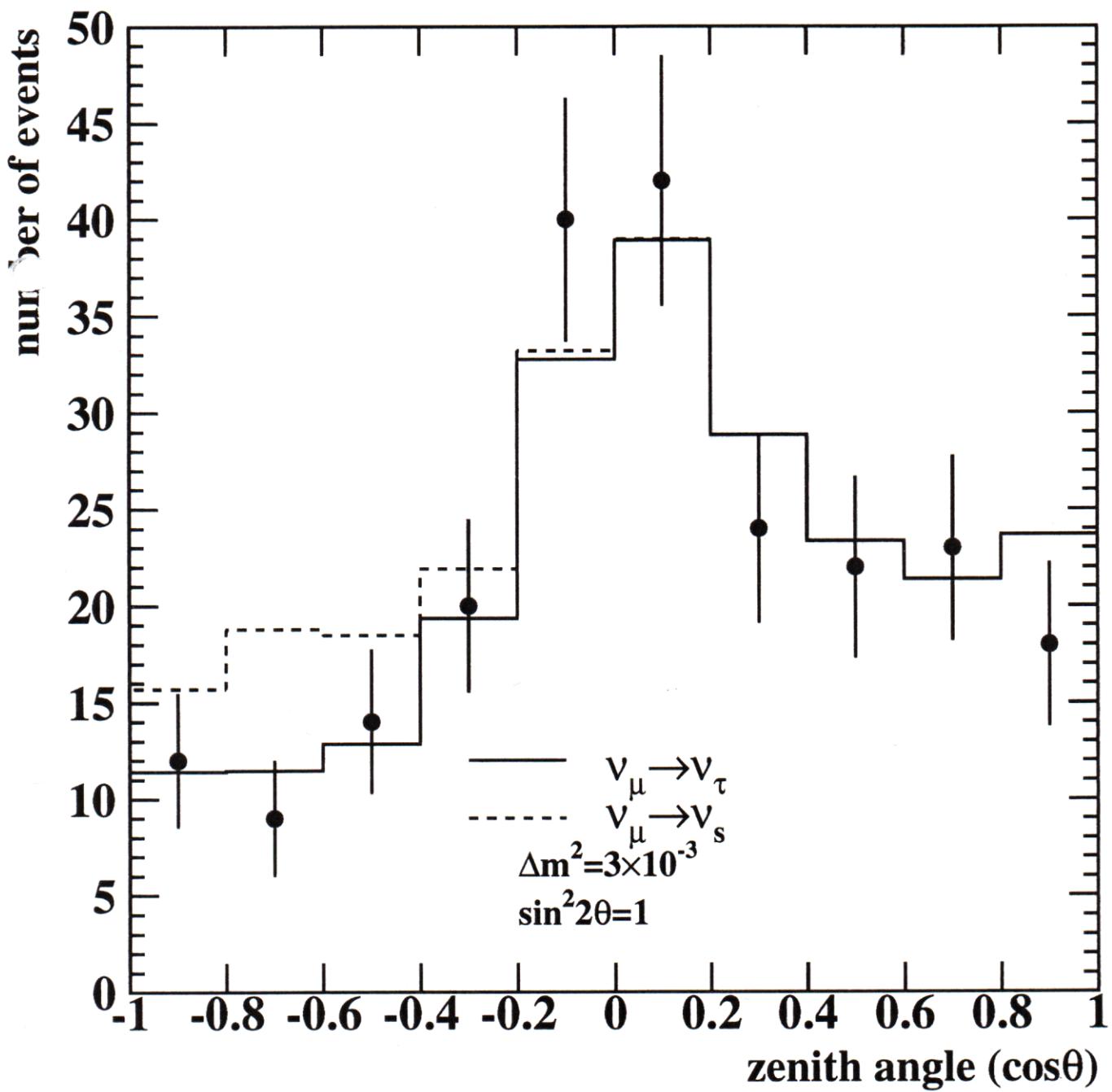
... DO THIS OR THAT

eg : MAXIMAL V-MIXING

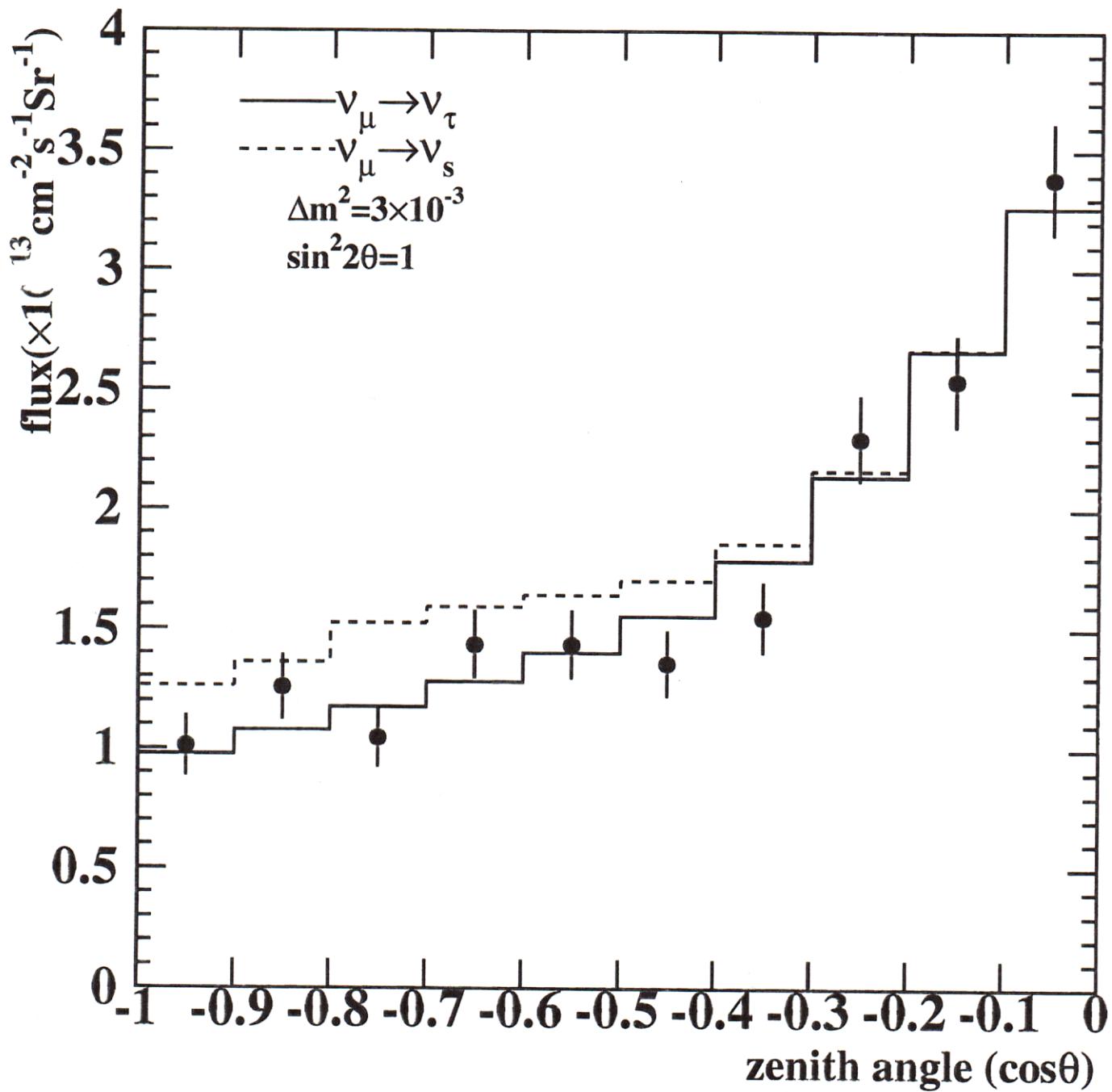
**zenith angle distribution of N.C. enriched multi-ring events (990days)**



zenith angle distribution of high E ( $E_{\text{vis}} > 5 \text{ GeV}$ ) PC events (990 days)



**zenith angle distribution of upward through going  $\mu$  events (1070days)**



# Super-Kamiokande

Run 8474 Event 5580445

100-03-21:14:00:07

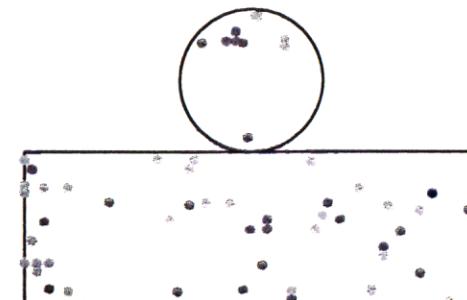
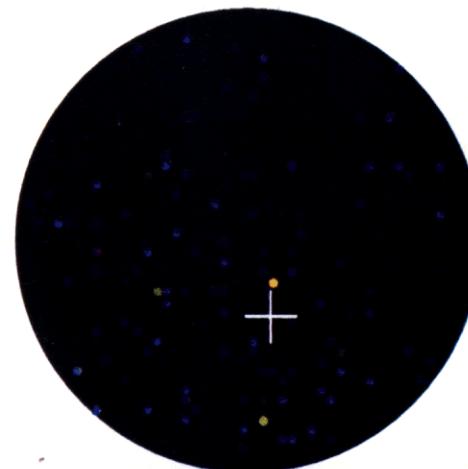
Inner: 1454 hits, 4484 pE

Outer: 1 hits, 0 pE (in-time)

Trigger ID: 0x07

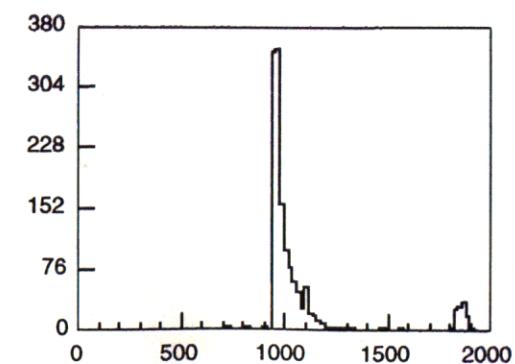
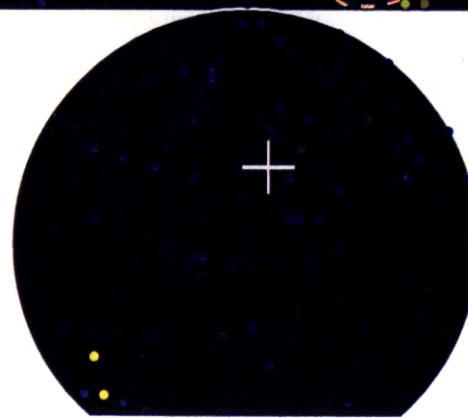
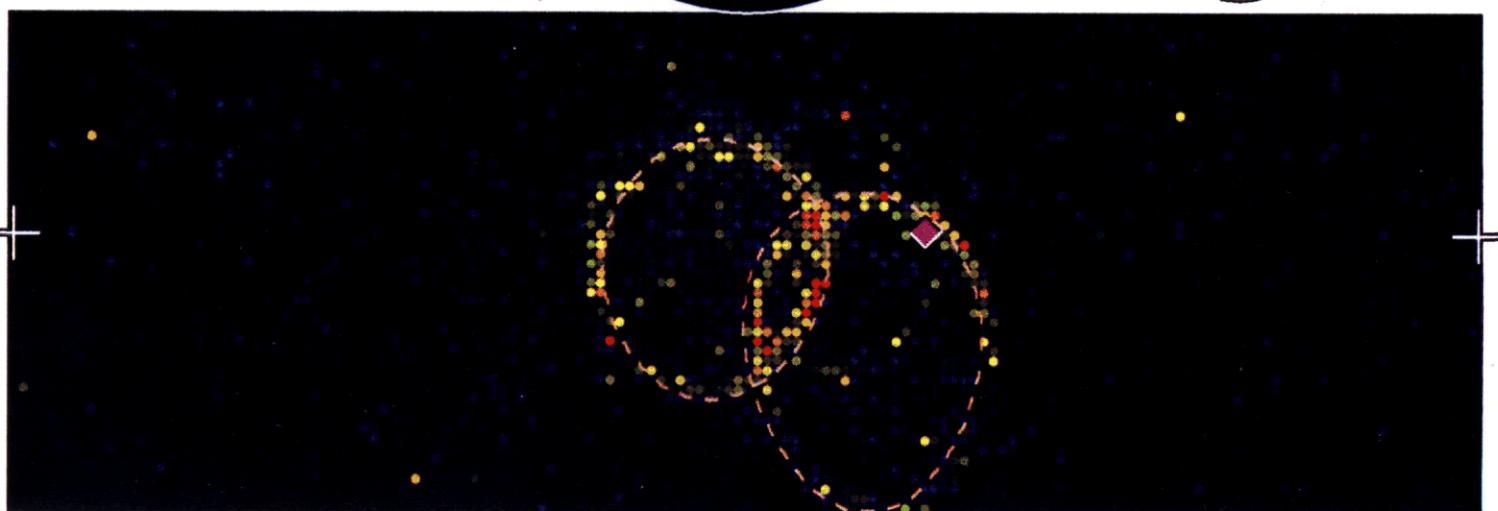
D wall: 1107.4 cm

