

Neutrino Factory Physics Update

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Fermilab

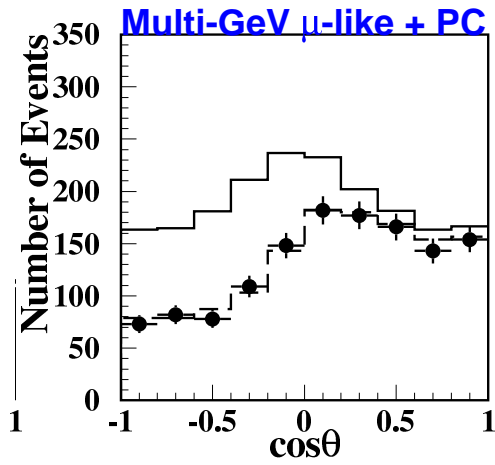
MUTAC Meeting
April 28-29, 2004
Brookhaven National Laboratory

Outline

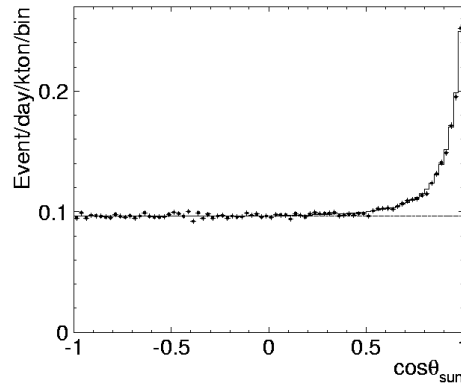
- Intro—what do we know, what do we want to know?
- ν beam choices
- Detector Choices
- From Probabilities to Mixing Angles
- Conventional ν Beam Proposals
 - Broad band, narrow band
 - First Maximum, Second Maximum
 - Combining Conventional Experiments
- ν Factory
 - Ultimate reach in $\sin^2 2\theta_{13}$
 - Combining ν factory measurements

From Discovery to Confirmation:

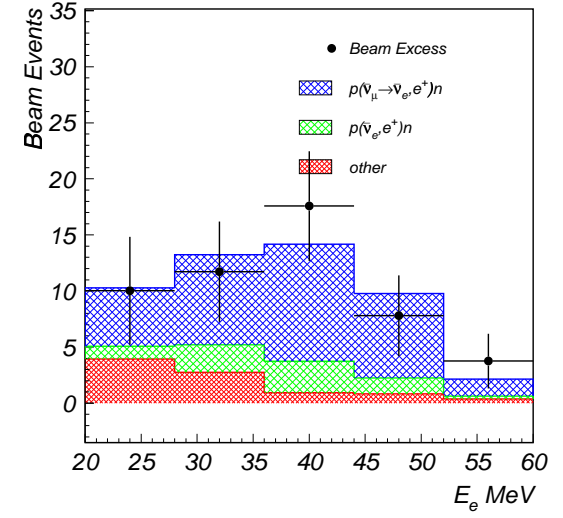
SuperK:



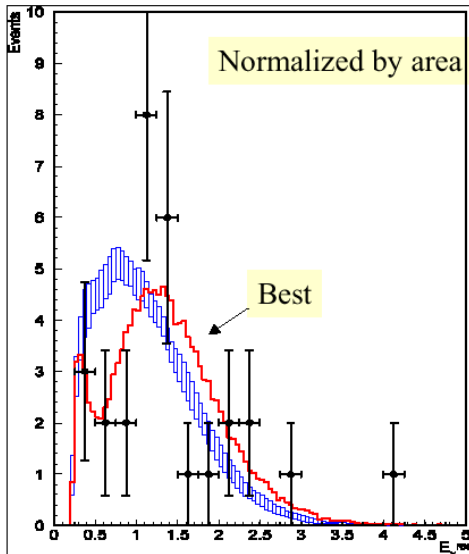
SuperK:



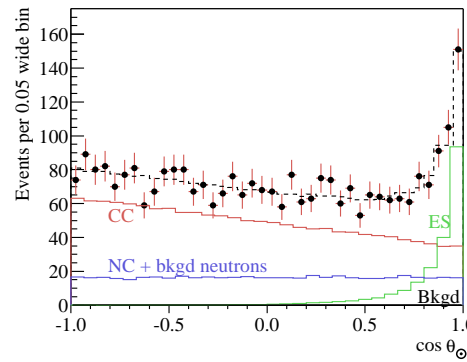
LSND:



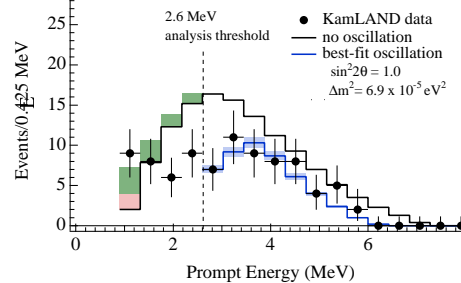
K2K:



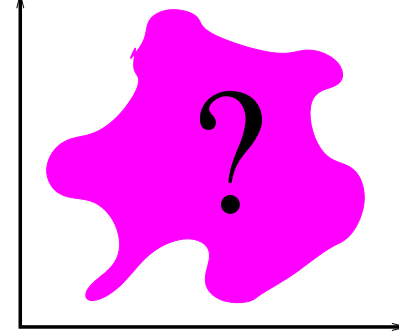
SNO:



KamLAND:



MiniBooNE:



From Confirmation to a Framework...

While MINOS and more KamLAND and SNO data take us from confirmation to precision, framework is being developed...

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \mathbf{U} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Produce Weak Eigenstate...
Propagate Mass

Detect Weak Eigenstate...
Eigenstate

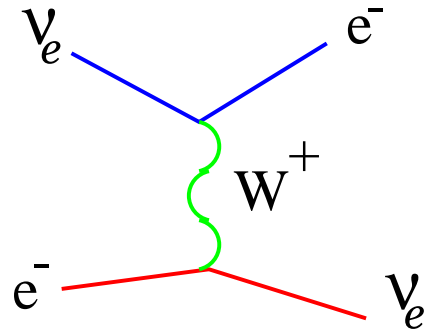
if $s_{ij} = \sin \theta_{ij}$, $c_{ij} = \cos \theta_{ij}$, then

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{-i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

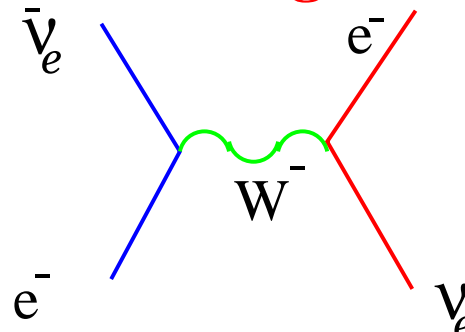
$$U = \begin{pmatrix} \text{Earth} \end{pmatrix} \begin{pmatrix} \nu_e \longleftrightarrow \nu_\mu \\ \text{at} \\ \text{atmospheric} \\ \Delta m^2 \end{pmatrix} \begin{pmatrix} \text{Sun} \end{pmatrix}$$

θ_{13} and δ : More than just an angle and phase...

• Are Neutrino Masses in the “wrong” order?



Coherent Elastic Scattering: $\nu_e \bar{\nu}_e$ only!

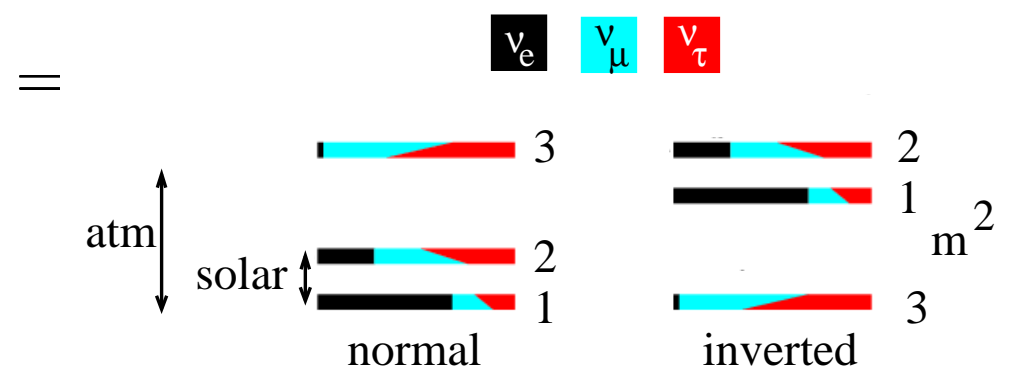


Changes ν_e oscillation probability

Wolfenstein, Phys. Rev. **D17** (1978)

$$\frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)} = \frac{2E_\nu}{E_R} \text{ for low } E_\nu$$

$$E_R = \frac{\Delta m_{atm}^2}{2\sqrt{2}G_F\rho_e} \approx 11\text{GeV}$$



from Mena and Parke, hep-ph/0312131

• CP violation → This could be what brings us all here...

Designing a Neutrino Experiment

- current designs: pin down (or eliminating) Δm^2
- next generation: Discovery of a non-zero θ_{13}
- from discovery to confirmation to precision...

