

Physics Opportunities at Neutrino Factory:

Stephen Parke
Fermilab

NFMCC: Jan 25, 2009

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That's Fit to Print"

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Mass Found in Elusive Particle; Universe May Never Be the Same

Discovery on Neutrino Rattles Basic Theory About All Matter

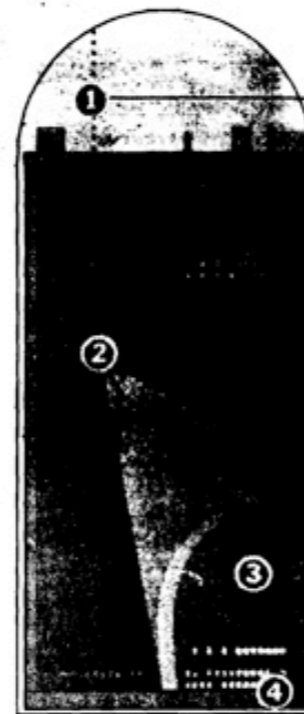
By MALCOLM W. BROWNE

TAKAYAMA, Japan, June 5 — In what colleagues hailed as a historic landmark, 120 physicists from 23 research institutions in Japan and the United States announced today that they had found the existence of mass in a notoriously elusive subatomic particle called the neutrino.

The neutrino, a particle that carries no electric charge, is so light that it was assumed for many years to have no mass at all. After today's announcement, cosmologists will have to confront the possibility that much of the mass of the universe is in the form of neutrinos. The discovery will also compel scientists to revise a highly successful theory of the composition of matter known as the Standard Model.

Word of the discovery had drawn some 300 physicists here to discuss neutrino research. Among other things, they said, the finding of neutrino mass might affect theories about the formation and evolution of galaxies and the ultimate fate of the

Detecting Neutrinos



Neutrinos pass through the Earth's surface to a tank filled with 12.5 million gallons of ultra-pure water . . .

. . . and collide with other particles . . .

. . . producing a cone-shaped flash of light.

The light is recorded by 11,200 20-inch light amplifiers that cover the inside of the tank.



LIGHT AMPLIFIER

And Detecting Their Mass

By analyzing the cones of light,

OKLAHOMA BLAST BRINGS LIFE TERM FOR TERRY NICHOLS

'ENEMY OF CONSTITUTION'

Judge Denounces Conspiracy and Hears From the Victims of a Terrifying Ordeal

By JO THOMAS

DENVER, June 4 — Calling him "an enemy of the Constitution," a Federal judge today sentenced Terry L. Nichols to life in prison without the possibility of parole for conspiring to bomb the Oklahoma City Federal Building, the deadliest terrorist attack ever on American soil.

In passing sentence after hearing from survivors of the blast and relatives of some of the 168 people who died in it, the judge, Richard P. Matsch of Federal District Court, said, "This was not a murder case."

He added: "It is a crime and the victims have spoken eloquently here. But it is not a crime as to them so much as it is a crime against the Constitution of the United States. That's the victim."



Bajram Curri, in no

Mixing Matrix:

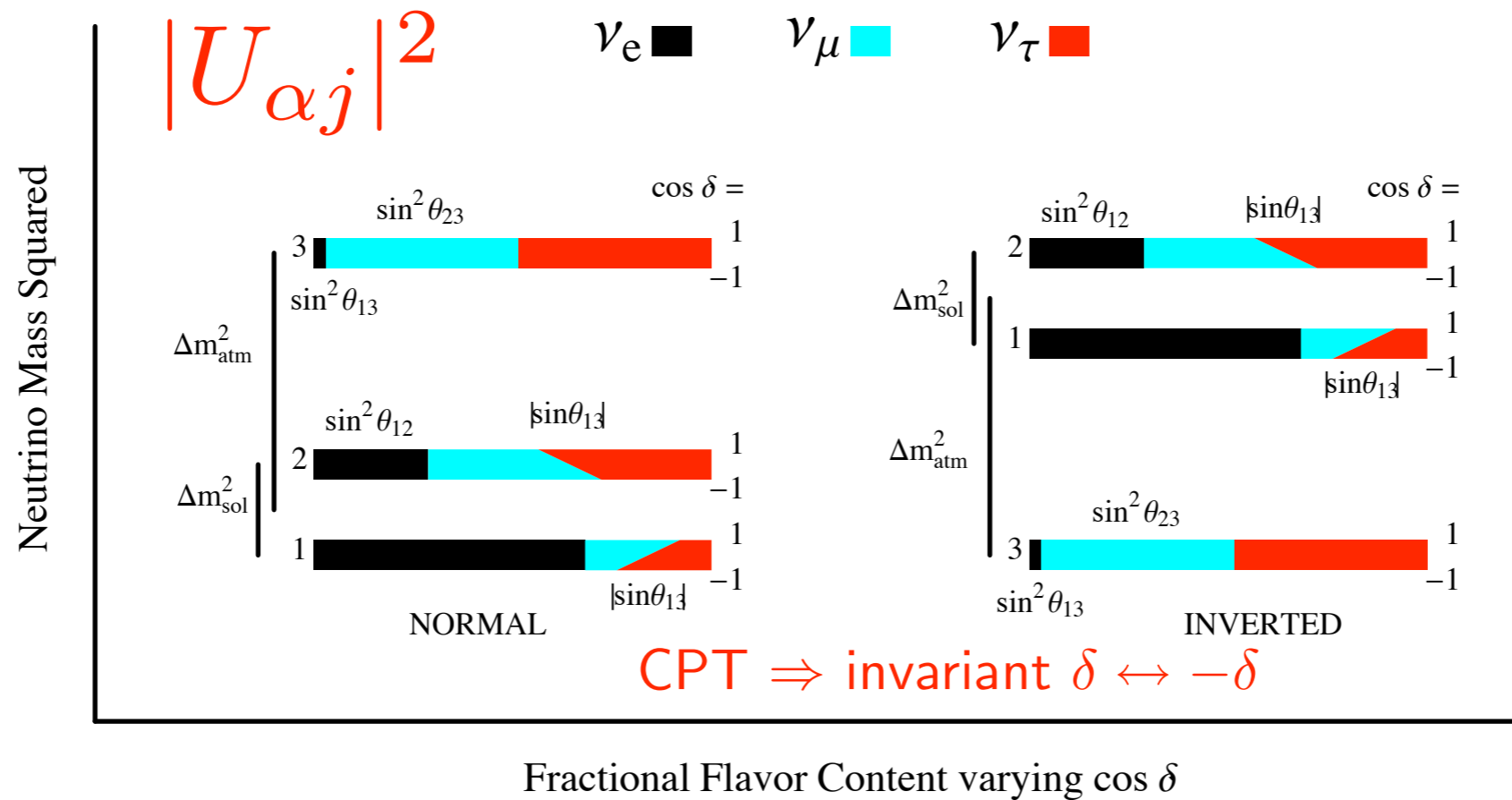
$$|\nu_e, \nu_\mu, \nu_\tau\rangle_{flavor}^T = U_{\alpha i} |\nu_1, \nu_2, \nu_3\rangle_{mass}^T$$

$$U_{\alpha i} = \begin{pmatrix} 1 & & & & & \\ & c_{23} & s_{23} & & & \\ & -s_{23} & c_{23} & & & \\ & & & c_{13} & s_{13}e^{-i\delta} & \\ & & & -s_{13}e^{i\delta} & c_{13} & \\ & & & & & 1 \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & \\ -s_{12} & c_{12} & \\ & & 1 \end{pmatrix} \begin{pmatrix} 1 & & \\ & e^{i\alpha} & \\ & & e^{i\beta} \end{pmatrix}$$

Atmos. L/E $\mu \rightarrow \tau$ Atmos. L/E $\mu \leftrightarrow e$ Solar L/E $e \rightarrow \mu, \tau$ $0\nu\beta\beta$ decay

500km/GeV

15km/MeV



$$\delta m_{sol}^2 = +7.6 \times 10^{-5} \text{ eV}^2$$

$$\sin^2 \theta_{12} \sim 1/3$$

$$|\delta m_{atm}^2| = 2.4 \times 10^{-3} \text{ eV}^2$$

$$\sin^2 \theta_{23} \sim 1/2$$

$$|\delta m_{sol}^2| / |\delta m_{atm}^2| \approx 0.03$$

$$\sin^2 \theta_{13} < 3\%$$

$$\sqrt{\delta m_{atm}^2} = 0.05 \text{ eV} < \sum m_{\nu_i} < 0.5 \text{ eV} = 10^{-6} * m_e$$

$$0 \leq \delta < 2\pi$$

One Global Fit:

Dominated by

parameter	best fit	2σ	3σ
Δm_{21}^2 [10^{-5}eV^2]	$7.65^{+0.23}_{-0.20}$	7.25–8.11	7.05–8.34
$ \Delta m_{31}^2 $ [10^{-3}eV^2]	$2.40^{+0.12}_{-0.11}$	2.18–2.64	2.07–2.75
$\sin^2 \theta_{12}$	$0.304^{+0.022}_{-0.016}$	0.27–0.35	0.25–0.37
$\sin^2 \theta_{23}$	$0.50^{+0.07}_{-0.06}$	0.39–0.63	0.36–0.67
$\sin^2 \theta_{13}$	$0.01^{+0.016}_{-0.011}$	≤ 0.040	≤ 0.056

KamLAND

MINOS

SNO

SuperK

Chooz

arXiv:0808.2016

Neutrino

Mass Spectrum:

- Quasi-Degenerate ?
- Hierarchical ?
- Normal or Inverted ?

Mixings:

- Deviations from $U_{Tri-Bi-Max}$
 $\sin^2 \theta_{13}, (\sin^2 \theta_{23} - 1/2), (\sin^2 \theta_{12} - 1/3)$
- Relationship between these deviations and $V_{CKM} - 1$
if any ?
- Magnitude and sign of CPV:
 $\propto \sin \theta_{13} \sin \delta$

eg: Sum rules

$$\theta_{13} \sim \theta_c / 3\sqrt{2} .$$

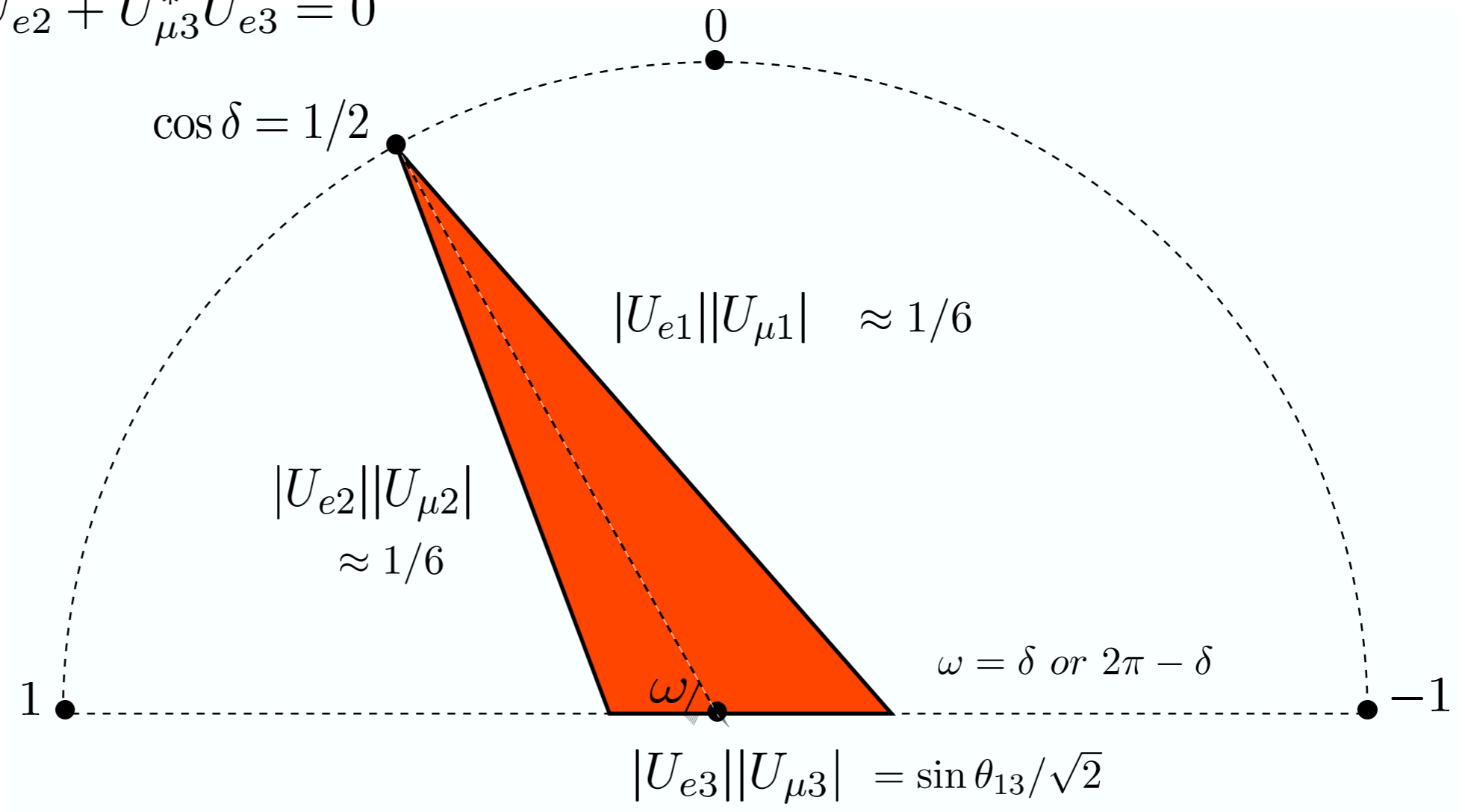
$$\tan^2 \theta_{\odot} \simeq \tan^2 \theta_{\odot, \text{TBM}} - \frac{1}{2} \theta_c \cos \delta ,$$