Fully-Contained

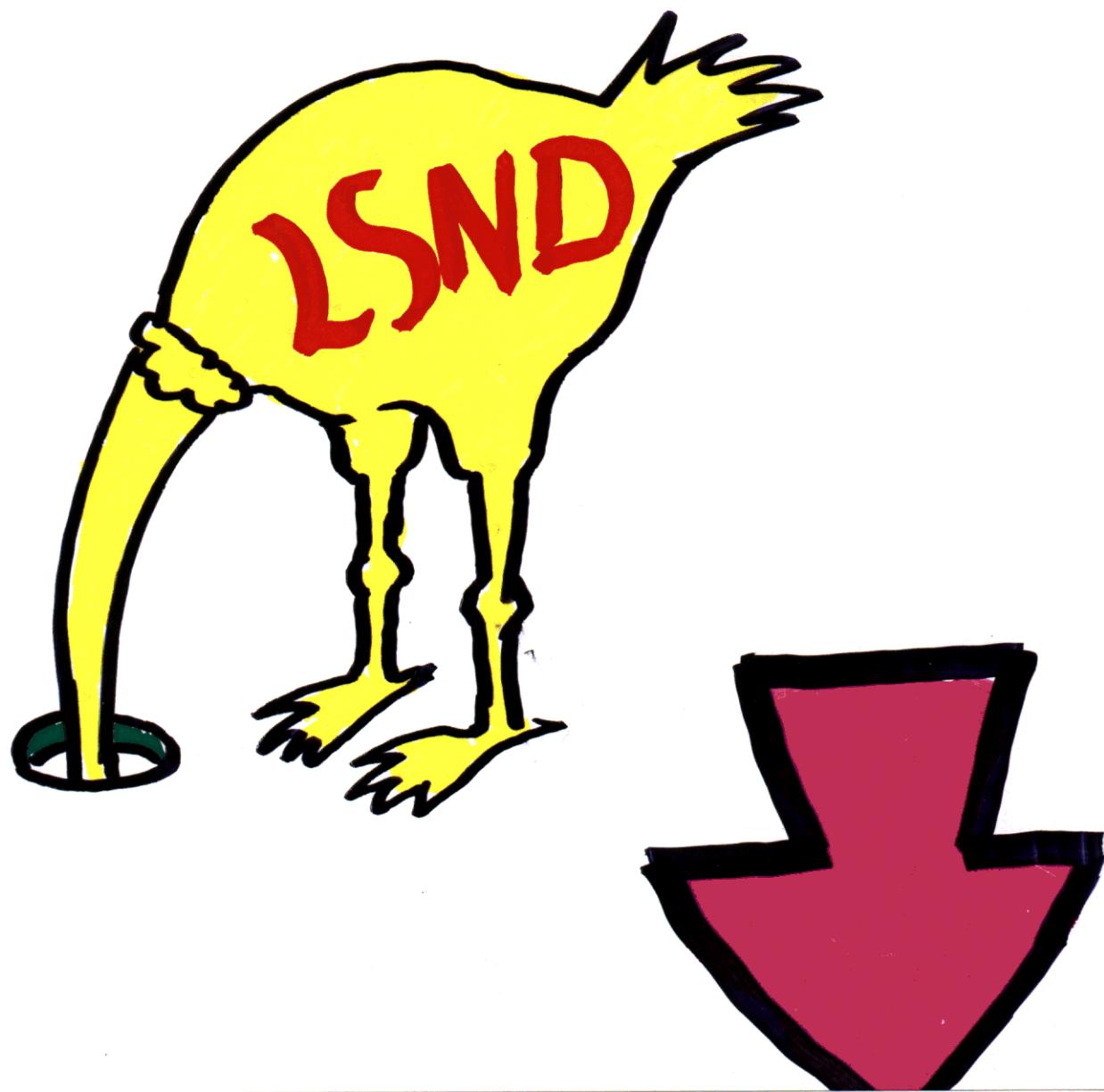


## Charge (pe)

- >26.7
- 23.3-26.7
- 20.2-23.3
- 17.3-20.2
- 14.7-17.3
- 12.2-14.7
- 10.0-12.2
- 8.0-10.0
- 6.2- 8.0
- 4.7- 6.2
- 3.3- 4.7
- 2.2- 3.3
- 1.3- 2.2
- 0.7- 1.3
- 0.2- 0.7
- < 0.2



**FORGET  
STERILE**



## WHAT WE KNOW FOR SURE

$$\Delta m_{23}^2 \sim 3 \cdot 10^{-3} \text{ eV}^2 \quad \theta_{23} \sim 45^\circ \quad \text{ATMS}$$

$$\theta_{13} < 13^\circ \quad (\text{CHOOZ})$$

$$c_{ij} = \cos \theta_{ij} \quad s_{ij} = \sin \theta_{ij}$$

$$\Delta m_{ij}^2 = m_j^2 - m_i^2$$

$$\Delta m_{12}^2, \Delta m_{23}^2$$

$$\theta_{12}, \theta_{13}, \theta_{23}, \delta$$

$$U \equiv U_1 U_2 U_3 \equiv$$

$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13} e^{i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

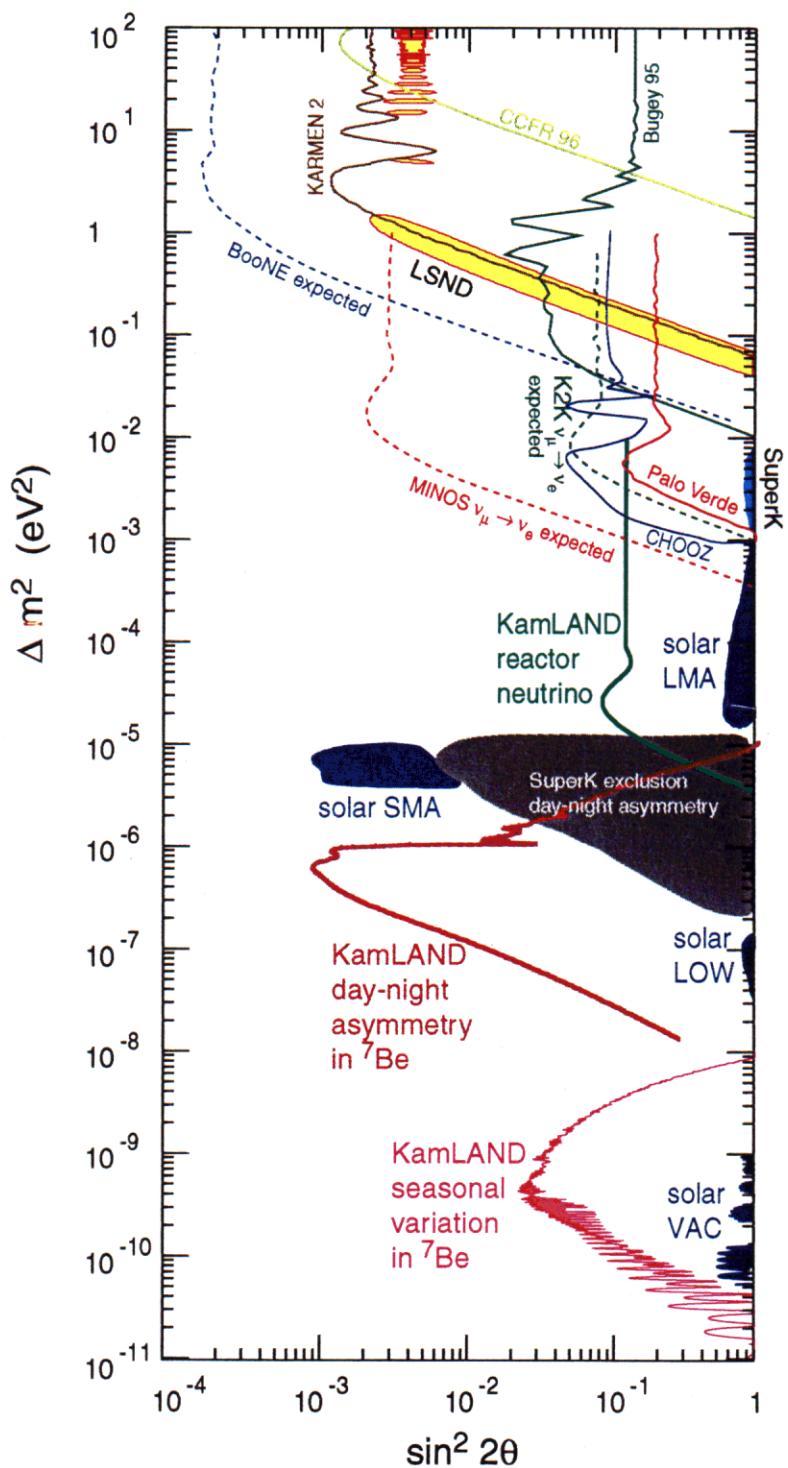


Figure 14: The current and expected limits at some of the future neutrino oscillation experiments. Note that different oscillation modes are shown together.

IFF  $\gamma_0$  

V.O. or LMA-MSW

Bi-MAXIMAL  $U \sim \begin{pmatrix} \frac{\sqrt{2}}{2} & -\frac{\sqrt{2}}{2} & S_{13} \\ \frac{1}{2} & \frac{1}{2} & -\frac{\sqrt{2}}{2} \\ \frac{\sqrt{2}}{2} & \frac{1}{2} & \frac{\sqrt{2}}{2} \end{pmatrix} * CP(\delta)$

IFF LMA-MSW

$$\Delta m_{12}^2 \sim 10^{-4} \div 10^{-5} \text{ eV}^2$$

in

$$\Delta m_{12}^2 \gtrsim \text{a few } 10^{-5} \text{ eV}^2$$

THE BEST OF ALL WORLDS

EVEN  $\delta$  CAN BE MEASURED  
OR CONSTRAINED AT A  $\nu$ -FAC

# TWO 'EXTREME' PHILOSOPHIES

## ① ☺ VALIANT; ☹ OVEROPTIMISTIC

e.g. CERVERA et al. [THE MAGNIFICENT SEVEN]

- { •  $E_\mu = 50 \text{ GeV}$
- 40kT Fe-Sci  $\mu^+, \mu^-$ ;  $e^\pm + \text{NC}$
- $10^{21} \mu^+$  and  $10^{21} \mu^-$  USEFUL DECAYS (54?)
- 4MW beam on target  $\rightarrow 10^{21} \mu/\text{g}$  (25% USED)

## ② ☹ COWARDLY; ☺ REALISTIC

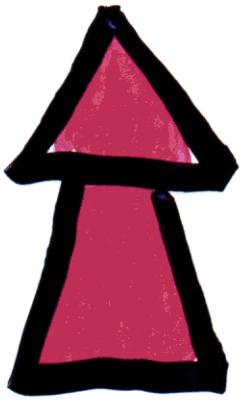
e.g. BARGER et al

- { •  $E_\mu \geq 20 \text{ GeV}$
- 50kT DETECTOR
- $2 \cdot 10^{18} \mu\text{-DECAYS}/\text{y}$

## 1.5 ☺, ☹ SOPHISTICATED DETECTOR

e.g. BUENO et al.

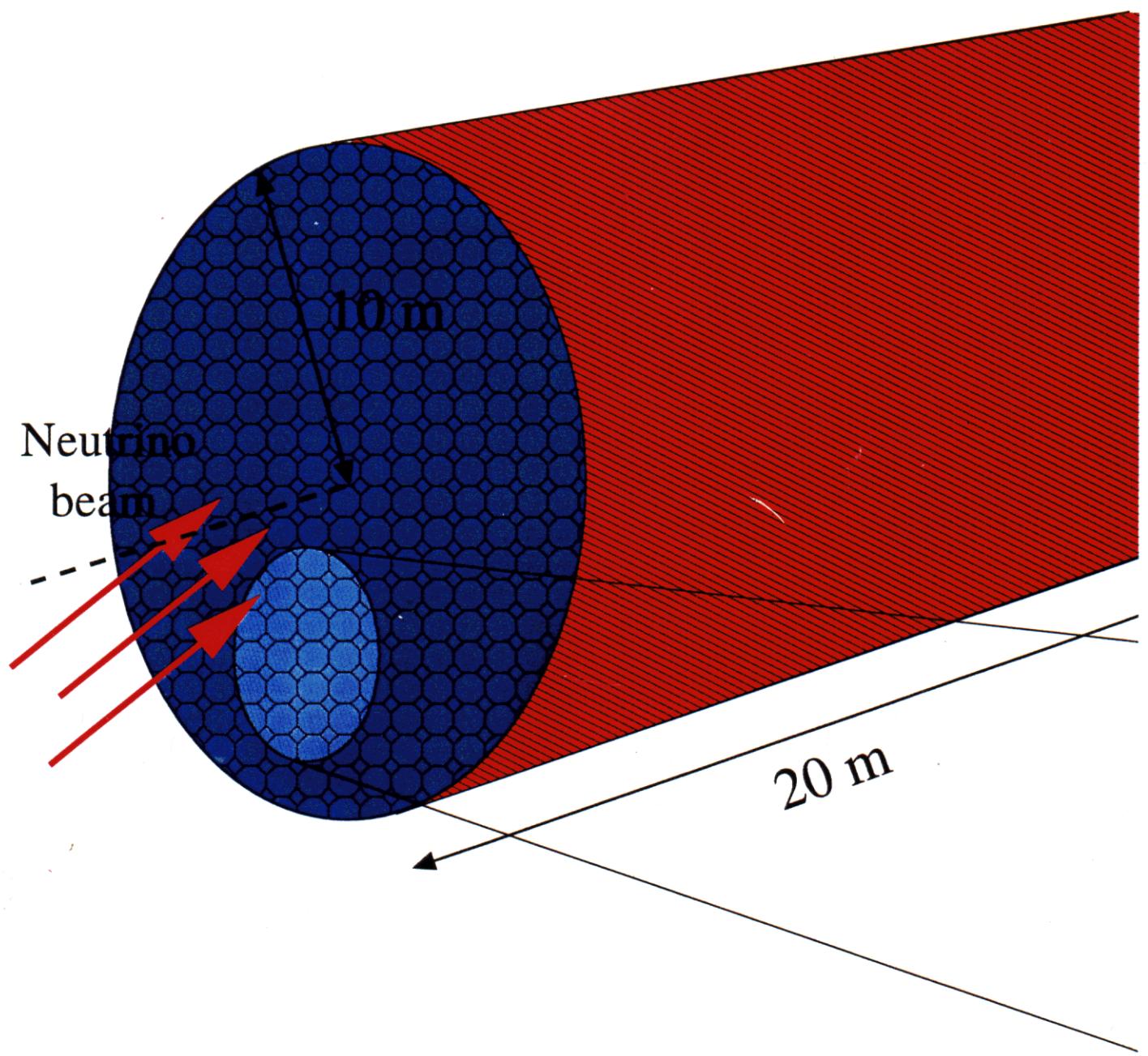
- { •  $E_\mu = 36 \text{ GeV}$
- 10kT ICANOE + DOWNSTREAM  $\mu^{+/-}\text{-DET.}$
- $10^{21} \mu^+$  and  $\mu^-$  USEFUL DECAYS



## MATTER OF DEBATE AND OPINION

MY O.: IT IS FAAAAAAAR  
TOO EARLY NOT TO BE  
OPTIMISTIC, BOTH MACHINE-  
AND DETECTOR-WISE

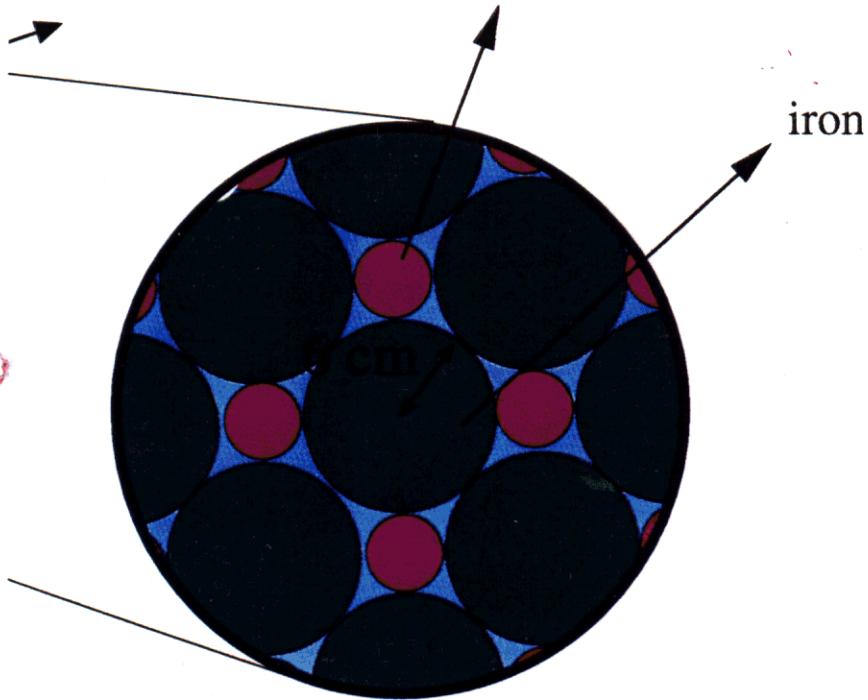
IT IS, HOWEVER, USEFUL  
TO KNOW HOW MUCH LESS  
ONE COULD DO IF ONE  
WAS FORCED TO SAVE \$, SF, £  
...