→ $P(\nu_\mu \rightarrow \nu_e) = A_\pm \sin^2 2\theta_{13} \pm B_\pm \sin \theta_{13} \sin \delta + \dots$

→ CP Violation without matter effects:

$$\frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)} \propto \frac{\Delta m_{sol}^2 L}{E} \frac{\sin \delta}{\sin \theta_{13}}$$

→ Matter effects without CP violation:

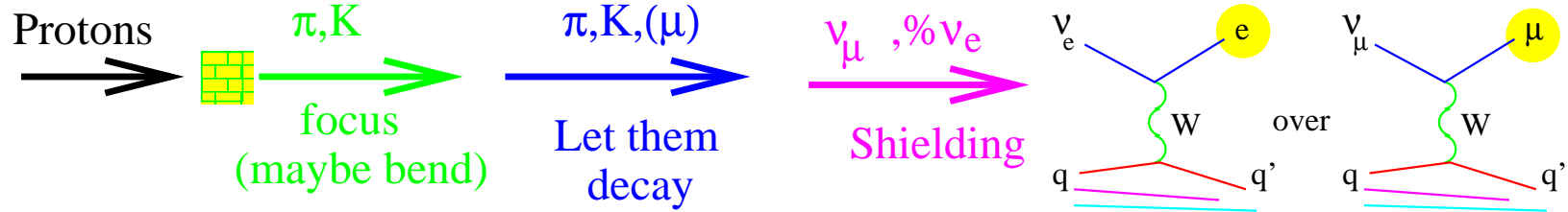
Because earth is filled with electrons...

$$\frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)} = \frac{2E_\nu}{E_R} \quad \text{for low } E_\nu$$

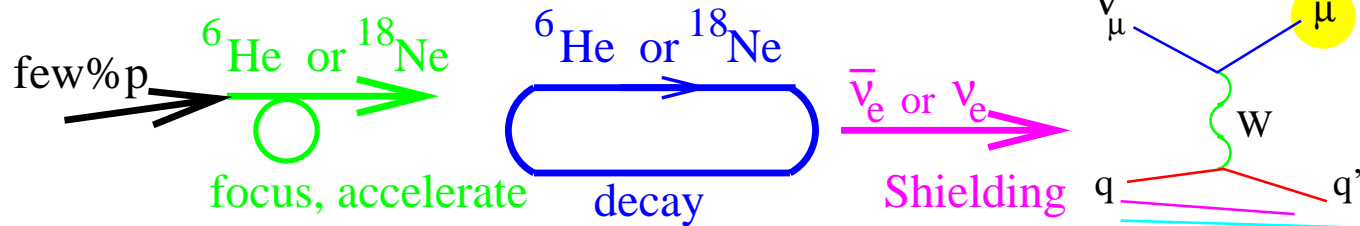
$$E_R = \frac{\Delta m_{atm}^2}{2\sqrt{2}G_F\rho_e} \approx 11\text{GeV}$$

Neutrino Beam Choices

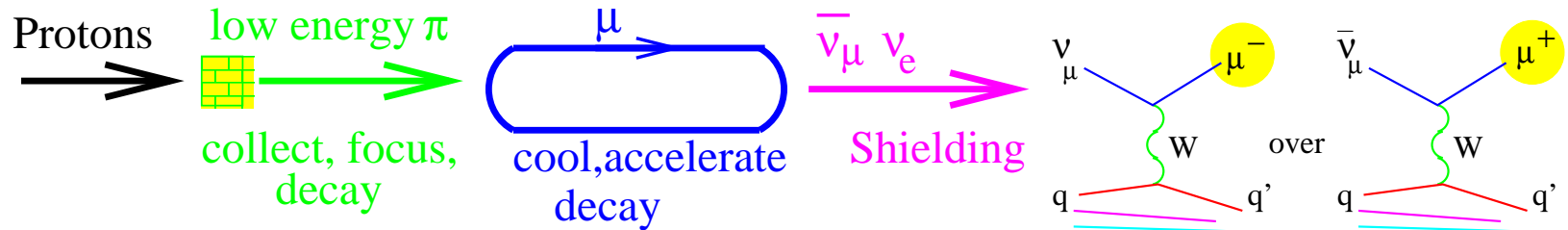
- Conventional Beam



- Beta Beam



- Neutrino Factory

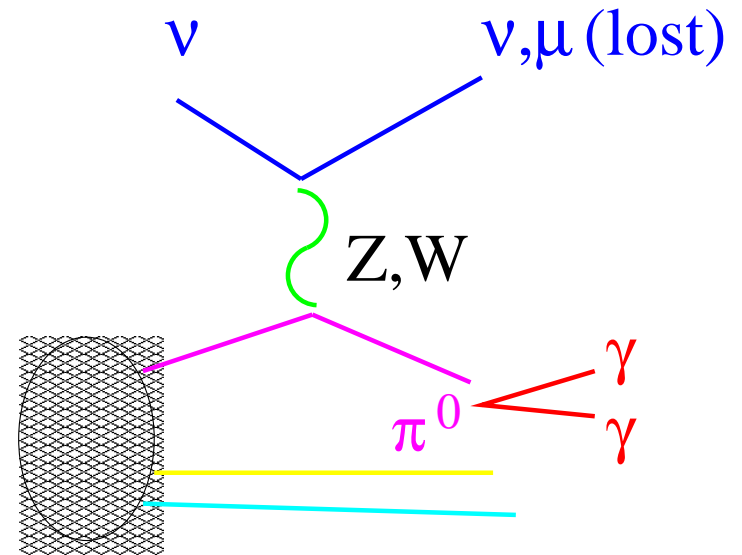


Events/parent at far detector:

$$N \propto \frac{\gamma m_{parent} K}{4\pi L_{det}^2} \frac{4\gamma^2}{(1 + \gamma^2 \theta_{\nu_{parent}}^2)^3}$$

Why is $\nu_\mu \rightarrow \nu_e$ hard? It's the Detector...

- We already know it's $< 5\%$ effect (CHOOZ)
- Unavoidable ν_e contamination of $\mathcal{O}(\%)$
- Can mistake π^0, μ, π^\pm for e^-

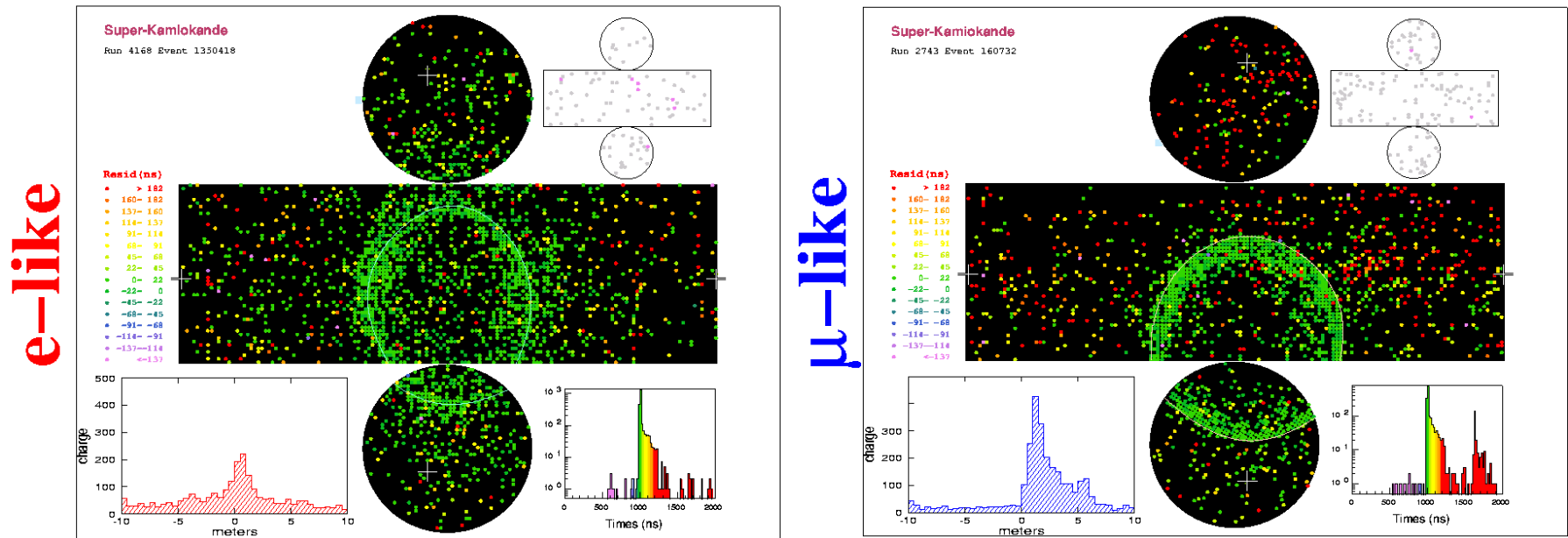


Why is $\nu_e \rightarrow \nu_\mu$ hard? It's the Beamline...