Sine/Signs

- $|U_{e3}|^2: \sin^2 \theta_{13}$
- Hierarchy: $sign(\delta m_{31}^2 \text{ or } \delta m_{32}^2)$
- CPV: $\sin \delta$
- Maximal Mixing: $\sin^2 \theta_{23} = \frac{1}{2}$
- Quadrant of δ : $\cos \delta = \pm \sqrt{1 - \sin^2 \delta}$
- Unitarity: lite sterile ν 's
- New Interactions and Surprises

Unitarity Triangle:

$$U_{\mu 1}^* U_{e 1} + U_{\mu 2}^* U_{e 2} + U_{\mu 3}^* U_{e 3} = 0$$



$$|J| = 2 \times \text{Area}$$

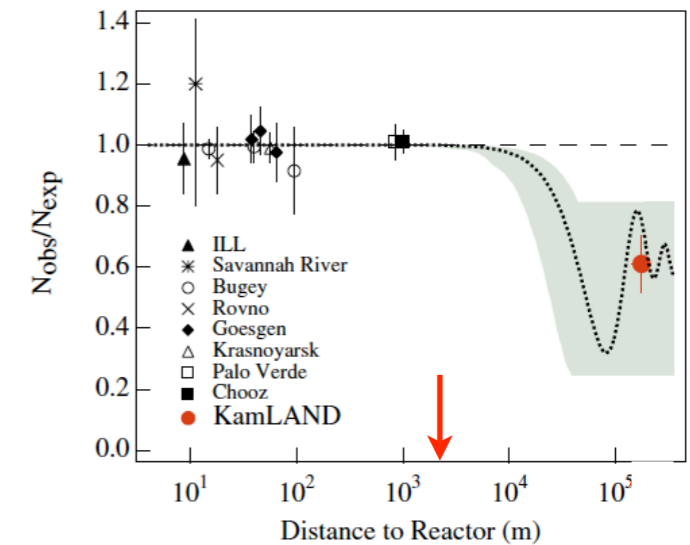
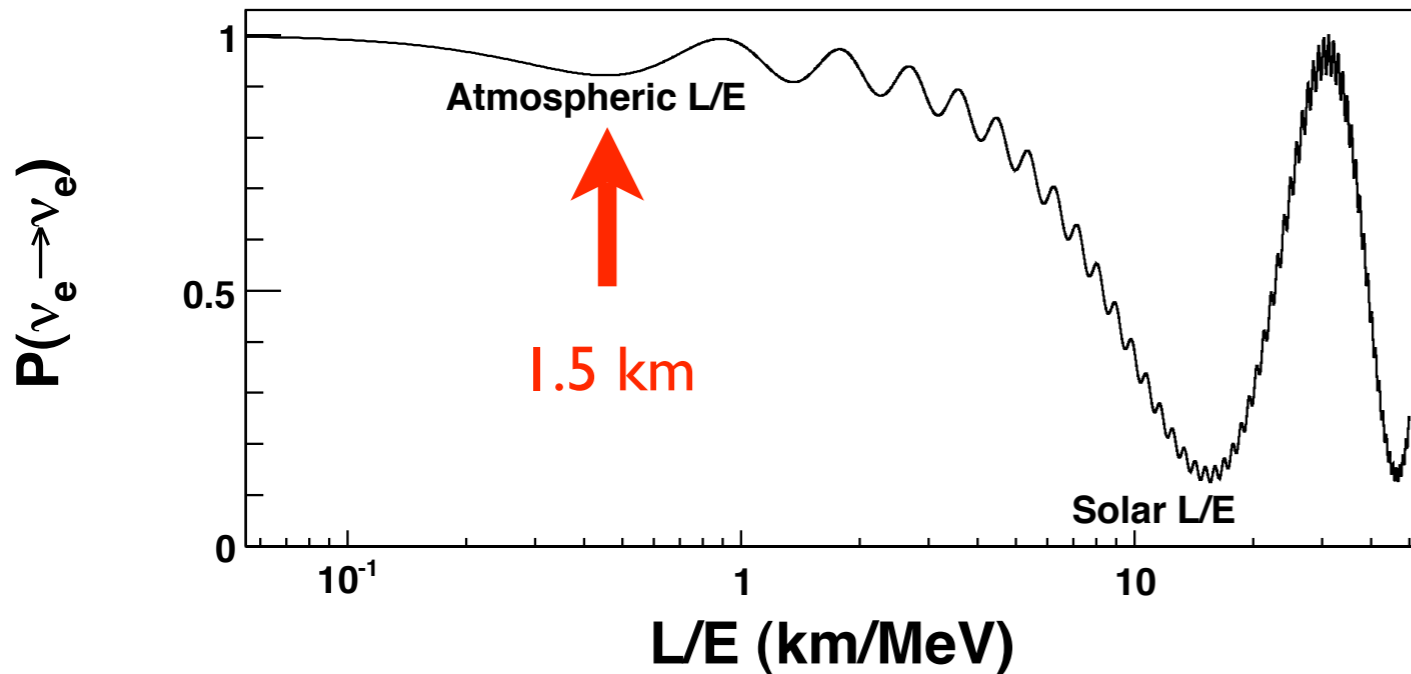
$$J = s_{12} c_{12} s_{23} c_{23} s_{13} c_{13}^2 \sin \delta$$

$\sin^2 \theta_{13}$ from Reactor Neutrinos:

kinematic phase:

$$\Delta_{ij} \equiv \frac{\delta m_{ij}^2 L}{4E}$$