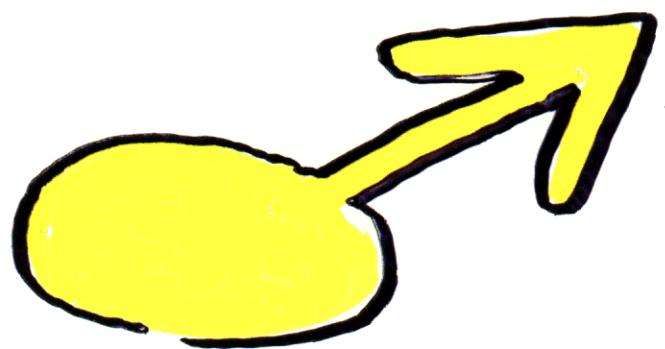


CERVERA, DYDAK, GOMEZ-CARDENAS



scintillator





SPARKLING  
SCIENTIST

OTHER DETECTORS e.g ICA<sup>RUS</sup>  
NOE

MAY BE VERY MUCH MORE  
SOPHISTICATED, BUT THEY  
ARE GOING AHEAD IN SIZE,  
TESTING, TECHNOLOGY and,  
HOPEFULLY, FUNDING

$\mu$  DISAPPEARANCE (NO  $\mu^{+/-}$  TELLING)

$\Delta m_{23}^2, \theta_{23}$  to  $\sim 1\%$

FREUND et al.  $s(\Delta m_{23}^2)$  may

ALSO ONLY REQUIRE  $\mu$ -DISAPP.

EARLY  
DESCENDING  
ATTITUDE



- 99.95 \$

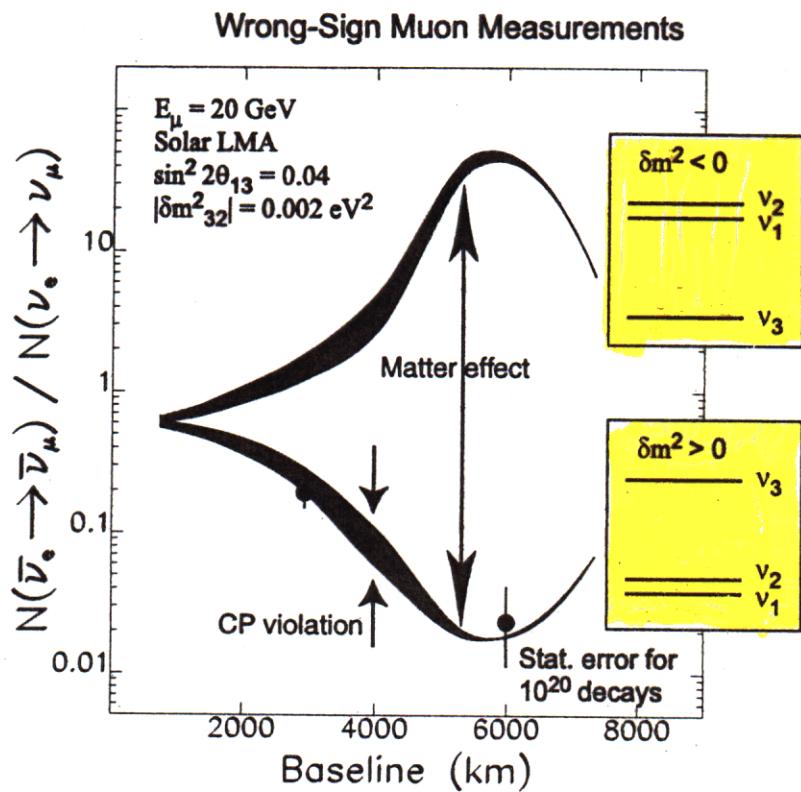
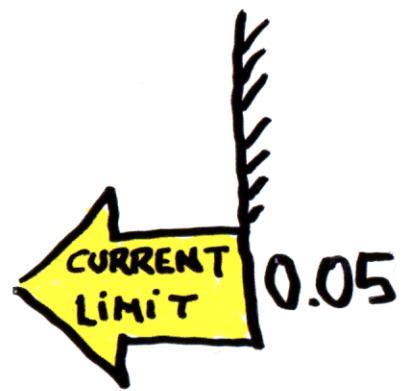
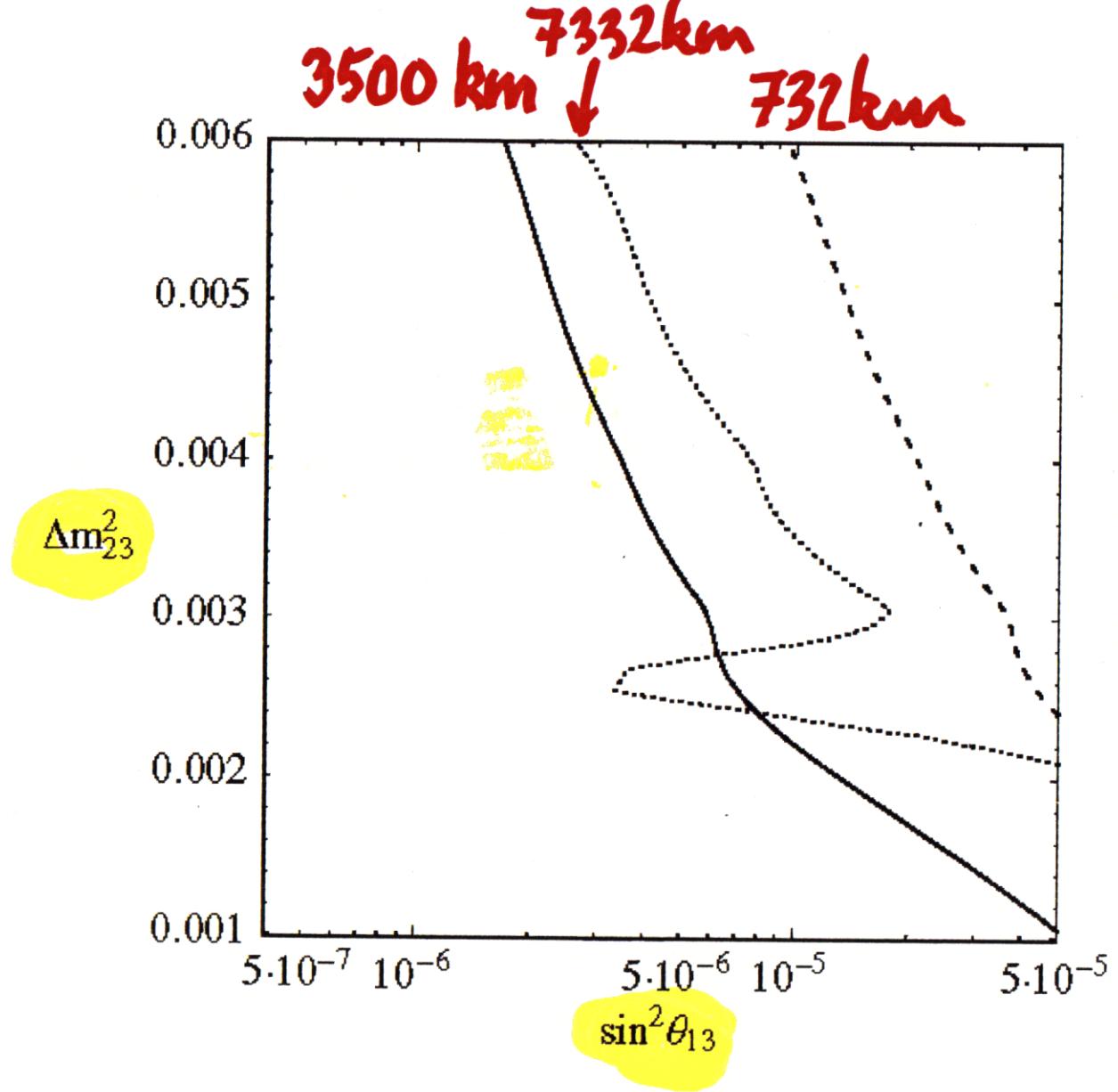


Figure I: Predicted ratios of  $\bar{\nu}_e \rightarrow \bar{\nu}_\mu$  to  $\nu_e \rightarrow \nu_\mu$  rates at a 20 GeV neutrino factory. The upper (lower) band is for  $\delta m^2_{32} < 0$  ( $\delta m^2_{32} > 0$ ). The range of possible CP violation determines the widths of the bands. The statistical error shown corresponds to  $10^{20}$  muon decays of each sign and a 50 kt detector. Results are from Ref. 51.

→ PRESUMABLY BARGER/GEER/RFJA/WHISNANT

COMPREHENSIVE  
STUDY

As in Fig. [12], including as well background errors and detection efficiencies.