- Have to make a ν factory or β -beam

Why is $\nu_e \rightarrow \nu_\tau$ hard? Beamline and Detector

Detector Choices: Water Cerenkov



Courtesy Mark Messier

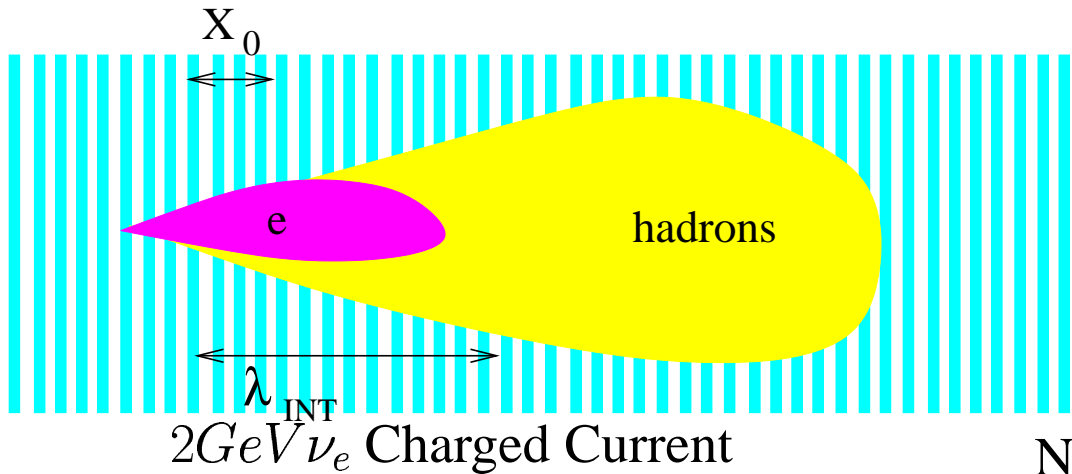
Extremely good separation for Single Particle Events

ν Quasielastic Energy Reconstruction works very well...

$$E_\nu = \frac{m_N E_\ell - m_\ell^2/2}{m_N - E_\ell + p_\ell \cos\theta}$$

But this equation fails for inelastic processes: enter backgrounds...

Detector Choices: Fine-Grained Calorimetry

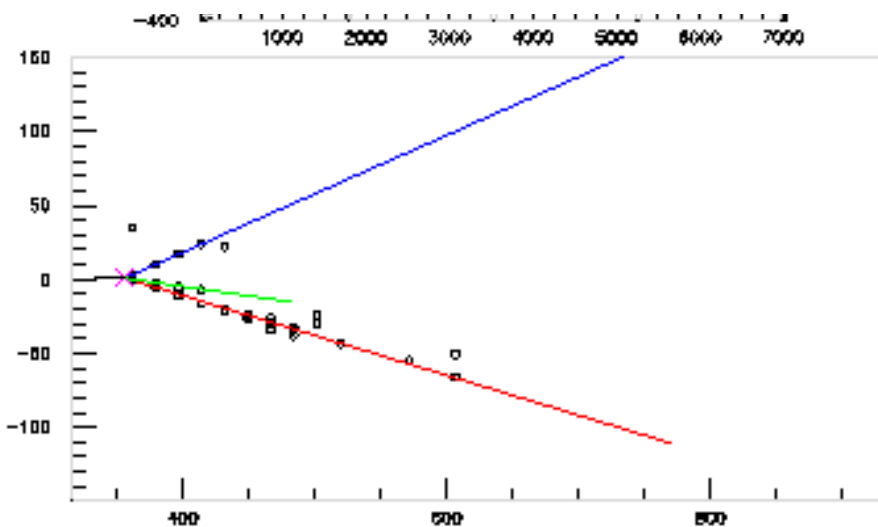


Tried and True Technology

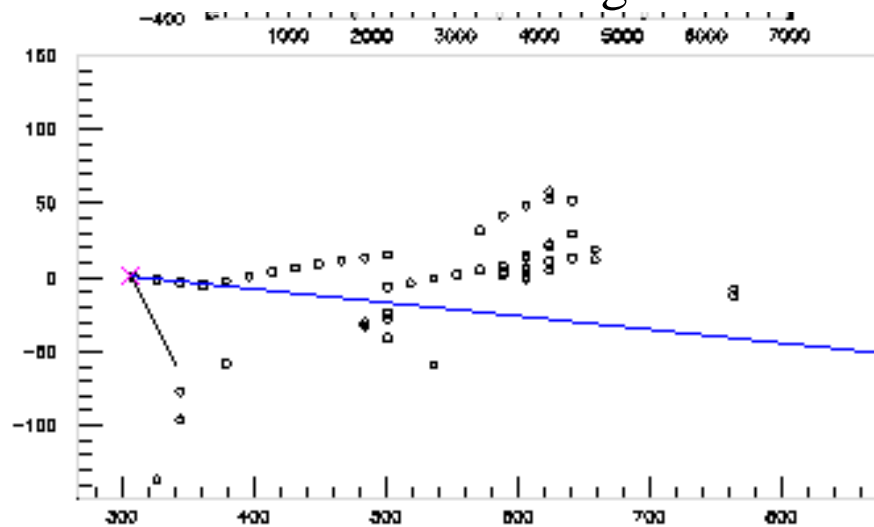
$1/3 X_0$ sampling calorimeter

Hit counting:

$$\sigma(E)/E = 12 - 15\% \text{ } 2\text{GeV}$$

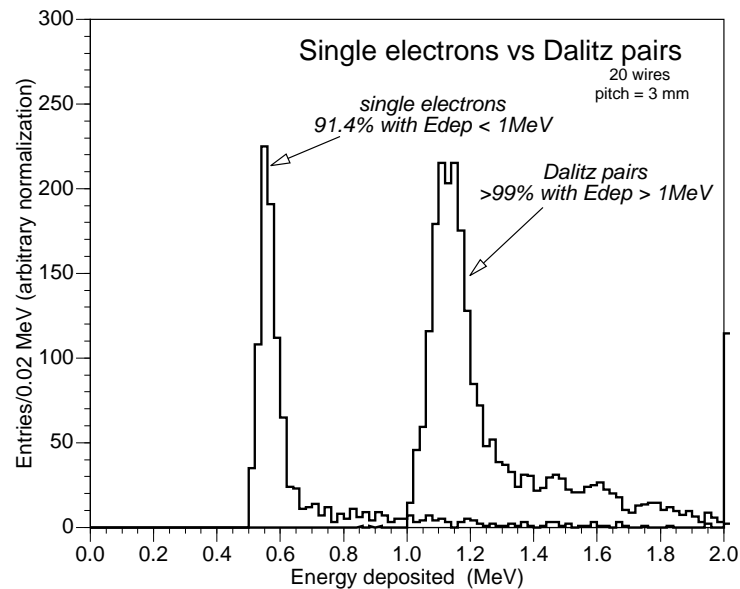
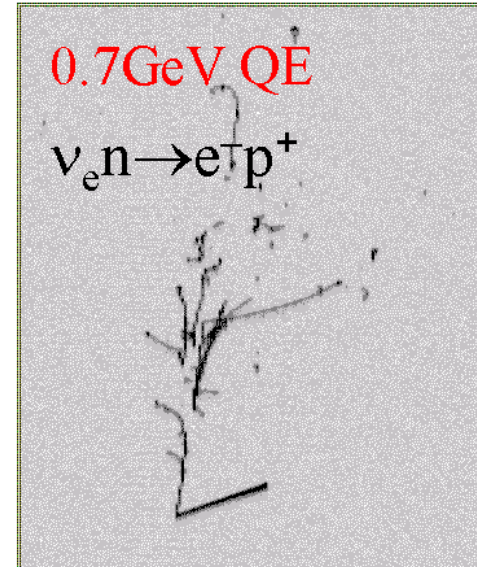
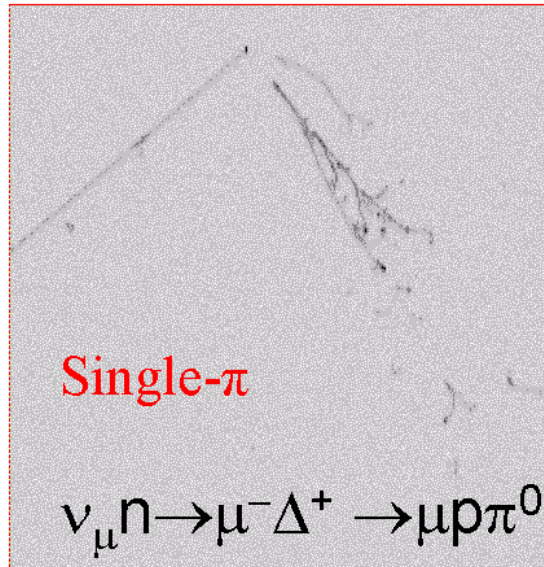


Neutral Current Background



few $\times 10^{-3}$ NC rejection with 35-40% signal acceptance at 2GeV
 Remaining backgrounds: CC and NC π^0 production

Detector Choices: LAr TPC



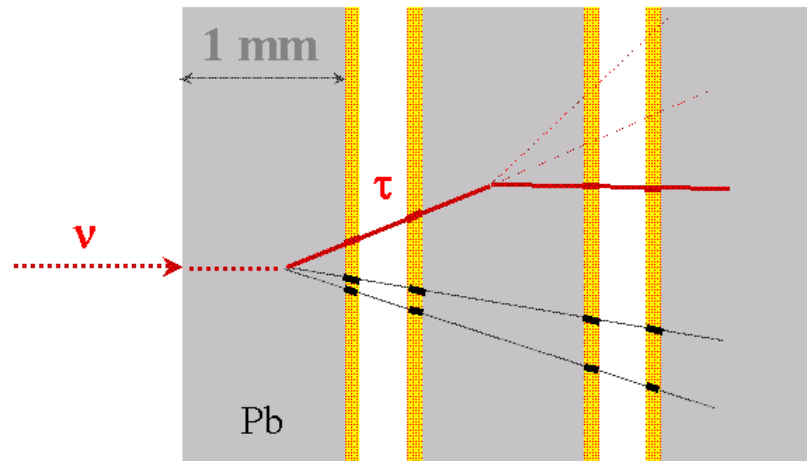
Excellent PID and energy reconstruction
Ref: ICANOE proposal LNGS P21/99
and

www.aquila.infn.it/icarus/
Events courtesy A. Rubbia

By far the smallest detector-related
backgrounds here, but... by far the most
technically challenging detector