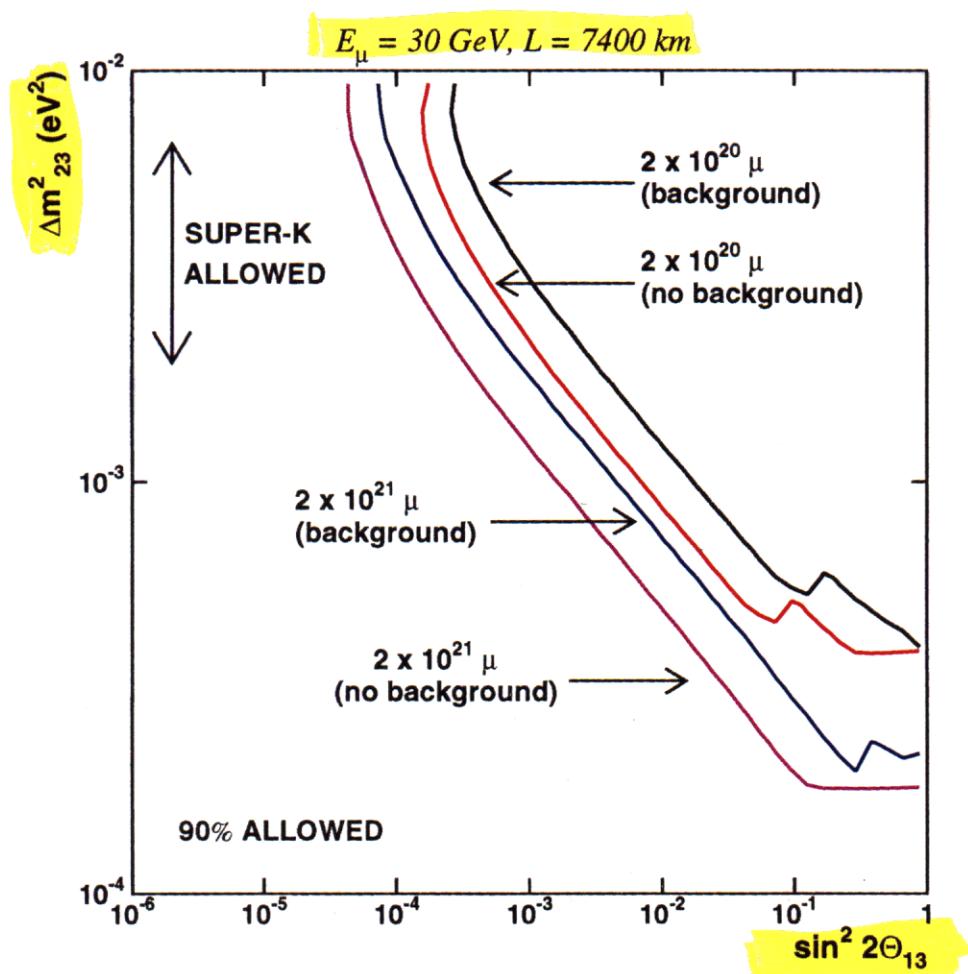


Figure 9: Sensitivity on  $\theta_{13}$

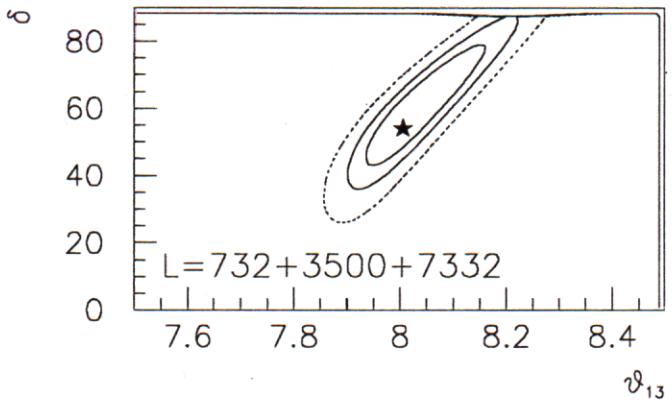
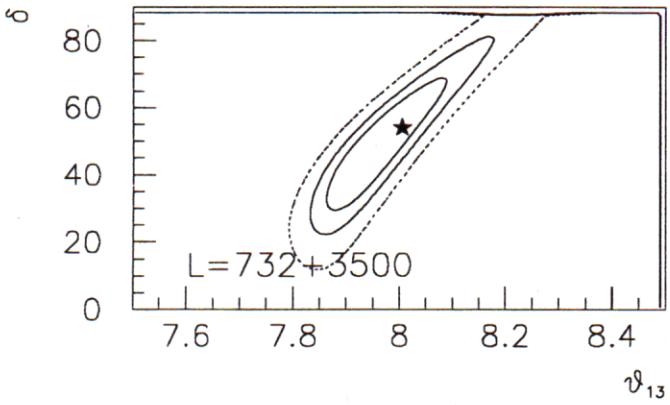
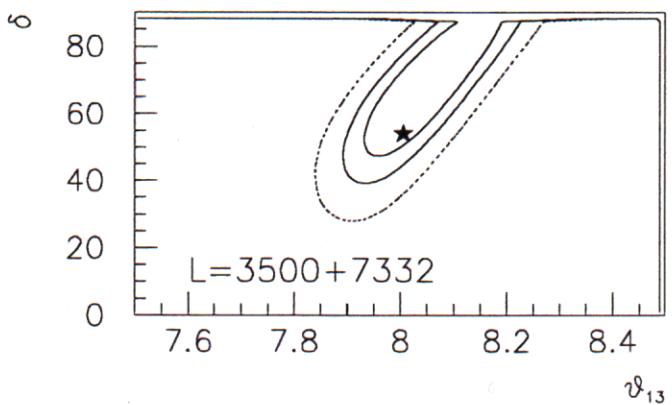
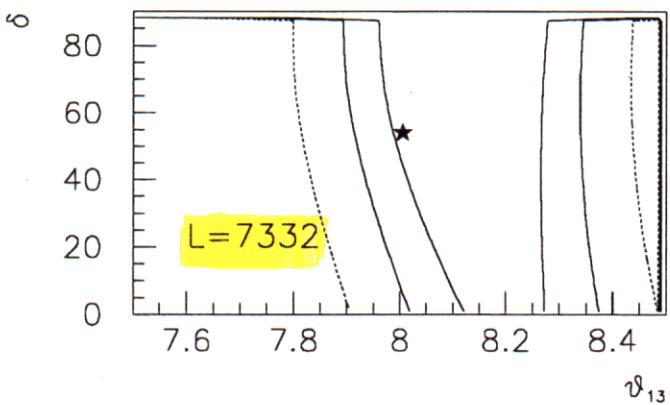
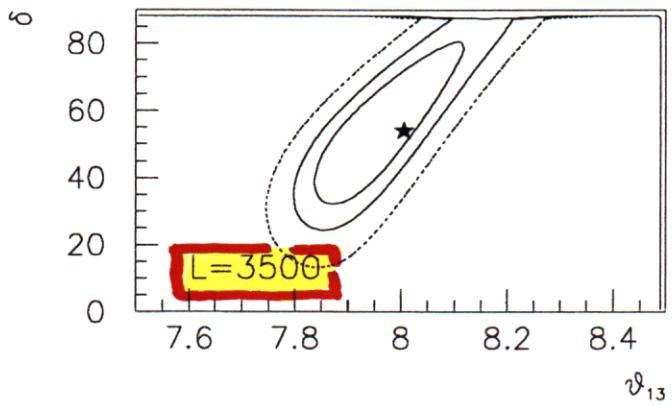
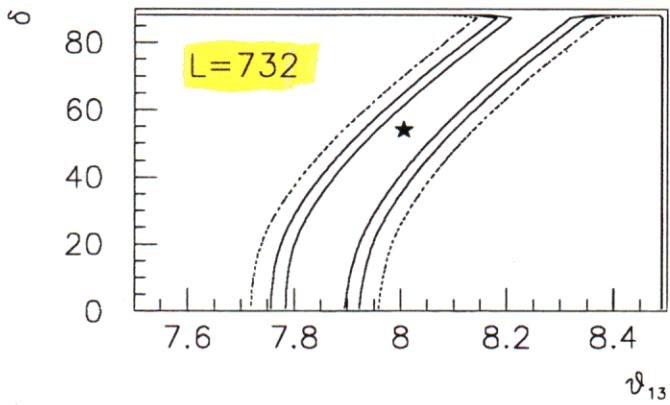
Bueno et al.

MID-RANGE LMA-MSW

$\delta$  ↑  $\theta_{13}$

$$\Delta m_{12}^2 = 5 \cdot 10^{-5} \text{ eV}^2$$

(NOT MOSTOPTIMISTIC)

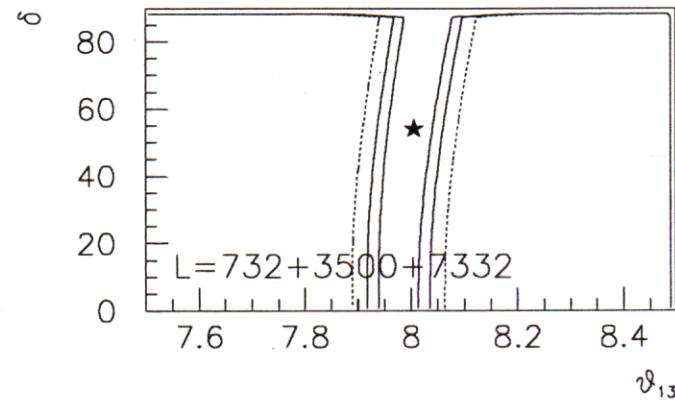
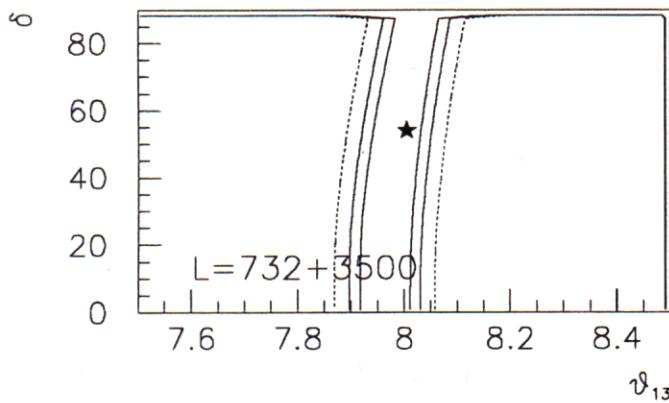
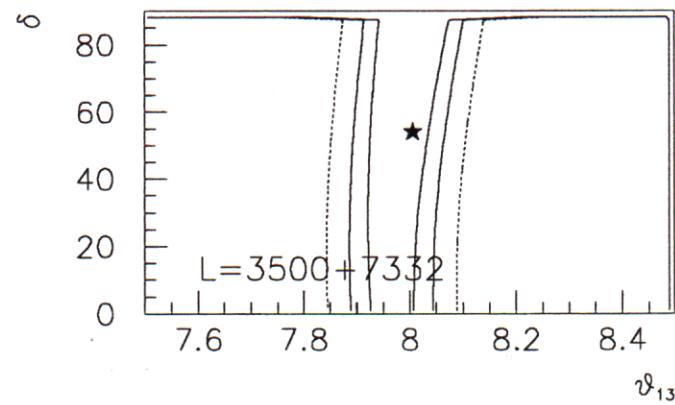
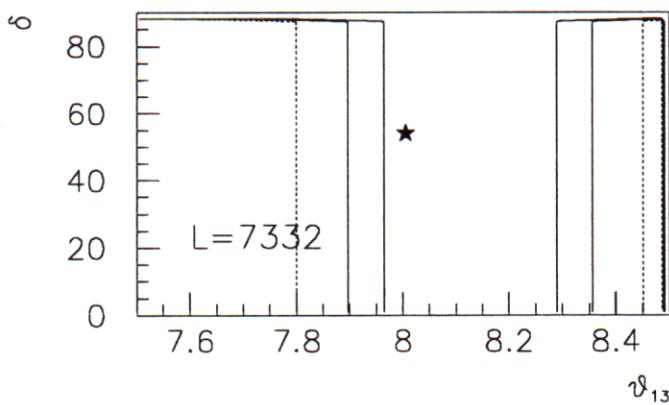
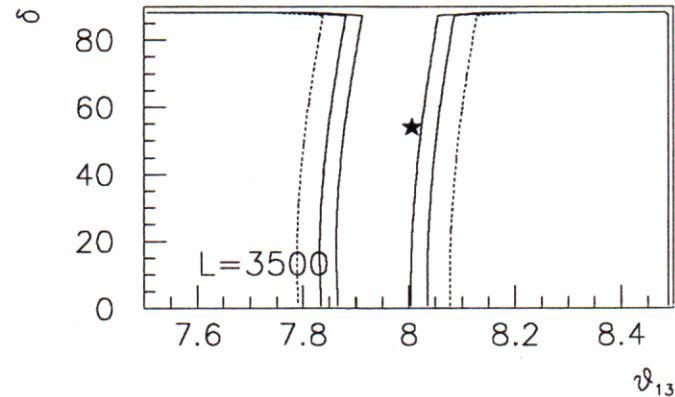
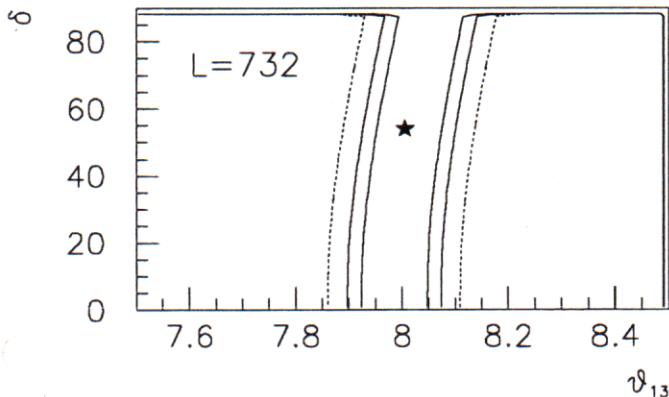


Cervera et al.

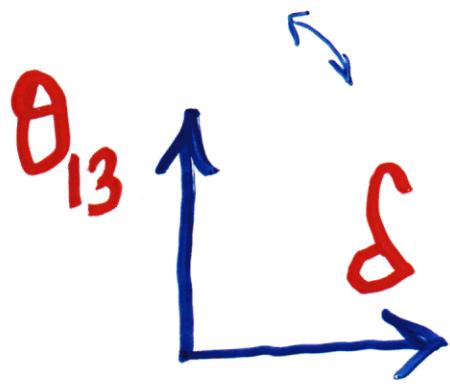
$$\delta \uparrow \quad \Theta_{13}$$

LOW-ENO LMA-NSW

$$\Delta m_{12}^2 = 10^{-5} \text{ eV}^2$$



Gerwera et al.



$$\Delta m_{12}^2 = 10^{-4} \text{ eV}^2$$

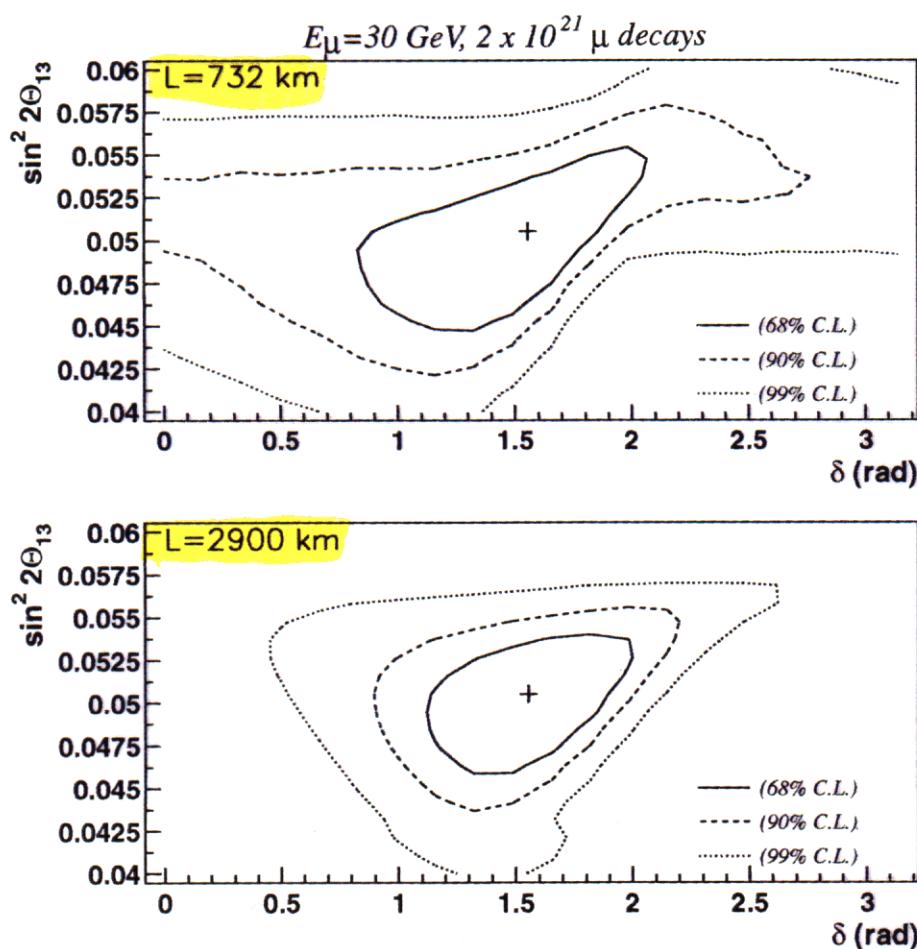


Figure 15: Correlation between  $\theta_{13}$  and CP phase  $\delta$  for two different baselines and  $2 \times 10^{21}$  decays.

Bueno et al.

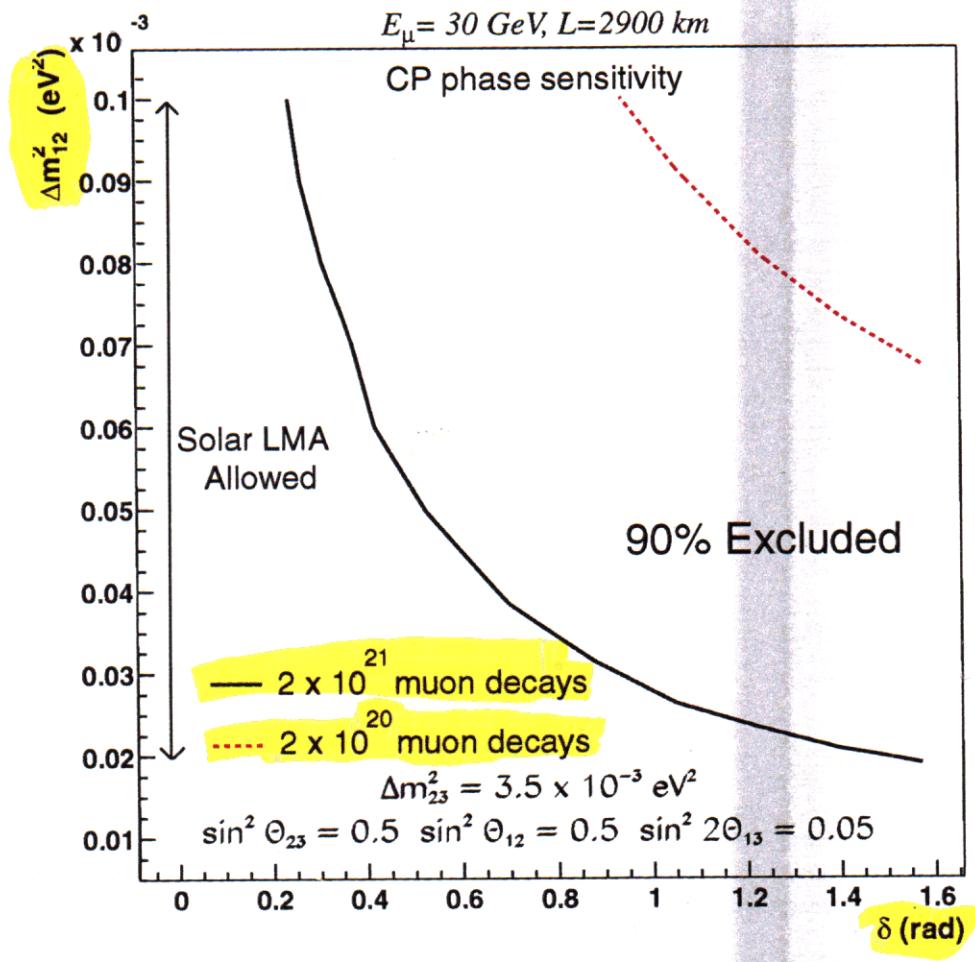


Figure 17: 90% C.L. sensitivity on the CP phase  $\delta$  as a function of  $\Delta m_{12}^2$  for two different normalizations: solid (dashed) line corresponds to  $10^{21}$  ( $10^{20}$ ) muons decays of each polarity.

Bueno et al.

Bruno et al. ALSO EMPHASIZE:

WRONG SIGN MUONS ARE OPTIMAL

TOOL, EXCEPT FOR SMALL  $\Theta_{13}$

→  $\tau$ -APPEARANCE IS BETTER

$\nu$ -FACTORY (AS B-FACTORIES)

SHOULD AIM AT OVERCONSTRAINING

THE PARAMETERS

→  $\tau$  NEEDED

$$1\nu F < \frac{1}{2} 12 \text{ BF}$$

ALL STUDIES OF  
SIGNALS, S/BACKGROUNDS

CONCUR:

$L \sim 3000$  km OPTIMAL

$L \sim 6000$  km USEFUL

(as an ~~ADD-ON~~)

$L \approx 732$  km ☹

IS NOT ENOUGH

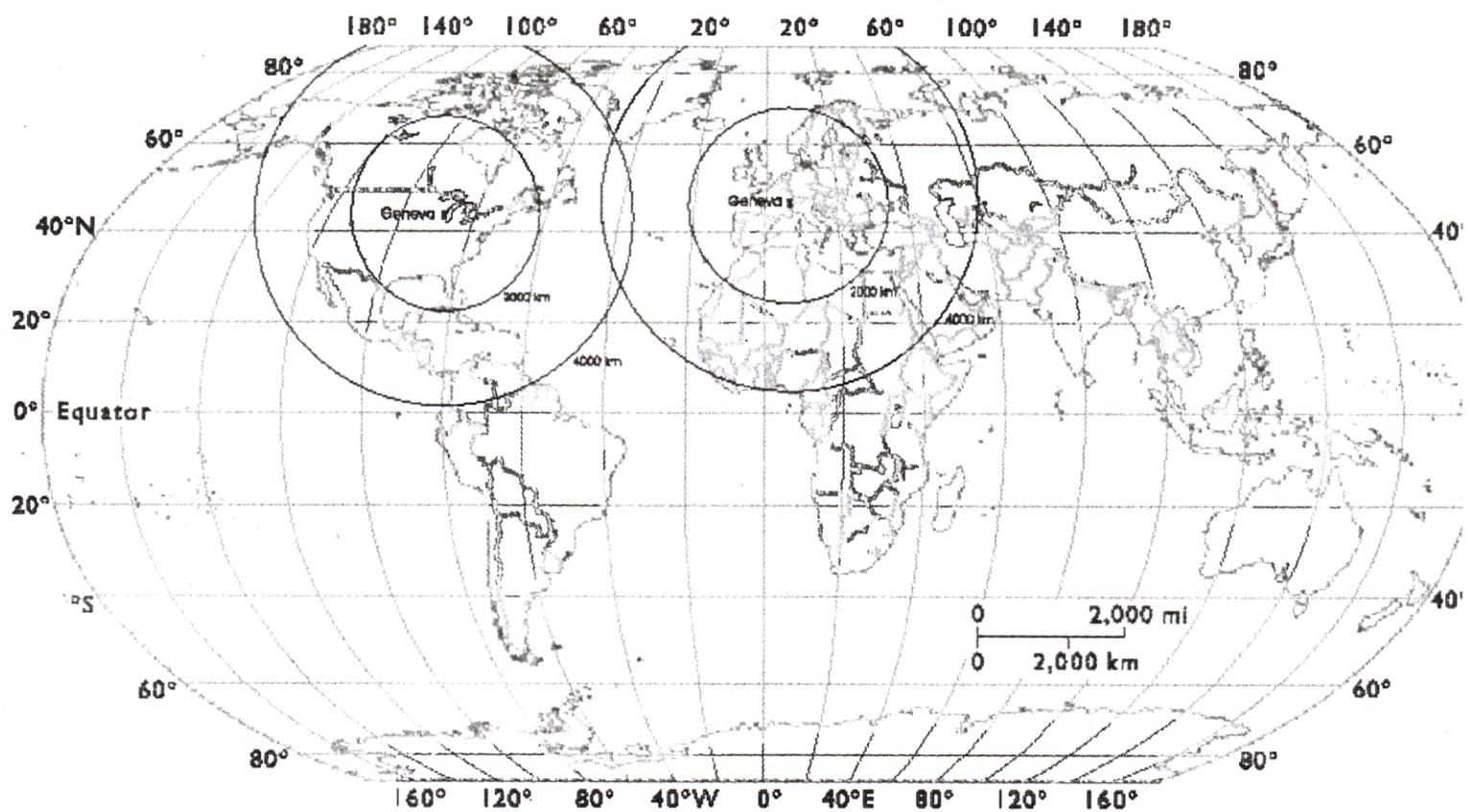
2000-4000 km CIRCLES CENTERED  
AT GENEVA (ILL) AND GENEVA (CH)



GOOD DISTANCE TO WEST COAST  
BUT

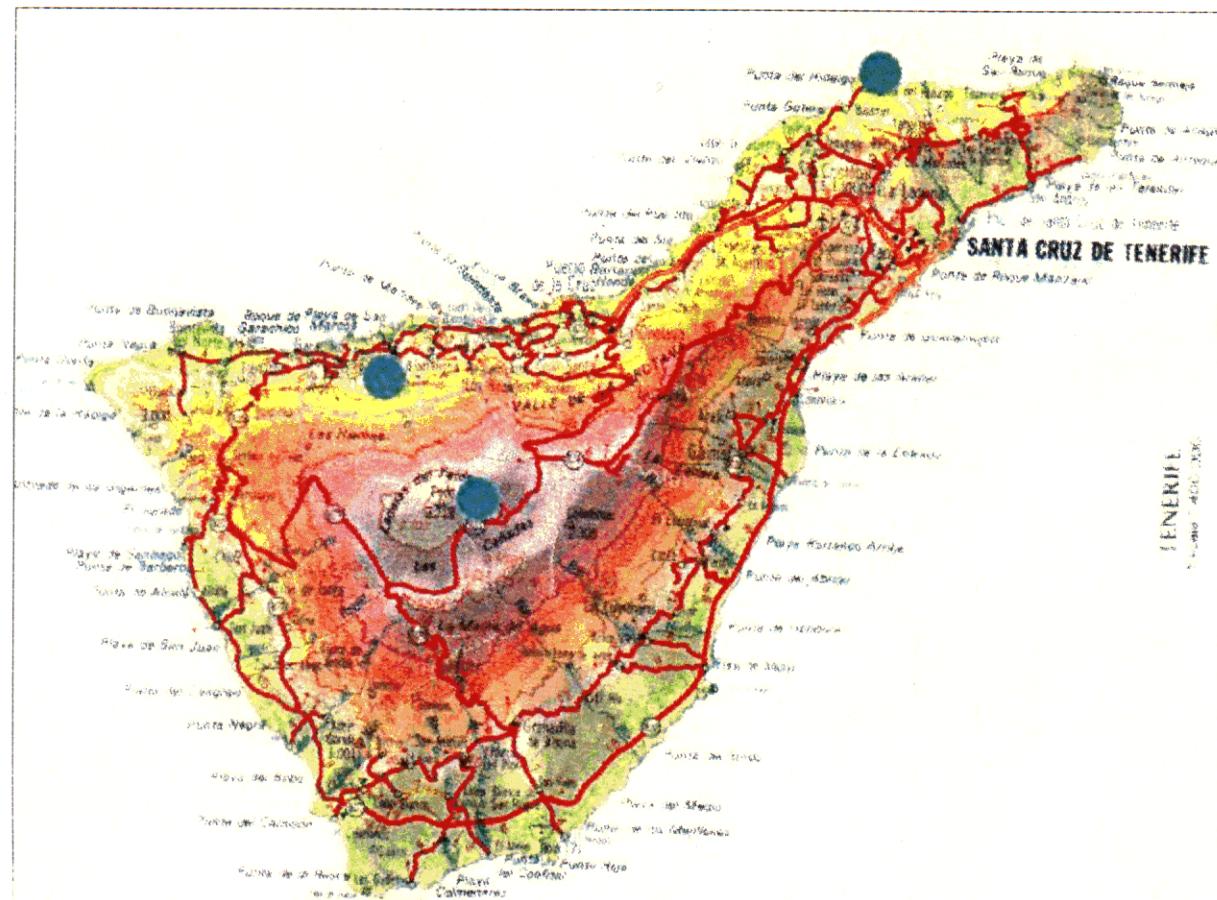
NO SINGLE, PEACEFUL, TECHNOLOGICALLY  
REASONABLE, NON-FREEZING PLACE  
WITHIN  $3000 \pm 1000$  km FROM  
GENEVA (CH)

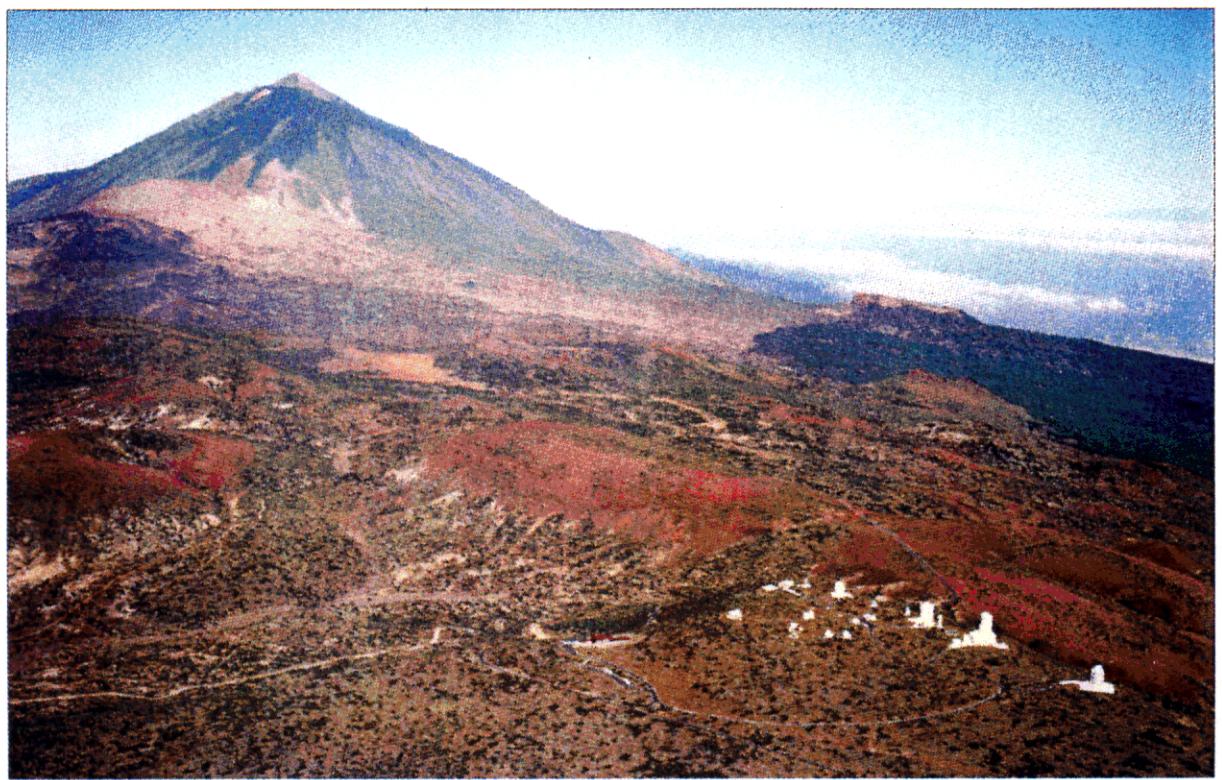
EXCEPT!