Emulsion Detectors for τ Appearance

The Emulsion Cloud Chamber (ECC)

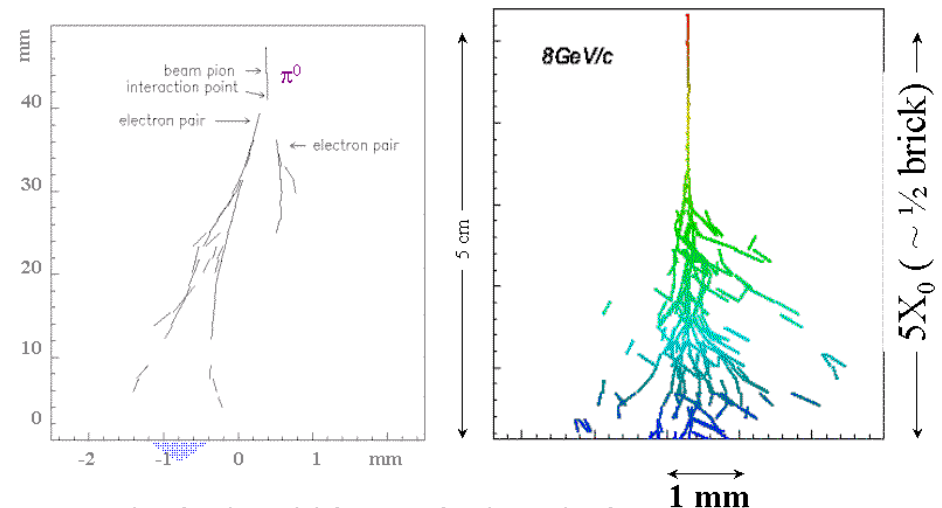


↳ $< \mu\text{m}$ space res. Emulsion layers
 ↓
 track segments

- What if MiniBooNE Confirms LSND??
- What about measuring $\nu_e \rightarrow \nu_\tau$?
- OPERA Design:
 Pb-Emulsion Sandwich

Event reconstruction with an ECC

ECC exposure at CERN-PS



Topological and kinematical analysis event by event

- Performance measured in e^- test beam at CERN
- Want to test this in known ν beam too!

In Praise of Near Detectors

If you can't remove all the backgrounds...

- measure them precisely in a near detector
- understand processes well enough to make far detector prediction

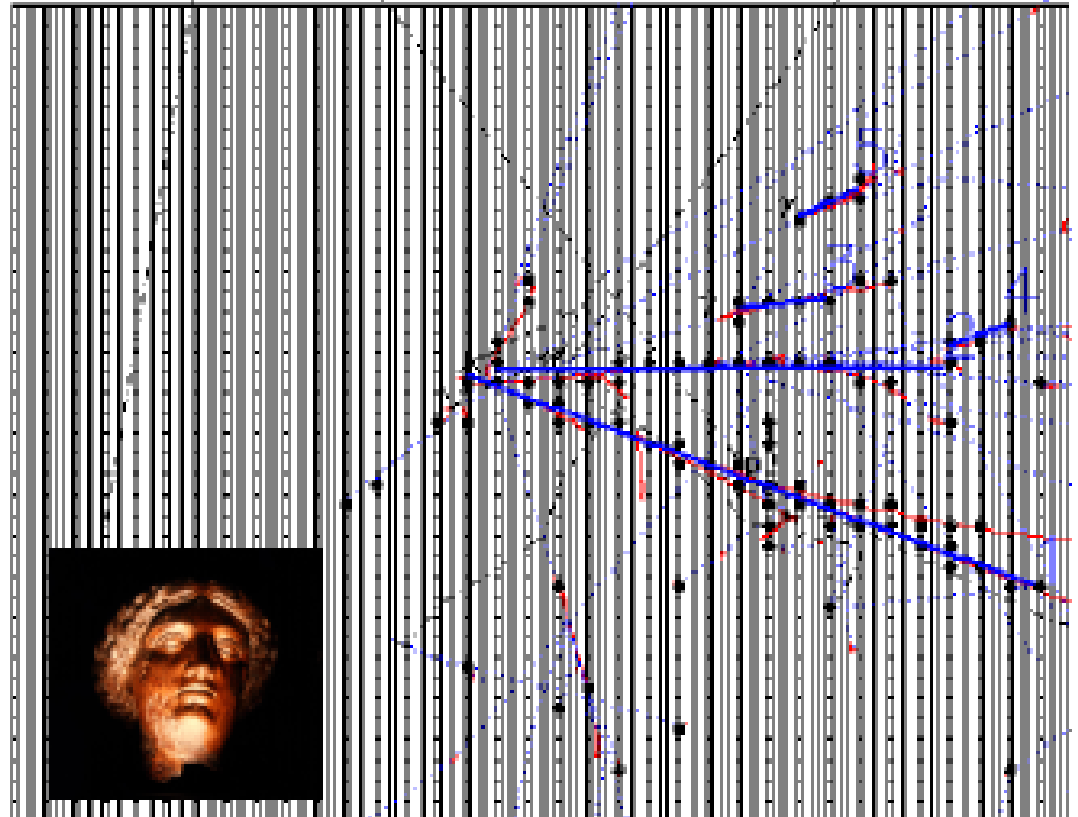
One example:

MINER ν A in NuMI/MINOS
joint Nuclear/Particle project to

- measure cross sections
- validate models based on
nuclear physics data
(e.g. JLab CLAS, Hall C)

[http://
www.pas.rochester.edu
/minerva](http://www.pas.rochester.edu/minerva)

MINER ν A Event (side view, active detector)



Conventional Beamlines will require better knowledge of cross sections...then again
so does nuclear physics...

Getting to θ_{13} in 3 generations... $P(\nu_\mu \rightarrow \nu_e) = \sum_{i=1,4} P_i$

$$P_1 = \sin^2 \theta_{23} \sin^2 2\theta_{13} \left(\frac{\Delta_{13}}{B_\pm} \right)^2 \sin^2 \frac{B_\pm L}{2} \quad \text{atmospheric part}$$

$$P_2 = \cos^2 \theta_{23} \sin^2 2\theta_{12} \left(\frac{\Delta_{12}}{A} \right)^2 \sin^2 \frac{AL}{2} \quad \text{solar part}$$

$$P_3 = J \cos \delta \left(\frac{\Delta_{12}}{A} \right) \left(\frac{\Delta_{13}}{B_\pm} \right) \cos \frac{\Delta_{13}L}{2} \sin \frac{AL}{2} \sin \frac{B_\pm L}{2} \quad \text{interference}$$

$$P_4 = \mp J \sin \delta \left(\frac{\Delta_{12}}{A} \right) \left(\frac{\Delta_{13}}{B_\pm} \right) \sin \frac{\Delta_{13}L}{2} \sin \frac{AL}{2} \sin \frac{B_\pm L}{2} \quad \text{interference}$$

where

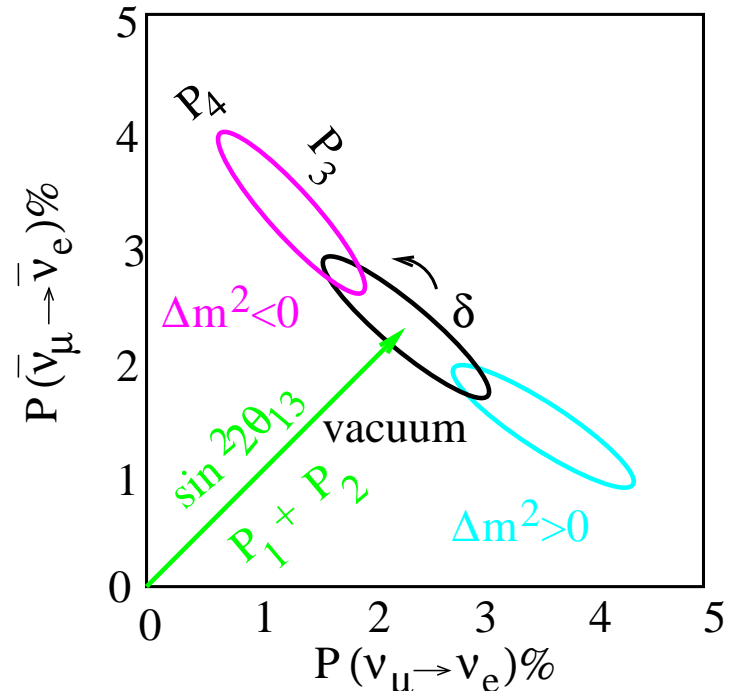
$$\Delta_{ij} = \frac{\Delta m_{ij}^2}{2E_\nu}$$

$$A = \sqrt{2} G_F n_e$$

$$B_\pm = |A \pm \Delta_{13}|$$

$$J = \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23}$$

and the \pm signifies neutrinos or antineutrinos



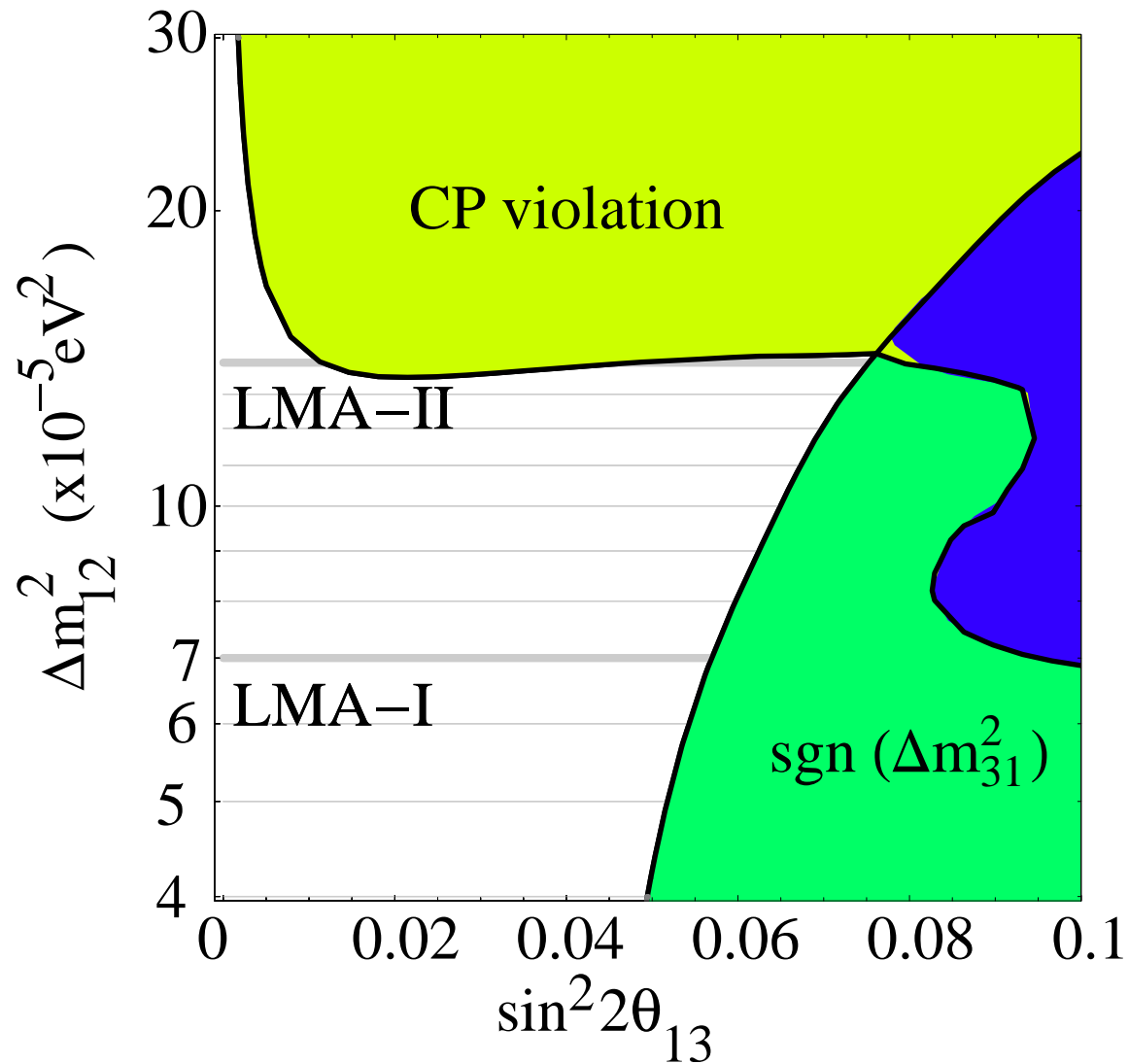
Scales of $\sin^2 2\theta_{13}$

	10^{-1}	3×10^{-2}	10^{-2}	10^{-3}	10^{-4}		
Finding $\theta_{13} \neq 0$	MINOS and CNGS	Conventional Phase I Superbeams					
Mass Hierarchy	Combinations of conventional Phase I SB's						
Evidence for \mathcal{CP}	Combinations of conventional Phase I SB's						
Beam Name	Peak Energy (GeV)	Detector	Mass (kton)	Power (MW)	$\sin^2 2\theta_{13}$ sens. ^a	δ^b	Matter Effect
OPERA	17	Pb-Emulsion	1.8	0.15	0.04	-	
ICARUS	17	LAr TPC	2.4	0.15	0.03	-	
MINOS	3.5	Steel-Scint.	5	0.4	0.05	-	
T2K	.7	H_2O Υ	22.5	0.8	0.006	-	-
NO ν A	2	Wood-Scint.	50	0.4	0.004	-	$\geq \mathcal{CP}$

Phase I Superbeam Combinations:

Minakata, Nunokawa, Parke, 2002; Huber, Lindner, Winter, 2002

What can you learn from T2K *and* NO ν A?



Winter, Huber, Lindner
Ref: Nucl. Phys. **F654**,
2003

With some regions
of parameter space
we might just see
hint of CP-violation

Scales of $\sin^2 2\theta_{13}$, continued

	10^{-1}	3×10^{-2}	10^{-2}	10^{-3}	10^{-4}
Finding $\theta_{13} \neq 0$	MINOS & CNGS	Conventional Superbeams Phase I	Conventional Superbeams Phase II		
Mass Hierarchy	Combinations of Phase I Superbeams	Combinations of Phase II Superbeams			
Evidence for \mathcal{CP}	Combinations of Phase I Superbeams	Combinations of Phase II Superbeams			

Beam Name	Peak Energy (GeV)	Detector	Mass (kton)	Power (MW)	$\sin^2 2\theta_{13}$ sens. ^a	δ^b	Matter Effect
T2HK	.7	H_2O Υ	450	4	0.001	$ \delta > 20^\circ$	$< \mathcal{CP}$
Super-NO ν A	2	Wood-Scint.	50+50	2	0.001	135 ± 20	$\geq \mathcal{CP}$
BNL2NUSL	1	H_2O Υ	500	1	0.004	45 ± 20	$> \& < \mathcal{CP}$
CERN SPL + β	.25	H_2O Υ	400	4	0.0016	90 ± 30	$\ll \mathcal{CP}$

Phase II combinations: Barger, Marfatia, Whisnant 2002

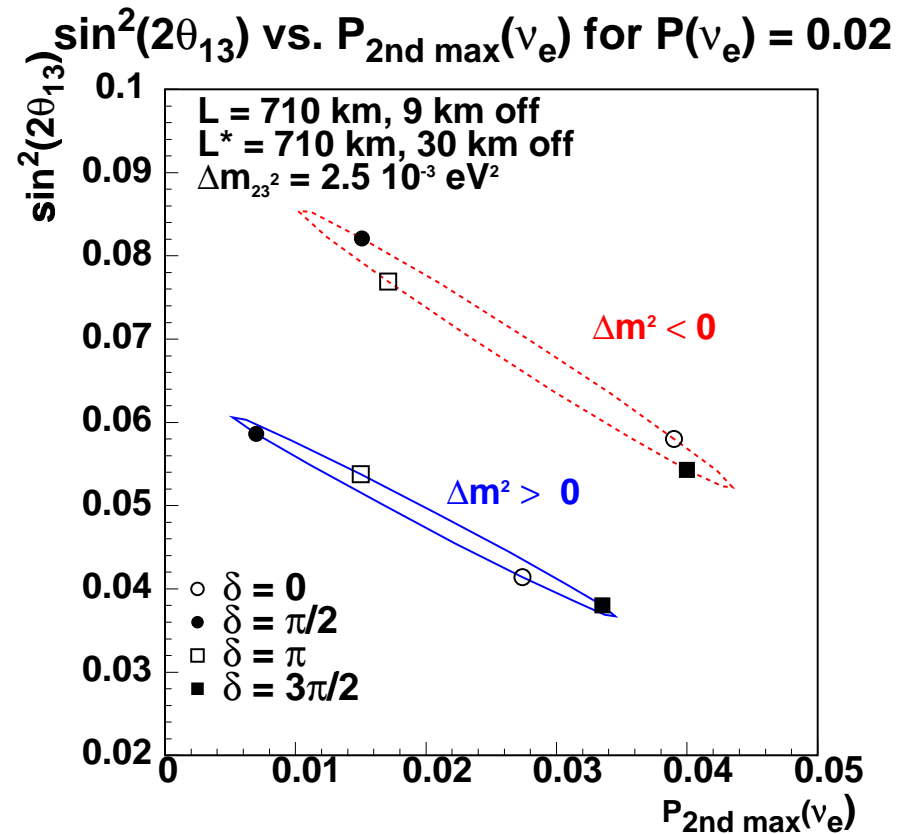
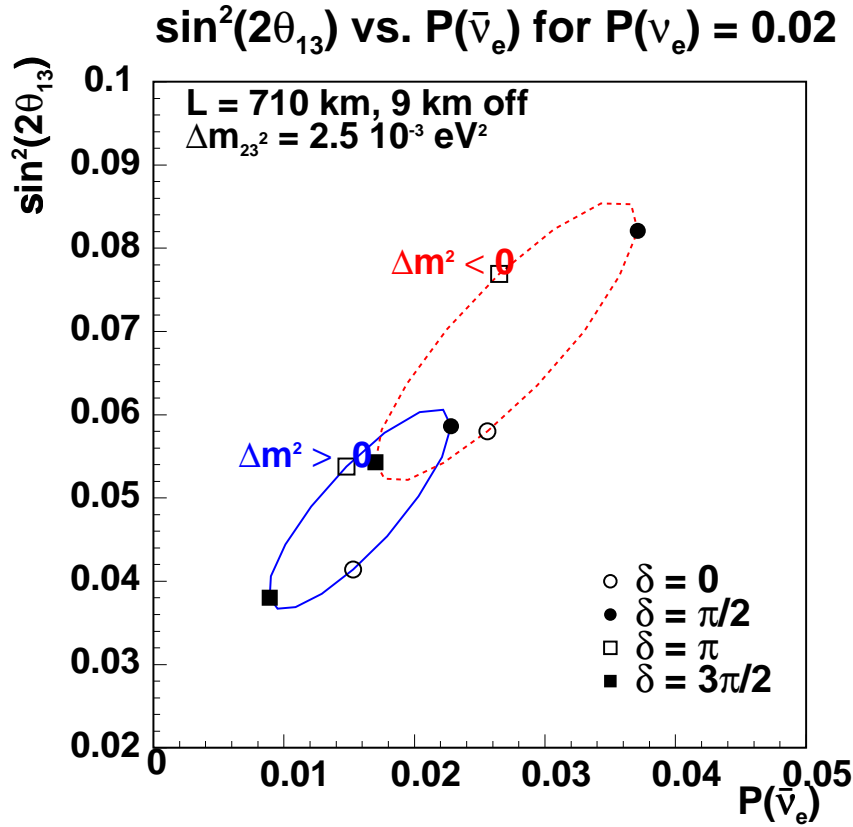
Running at Second Oscillation Maximum