## ISLAS CANARIAS

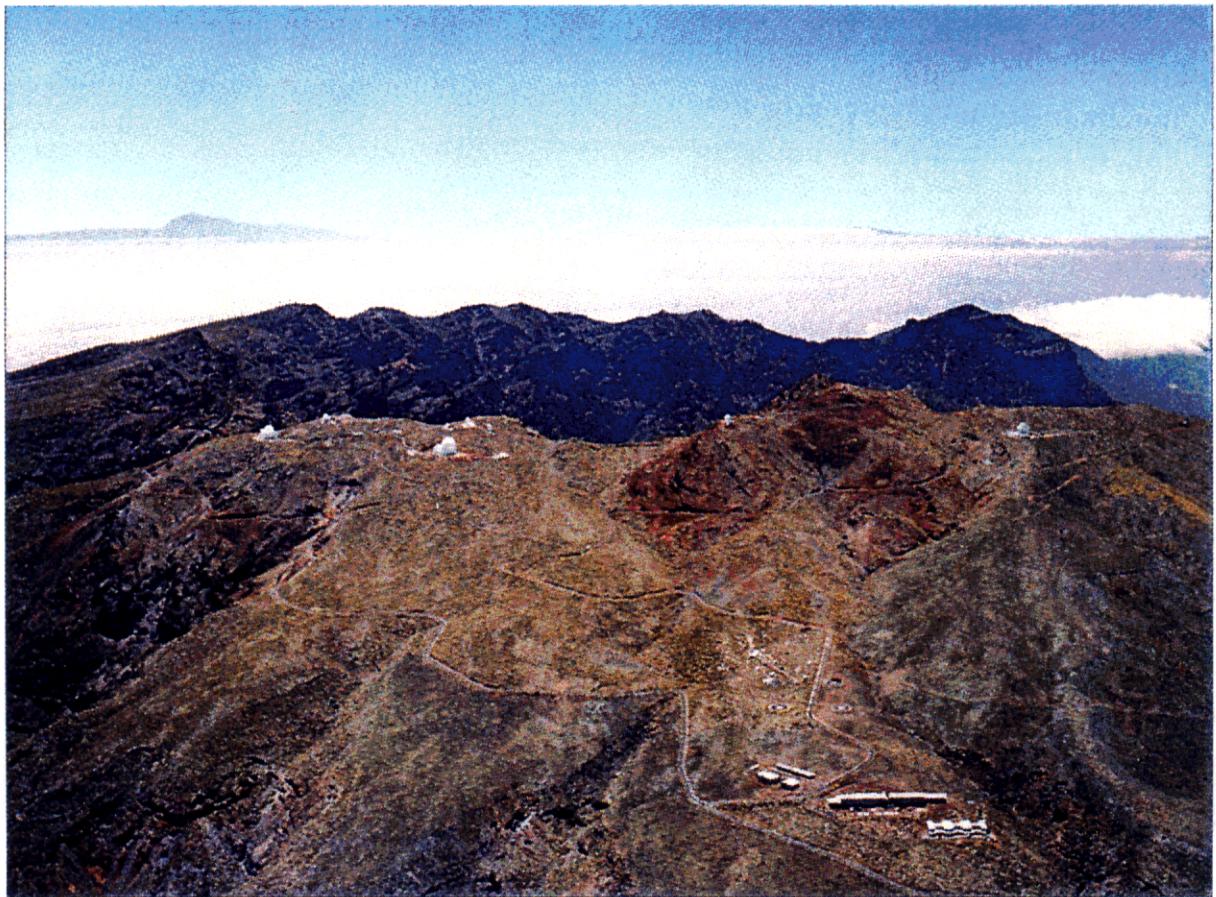






LA PALMA





km

FERMI → GRANSASSO 7332

FERMI → CANARIES ~6000

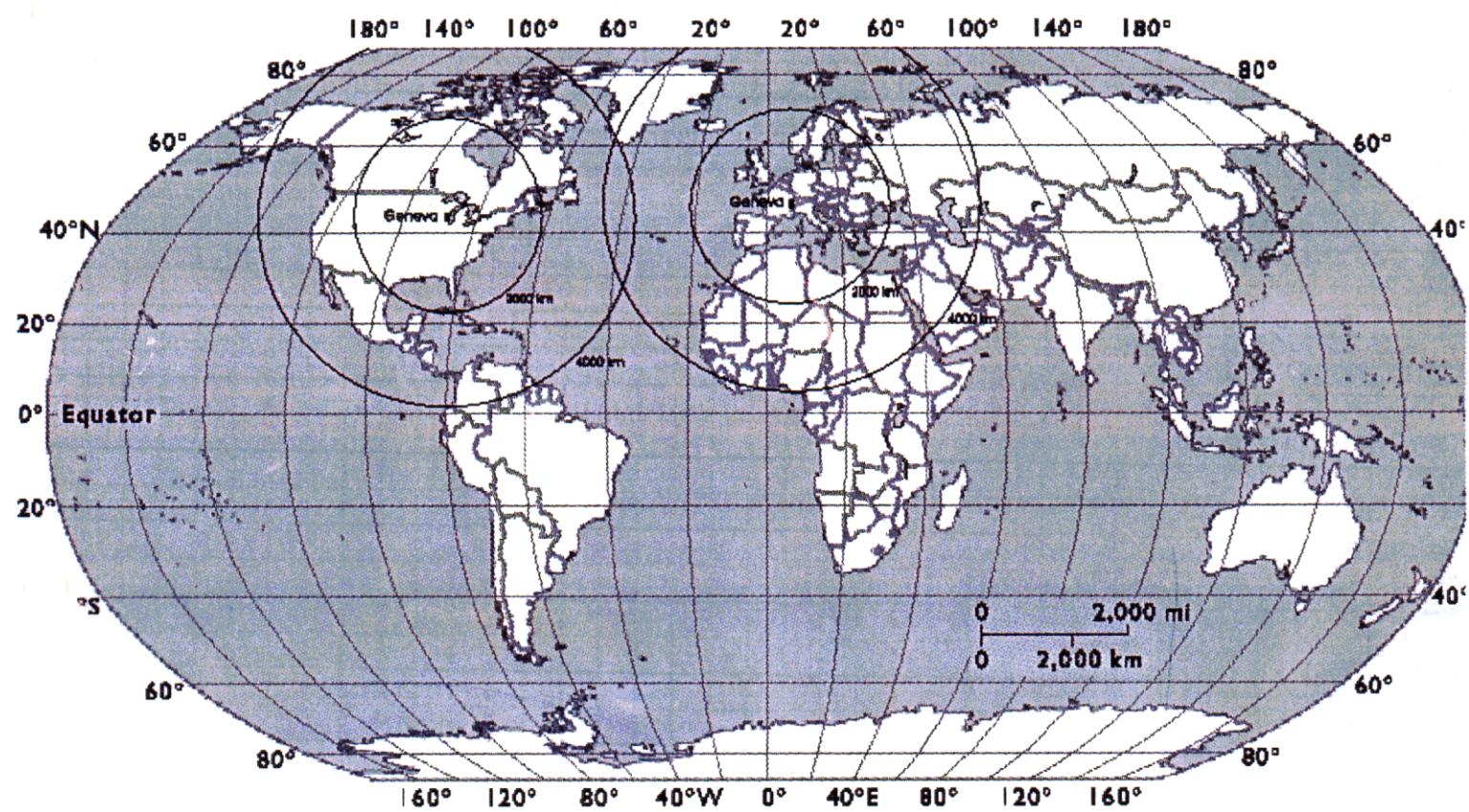
FERMI → WEST COAST ~3000  
[OPTIMAL]

CERN → GRANSASSO 732 ☹

CERN → FERMILAB ~6000

CERN → CANARIES ~3000

[OPTIMAL]



EITHER QVA  
TO A  
(MAGNETIZED)

SUPER K

~ 8500 km

**HOW  
TO  
CONVERT  
A  
D.G.?**

A.O.B.

## MATTER EFFECTS

$$\rightarrow \langle n_e \rangle |_{\nu \text{ TRAJ.}} \pm 10\%$$

FREUND et al.

FREUND et al. al.

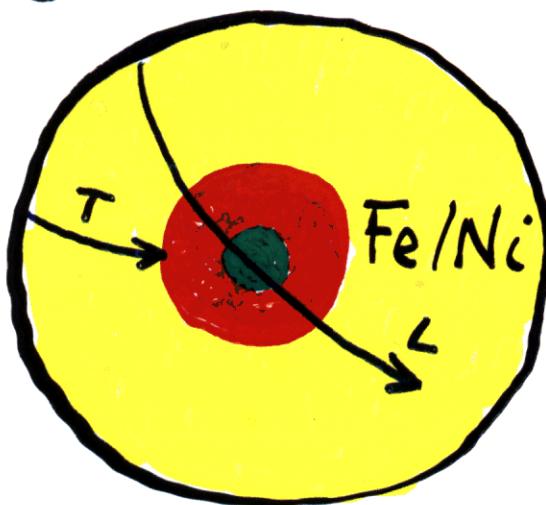
CERVERO et al.

BUENO et al.

BARGER et al.

ALBRIGHT et al.

## $\nu e$ FORWARD SCATTERING



# A.O.B.

## $\mu$ -POLARISATION

○ TURNING OFF  $\gamma_e$

CAN MODULATE S/B

○ SURROGATE FOR THE  
NEARLY IMPOSSIBLE  
 $e^+/-$  DISTINCTION

See CERN study

MORE OF A "WHY NOT"  
THAN A "MUST"

On Burt Richter's insistence

IS AN UPGRADED, ~~tl/k~~  
(CONVENTIONAL) BERM  
COMPETITIVE WITH  
A V-FACTORY



# NON-OSCILLATION $\nu$ -PHYSICS

e.g. : FERMILAB AND CERN REPORTS

CLOSE-BY LOCATION :

10's Mevents/kg /  $10^{21}$  injected  $\mu$ 's

(WITH A NATURALLY POLARISED  $\nu$ BEAM)

- $F_i, G_i(x, Q^2)$  IN p, n,  $\bar{N}$  QCD, NP
- PRECISION EW.  $\phi$ :  $\Gamma(\nu e)$   $10^5$  FOLD STAT.
- $D_0 - \bar{D}_0$  MIXING, TAGGED AT PROD.
- $\nu$  MAGNETIC MOMENTS

CATCHING SLOW BACKWARDS  $\mu$ 'S

$\mu \rightarrow e\gamma$

$\mu N \rightarrow e N$

$\mu \rightarrow eee$

$\mu^+ e^- \rightarrow \mu^- e^+$   $i$   $10^3 * \text{CURRENT}$   
 $i$   $10^3 * \text{STATISTICS}$

## Deep Inelastic Scattering Experiments

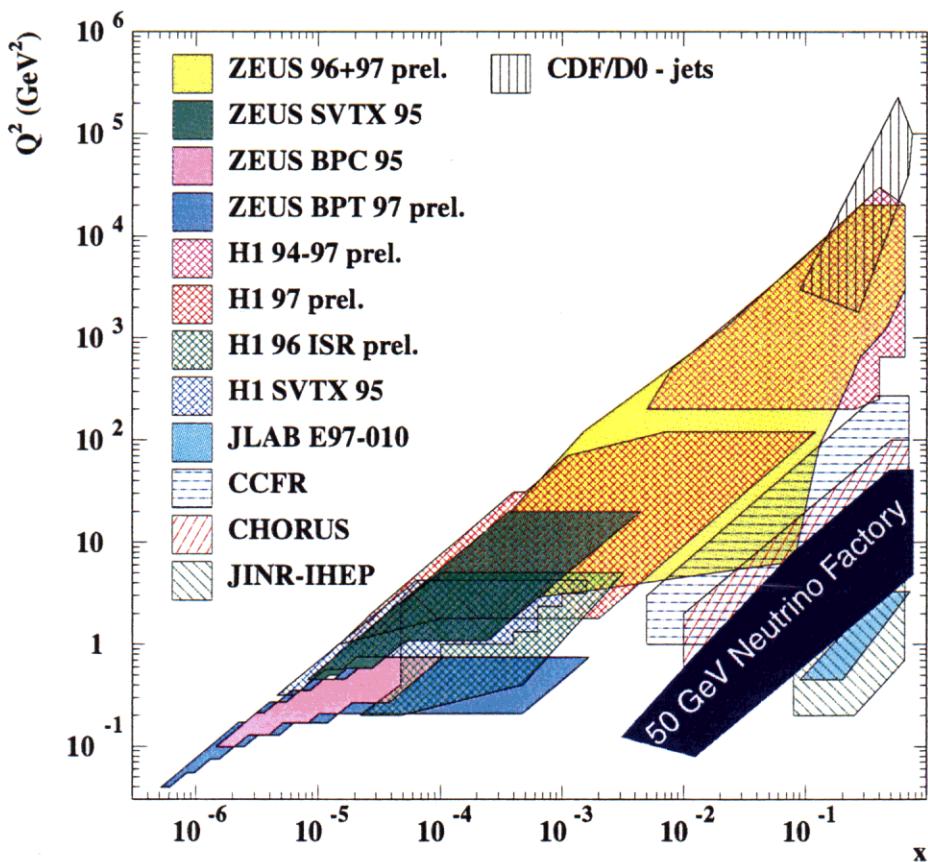


Figure 53: Comparison of kinematic ranges for present DIS experiments with a 50 GeV Neutrino factory.

FOR DECADES  
EXPERIMENTS SEEMED  
TO DO LITTLE BUT  
TO VERIFY THE  
PREDICTIONS OF THE  
STANDARD MODEL

( $N_y = 3$ , STANDARD EXCEPTION)

THIS SEEMED TO BE  
TAKING FOREVER

**FOREVER**

**IS A VERY  
LONG TIME**

**PARTICULARLY  
TOWARDS THE**

**END**

**U-OSCS. BROUGHT US  
BEYOND THE END**

**(OF BORING PHYSICS)**

$D \rightarrow (K, \pi) \mu \nu_\mu$

C  
↓

$$S_0 \sim s \cos \theta_c - d \sin \theta_c$$

DECOHERENCE





$$\nu_\mu = \alpha \nu_1 + \beta \nu_2 + \gamma \nu_3$$

COHERENT OSCILLATIONS

NATURE

HAS ALSO TUNED MANY OTHER PARAMS.  
FOLLOWING THE PRINCIPLE OF MINIMAL EFFORT (FOR PHYSICISTS)

○  $m_\nu$ 's

○  $E(\text{CRs})$ ,  $R_\odot$ ,  $P_\odot$

○  $E(\text{WRS})$ ,  $R_\odot$ ,  $P_\odot$

$d(\Theta)$  ?

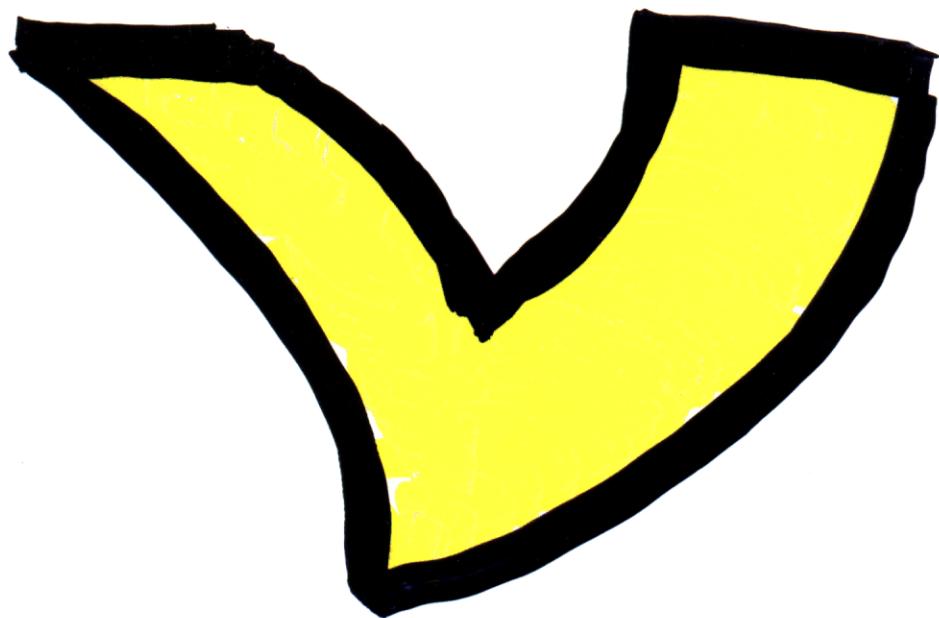
HOW SHOULD WE  
RECIPROCATE  
NATURE'S LOVING  
GENEROSITY

AND DO THE  
RIGHT THING

BY HER

???