Imagine NO ν A and/or T2K measures $P(\nu_\mu \rightarrow \nu_e) = 0.02\dots$

1st max: $CP \approx Matter$

2nd max: $CP \rightarrow 3 \times CP$

$Matter \rightarrow E_{2nd}/E_{1st}$

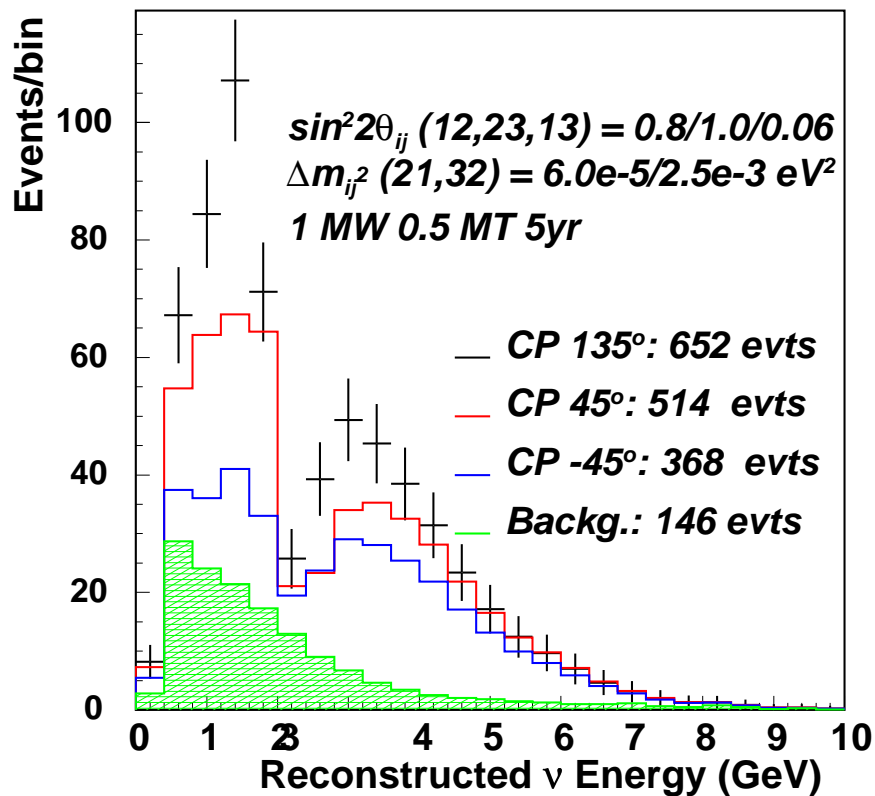


Maybe mass hierarchy can be determined, maybe not...

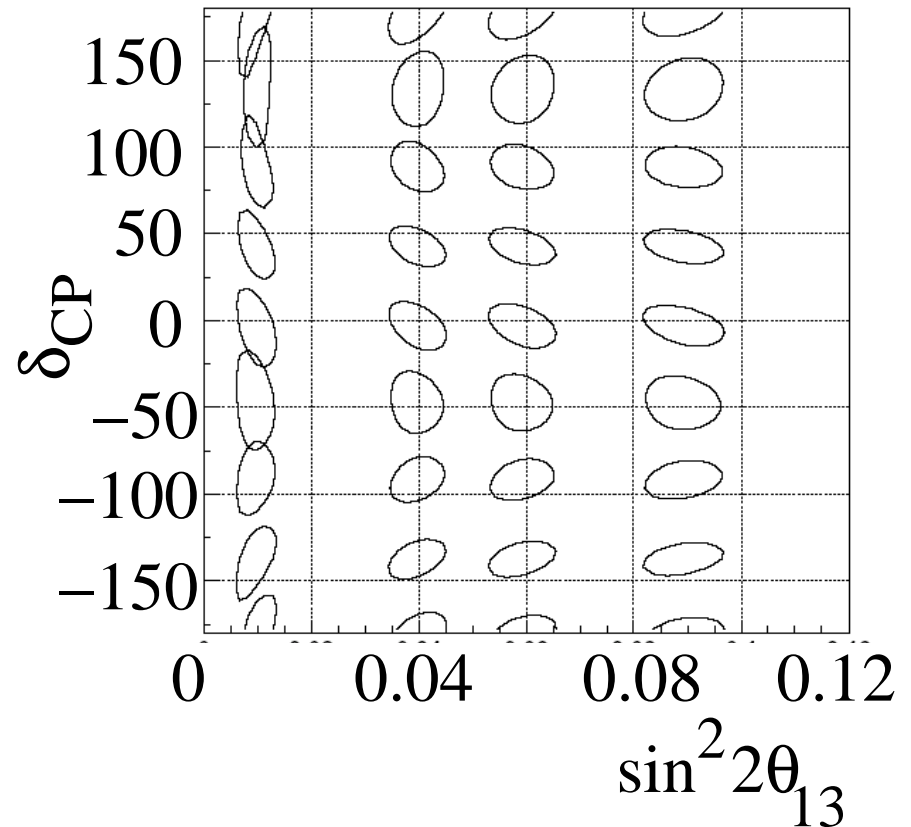
Now imagine it can't...

G.Feldman, HEPAP meeting

BNL LOI ν_e Appearance



Regular hierarchy $\nu\nu$ and $\text{Anti}\nu\nu$ running



again, ONLY 1-ring events in detector, 68% CL limits, different test points include π^0 backgrounds

CP violation effects huge at second oscillation maximum...