# BUILDING A



# FACTORY

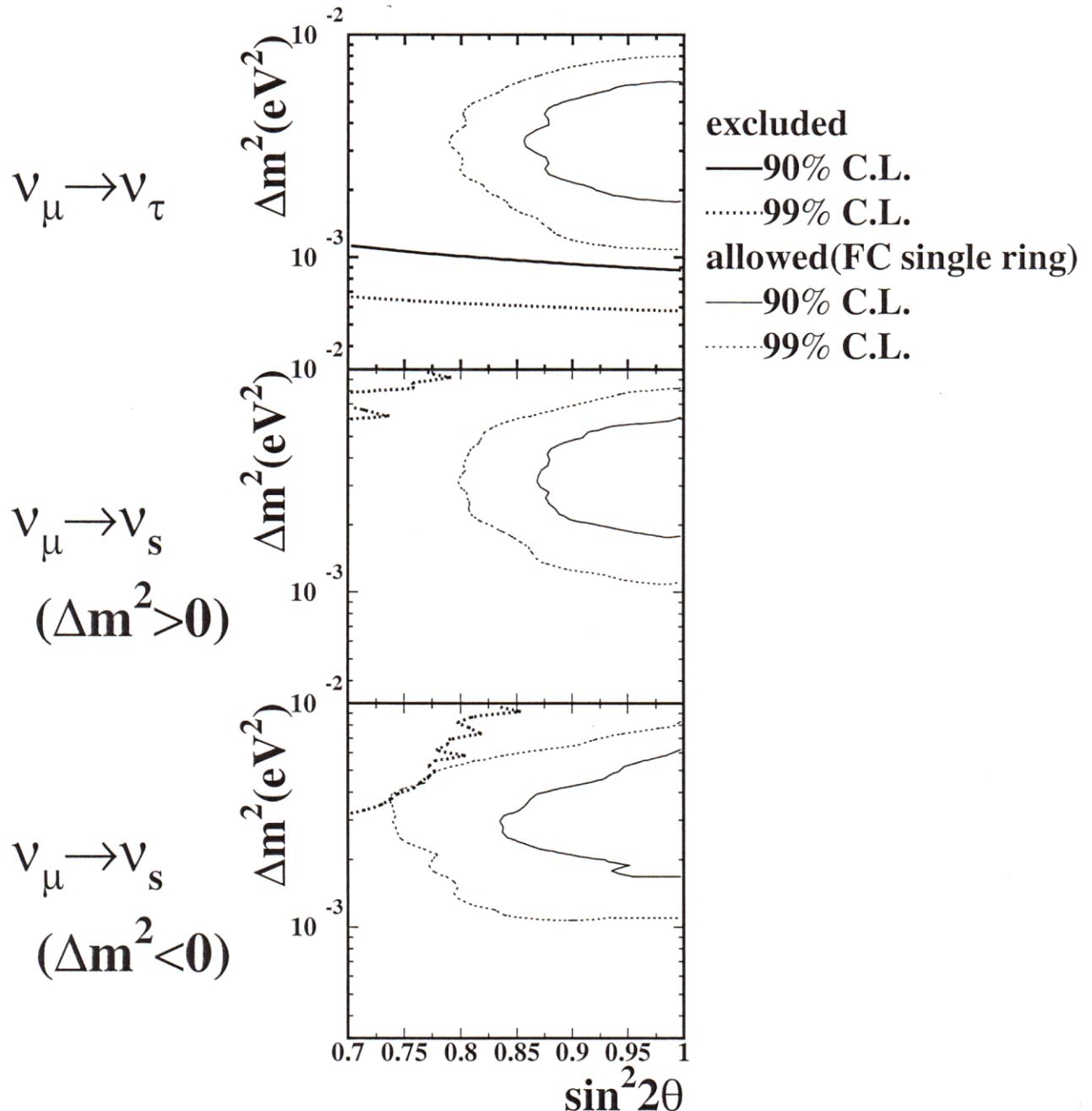
OTHERS e.g ICA<sup>RUS</sup>  
NOE may be orders of  
magnitude more sophisticated but are  
moving ahead in testing, tests and development  
stage,

ref?

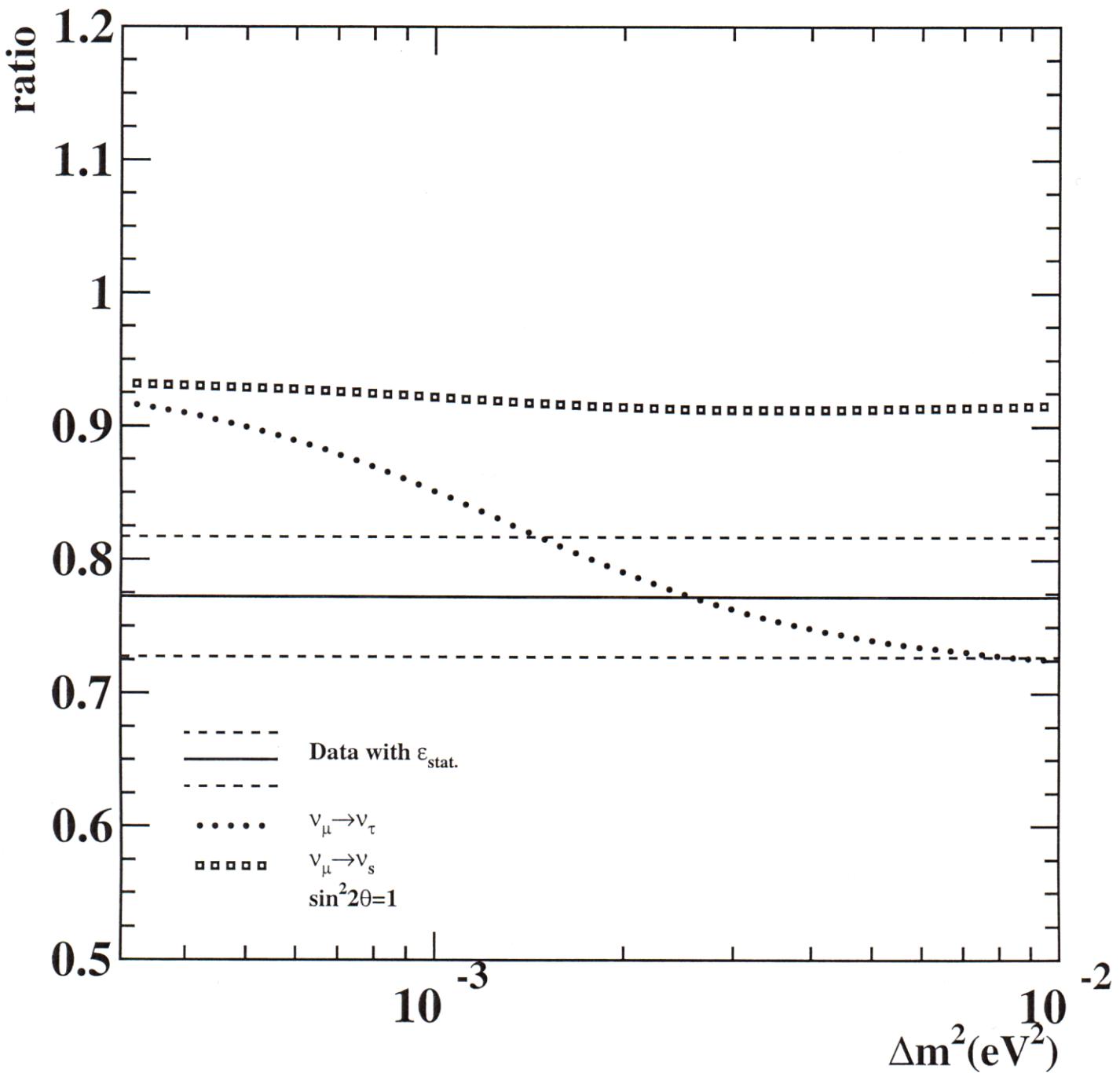
Some of these detectors are so  
large that they do not fit  
in a transparency

but they are quite simple quite simple  
(the Majorana  $\bar{F}_e$ , segmented simulation)

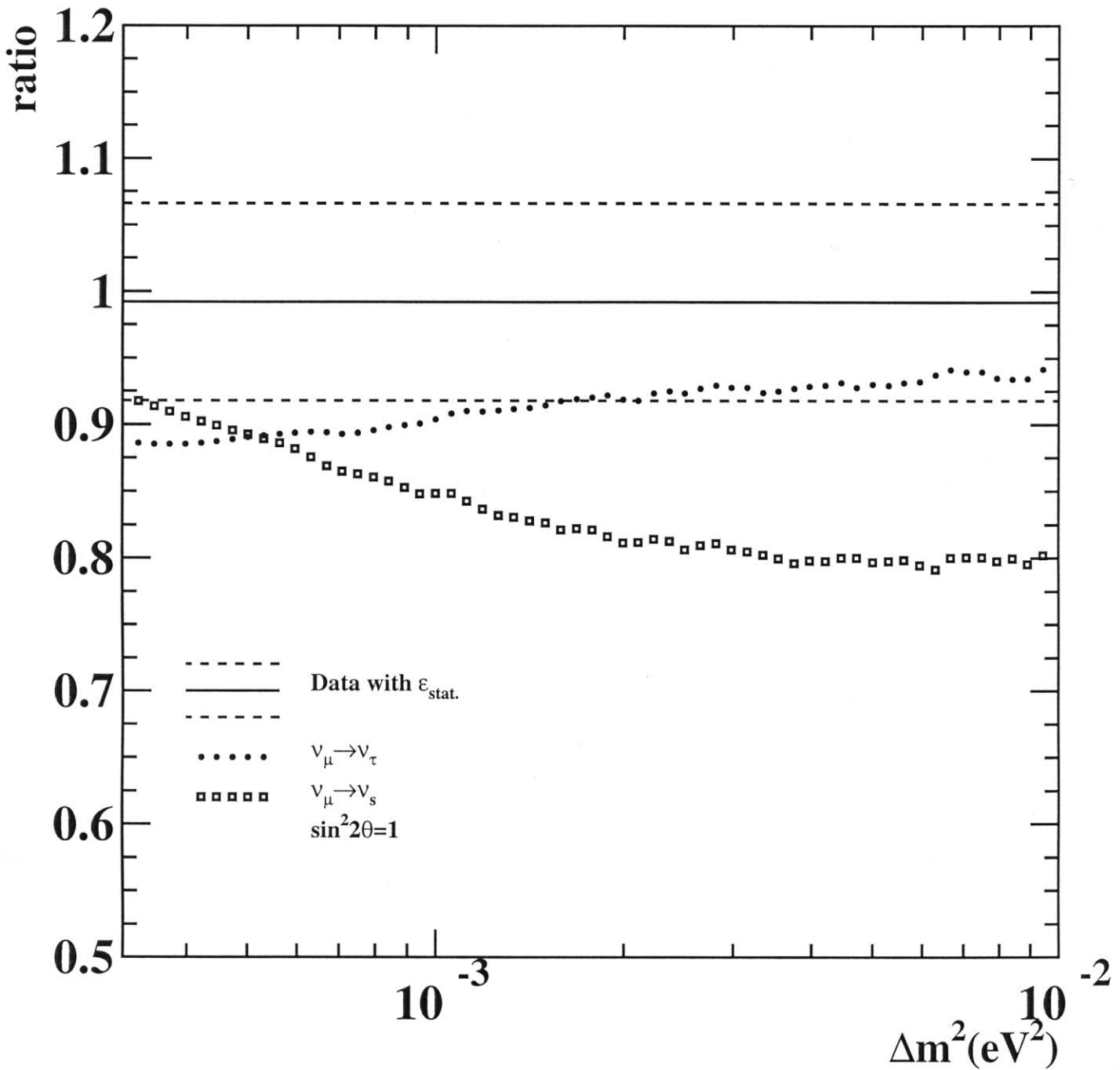
excluded region from combined analysis(multi+PC+up $\mu$ )