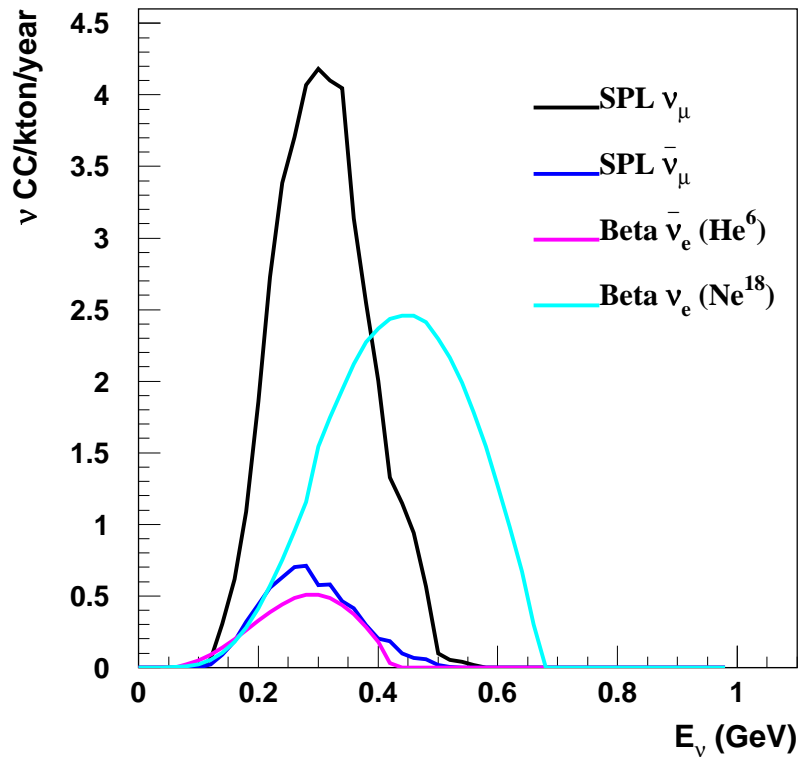
5 years ν running

5 years $\bar{\nu}$ running

similar for other hierarchy

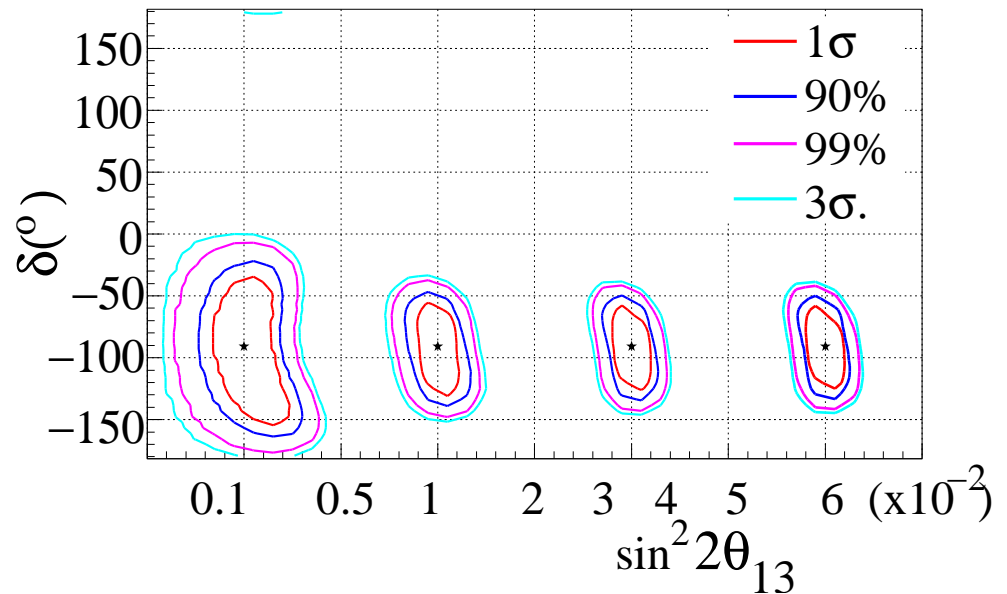
Diwan, 1/30/04 APS Study meeting

β and Conventional Beams at CERN



Conventional Ion production & acceleration, use same H_2O Υ detector in Frejus
 Run $\nu_e(Ne^{18})$ and $\bar{\nu}_e(He^6)$ in separate bunches, using timing to distinguish beams

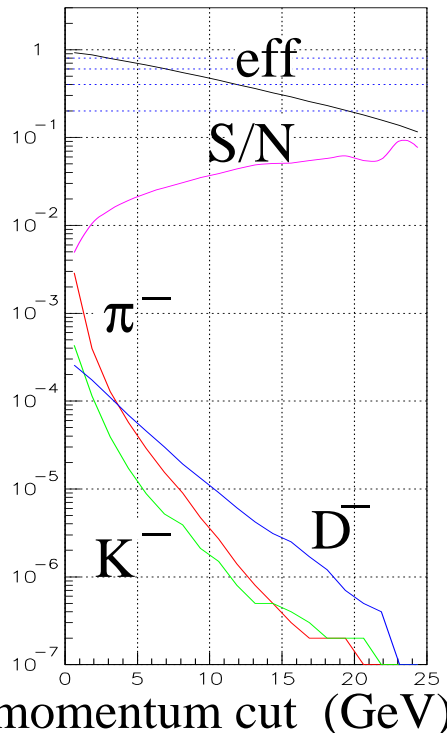
- Ability to see maximal CP violation:
- Cross-Check of Framework
 - $\nu_e \rightarrow \nu_\mu$ and $\bar{\nu}_e \rightarrow \bar{\nu}_\mu$
 - $\nu_\mu \rightarrow \nu_e$ and $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$
- added precision...
by including all measurements



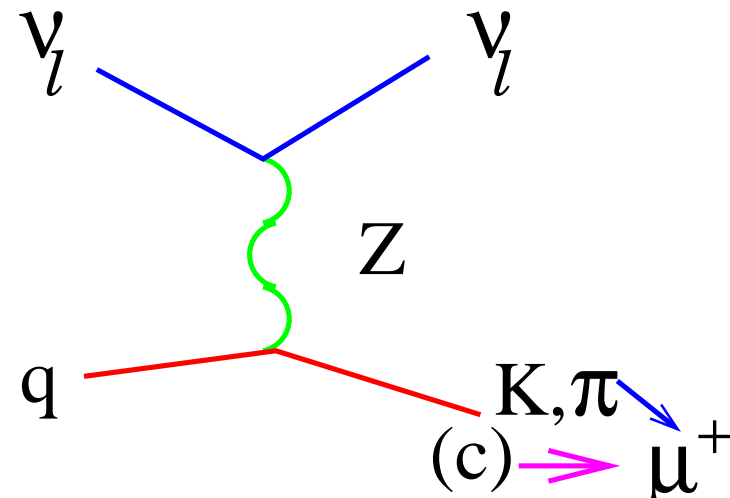
1° means $\sin^2 2\theta_{13} = 1.2 \times 10^{-3}$
 Ref: M. Mezzetto, NuFact03

Why is $\nu_e \rightarrow \nu_\mu$ at a ν Factory Easy?

- Neutrinos/MW proton power *cf* conventional beams $\propto (E_\mu/15)^3$
- No Intrinsic ν_μ in the beam, only $\bar{\nu}_\mu$'s
- Charge of Muon easier to measure than e/π^0 separation
- Detector Technology straightforward (see MINOS)
- Backgrounds at $\leq 10^{-4}$ level, not few $\times 10^{-3}$



Cervera *et al*, Nucl.Phys.**B579** 17,2000



Momentum cut on muon
easily removes
backgrounds

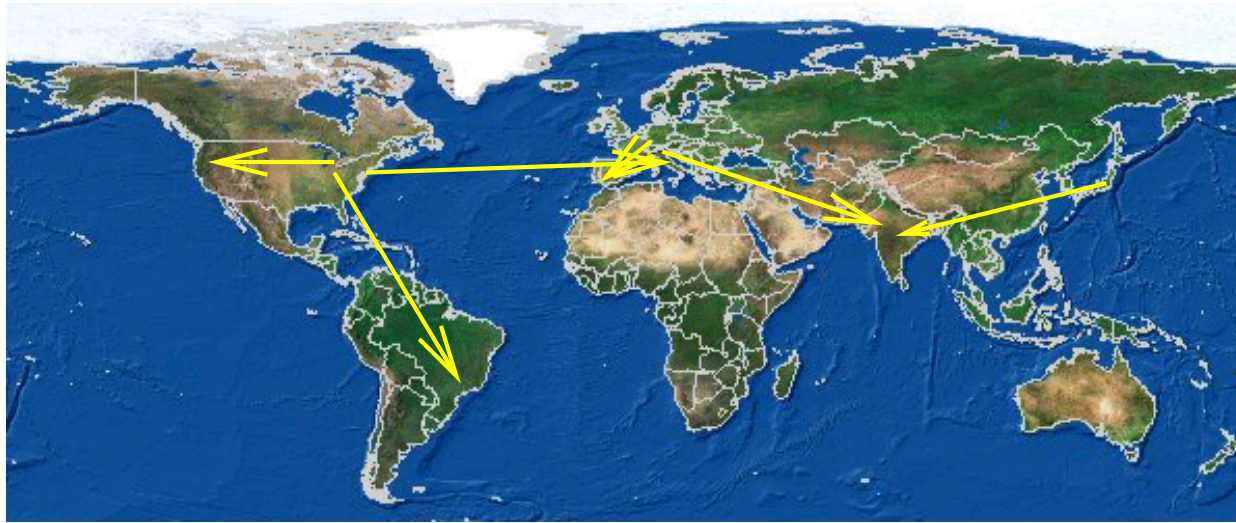
Scales of $\sin^2 2\theta_{13}$					
	10^{-1}	3×10^{-2}	10^{-2}	10^{-3}	10^{-4}
Finding $\theta_{13} \neq 0$	MINOS & CNGS	Conventional Superbeams Phase I	Conventional Superbeams Phase II	Combinations of ν Factories	
Mass Hierarchy	Combinations of Phase I Superbeams	Combinations of Phase II Superbeams	Combinations of ν Factory+SB	Combinations of ν Factories	
Evidence for $\epsilon\mathcal{P}$	Combinations of Phase I Superbeams	Combinations of Phase II Superbeams	Combinations of ν Factories		

Superbeam ν Fact combo: Burguet-Castell *et al*, 2002

ν Factory with τ ID: Donini, Meloni, Migliozi, 2002; Autiero *et al*, 2003

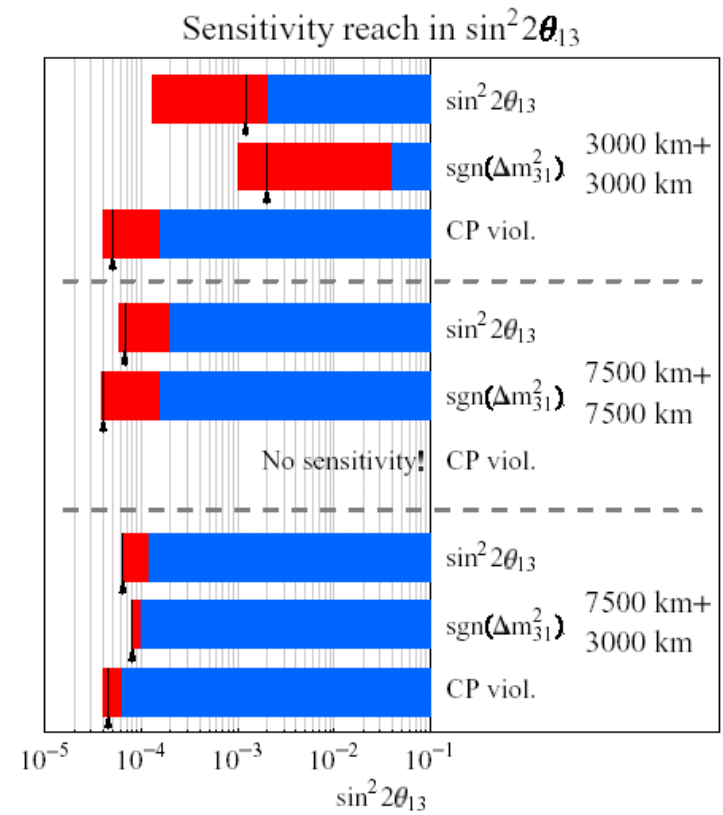
ν Factory “magic baseline”: Lipari, 2000; Burguet-Castell, 2001; Barger, Marfattia, Whisnant, 2002; Huber, Lindner, 2002; Huber and Winter, 2003, Asratyan *et al*, 2003