vertical/horizontal ratio of upward through going  $\mu$  events



## up/down ratio of N.C. enriched multi-ring events



## Deep Inelastic Scattering Experiments

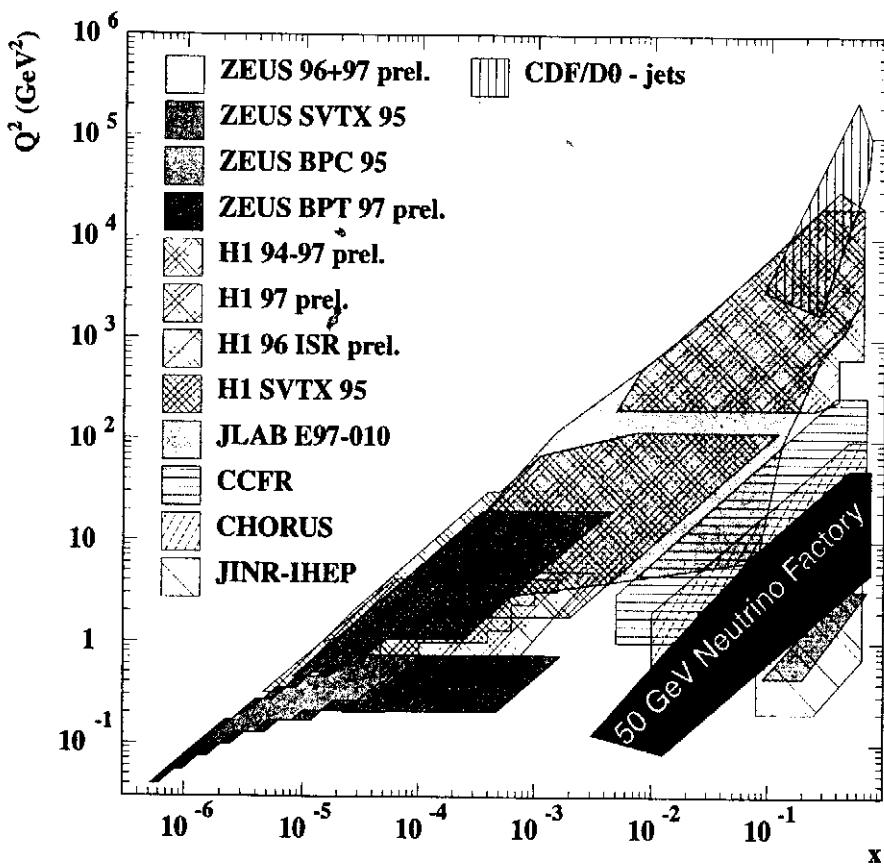
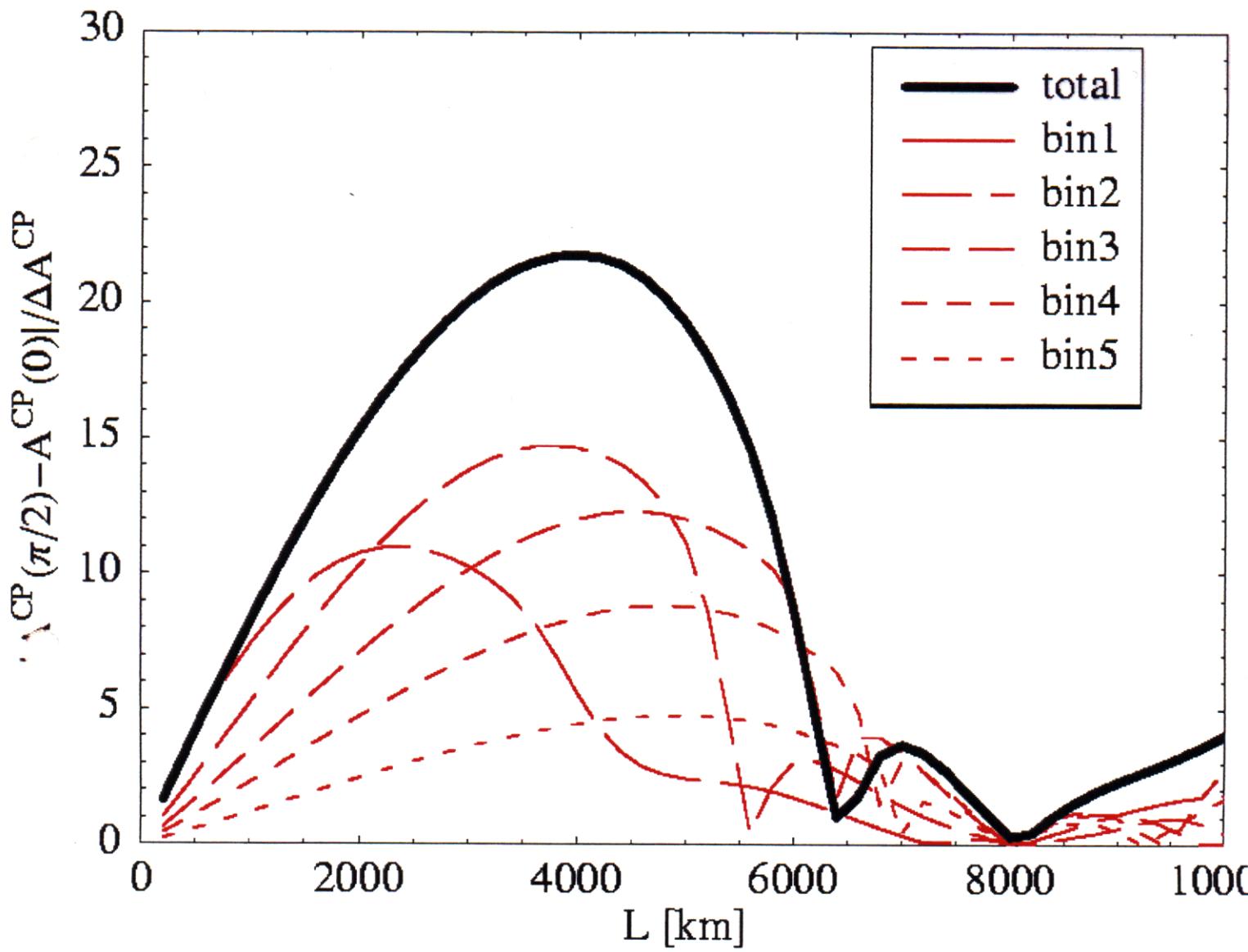


Figure 53: Comparison of kinematic ranges for present DIS experiments with a 50 GeV Neutrino factory.



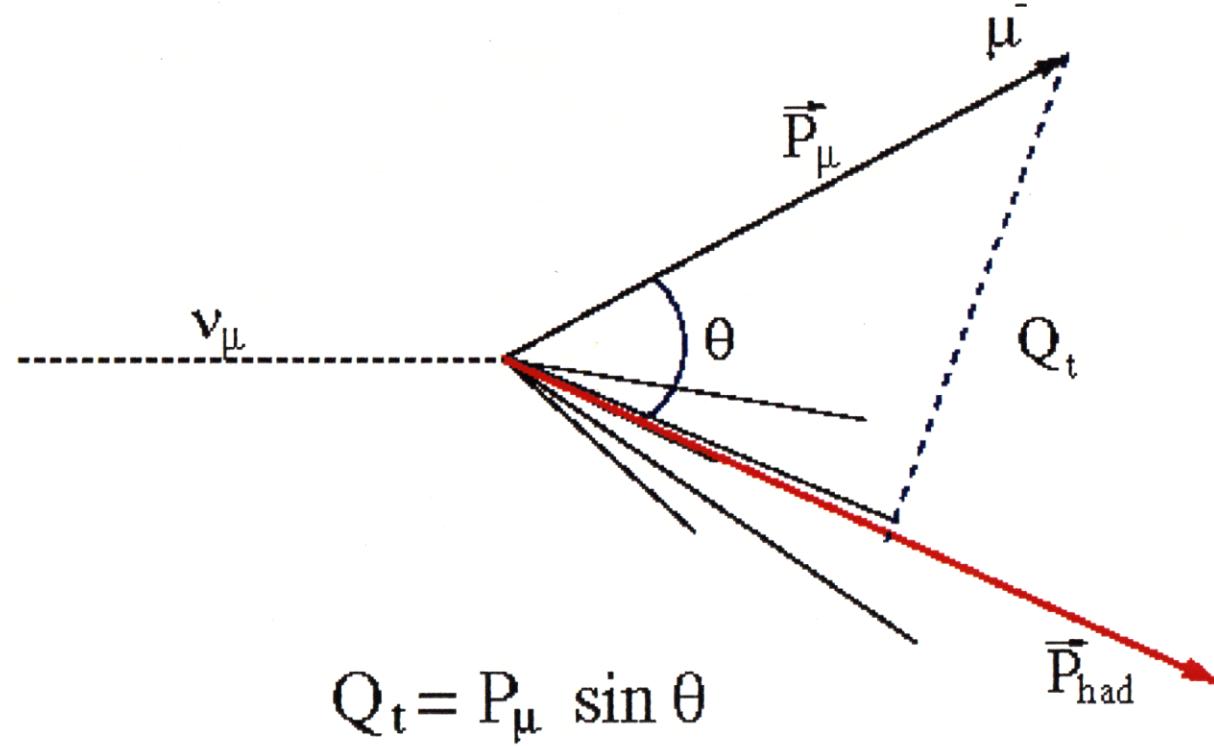


Figure 8: *Definition of the kinematical variables used in this study.*

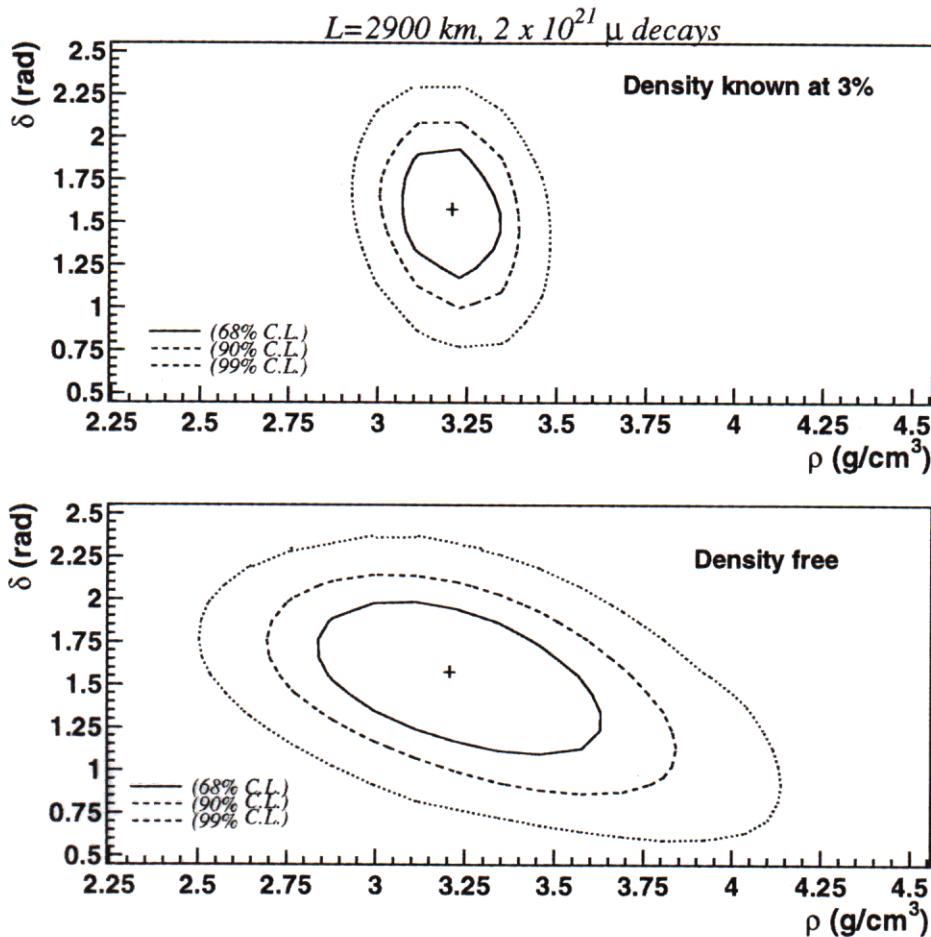


Figure 16: Correlation between the average matter density  $\rho$  and the CP phase  $\delta$  for  $L=2900 \text{ km}$ . In the lower plot we leave Earth's mean density as a free parameter in the fit. In the upper plot we assume that density is known within 3%. We see that the presence of matter does not spoil the possibility of performing a measurement of  $\delta$ .

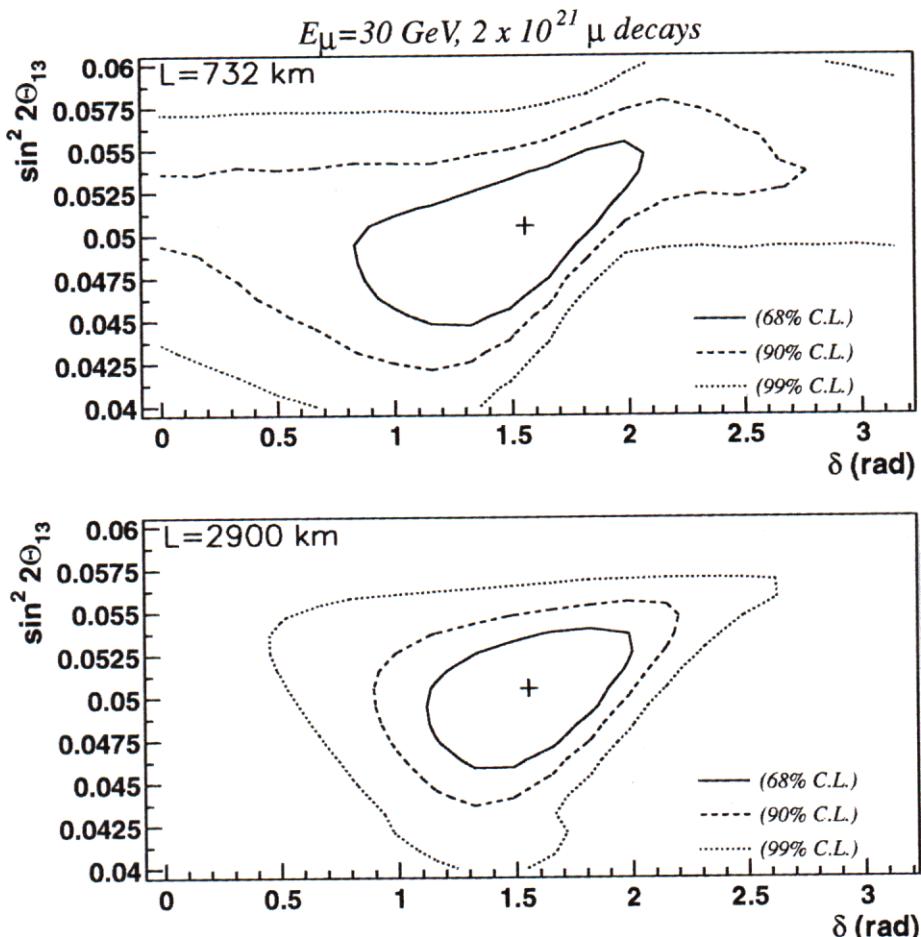


Figure 15: Correlation between  $\theta_{13}$  and CP phase  $\delta$  for two different baselines and  $2 \times 10^{21}$  decays.

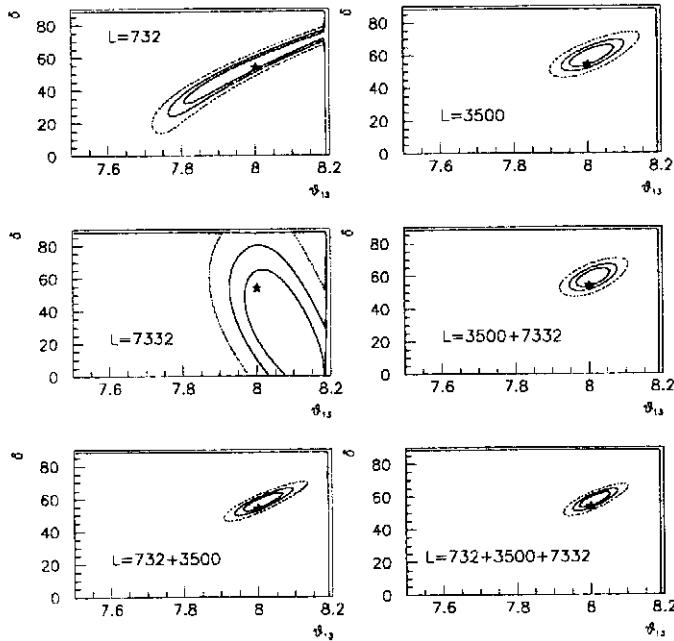


Figure 16: 68.5, 90, 99 % CL contours resulting from a  $\chi^2$  fit of  $\theta_{13}$  and  $\delta$ . The parameters used to generate the “data” are depicted by a star and the baseline(s) which is used for the fit indicated in each plot. Only statistical errors are included.