Ultimate ν Factory Reach



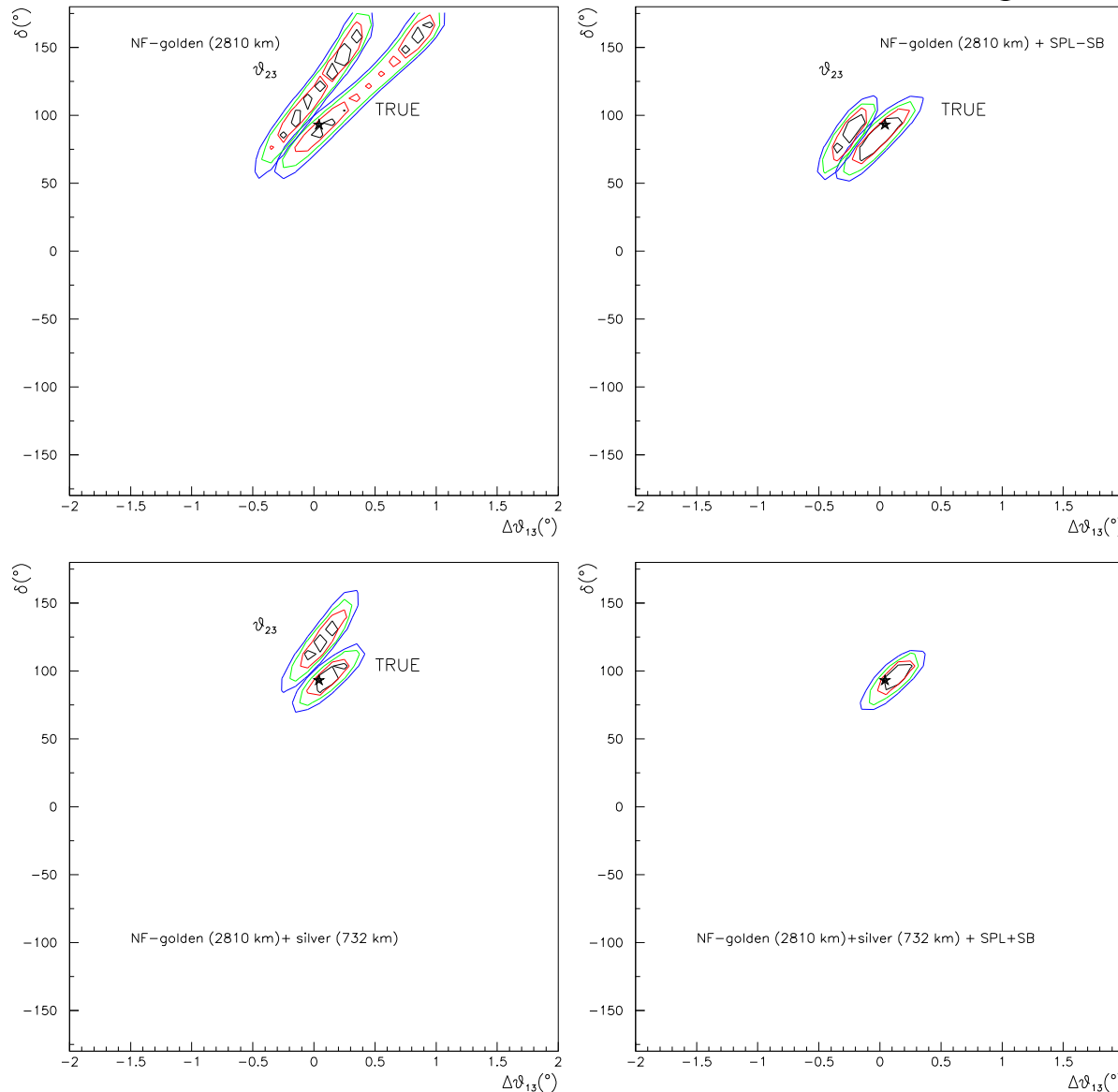
- Lots of work done in past
- apologies for not showing it
- New Studies: “magic baselines”
 $\sin \frac{\sqrt{2}G_F n_e L}{2} = 0$, for example $\frac{\sqrt{2}G_F n_e L}{2} = 2\pi$
 $L \sim 7250 km$
 only P_1 remains of P_1 through P_4

Asymmetry is close to 1 in this case:
 seeing $\nu_\mu \rightarrow \nu_e$ at this L gives
 θ_{13} and mass hierarchy!
 Huber, Winter Phys.Rev.**D68** 2003



Combining ν Factory and ...

From Donini, NuFact03: There were even ambiguities with the neutrino factory:



Getting to ultimate precision means combining data from several channels:

- Wrong-sign muons
- $\nu_e \rightarrow \nu_\tau$
- Conventional Beams

hep-ph/0310014

Summary of Sensitivities

Beam Name	Mass (kton)	Power (MW)	$\sin^2 2\theta_{13}$ sens. ^a	δ^b	Matter Effect
OPERA ^o	1.8	0.15	0.04	-	
ICARUS ^o	2.4	0.15	0.03	-	
MINOS ^m	5	0.4	0.05	-	
CNGS ^{**}	2.35	.15	$\sim 0.02^{**}$		\geq CP
T2K	22.5	0.8	0.006	-	-
NO ν A	50	0.4	0.004	-	\geq CP
T2HK	450	4	$\sim 0.001^s$	$ \delta > 20^\circ$	$<$ CP
Super-NO ν A	100	2	$\sim 0.001^s$	135 ± 20	\geq CP
BNL2NUSL	500	1	0.004	45 ± 20	$>$ & $<$ CP
CERN SPL	400	4	0.0016	90 ± 30	\ll CP
β Beam	400	.04		T viol.	\ll CP
ν Factory	50	4	$< 10^{-4}$	90 ± 20	huge!

^a at $\Delta m_{32}^2 = 3 \times 10^{-3} eV^2$, at 90%CL

^b all evaluated at different regions of parameter space!

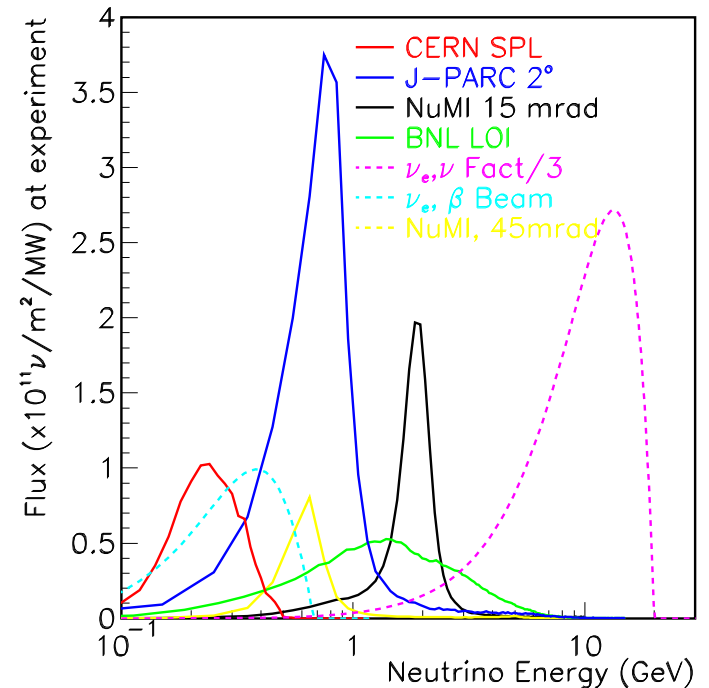
^o Komatsu, Migliozzi, Terranova J.Phys. **G29** 443, 2003 ^m

Diwan, Messier, Viren, L.Wai, NUMI-L-714

^s Assume 5% systematic uncertainty!

^{**}modified, Rubbia, Sala, hep-ph/0207084

Comparison of Fluxes
per MegaWatt
at each experiment:



Note ν Factory flux
divided by 3
to fit on graph!

Conclusions

Our understanding of the reach of conventional ν beam has evolved...

- Narrow band beams (“Off Axis”) mean lower backgrounds but...
- Going from Probabilities to Mixing angles not trivial
- Broad band beams (CERN,BNL) require MTon detectors for physics
 - Making PMT’s takes time
 - Even if water is cheap, MTon and MWatts are expensive units to work in
- if θ_{13} is right around the corner conventional beams may get us partly there

But the Physics Case for a ν Factory is alive and well:

- ν events/MW proton power still $\propto (E_\mu/10)^3$ cf conventional beams
- For precision measurements, want a precision machine!
 - Detector Technology for wrong sign muon appearance well-understood
 - Flux, cross section uncertainties still low
- Need to study many channels for real test of the framework...
- $\sin^2 2\theta_{13}$ and mass hierarchy reach still $100\times$ conventional beams...