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \Delta_{21} - \sin^2 2\theta_{13} (\cos^2 \theta_{12} \sin^2 \Delta_{31} + \sin^2 \theta_{12} \sin^2 \Delta_{32})$$



$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \approx 1 - \sin^2 2\theta_{13} \sin^2 \left(\frac{\delta m_{ee}^2 L}{4E} \right) - \mathcal{O}(\Delta_{21})^2$$

> 0.01

$$\delta m_{ee}^2 = \cos^2 \theta_{12} |\delta m_{31}^2| + \sin^2 \theta_{12} |\delta m_{32}^2|$$

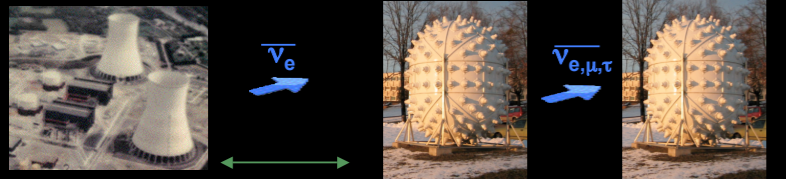
< 0.002

Double
Chooz:



90% C.L.
Sensitivities:

One nuclear plant & two detectors



Nuclear reactor
1,2 core(s) → ON/OFF : ok
4 cores → ON/OFF : no !

Near detector
5-130 tons
> 50 mwe

Far detector
5-130 tons
> 300 mwe

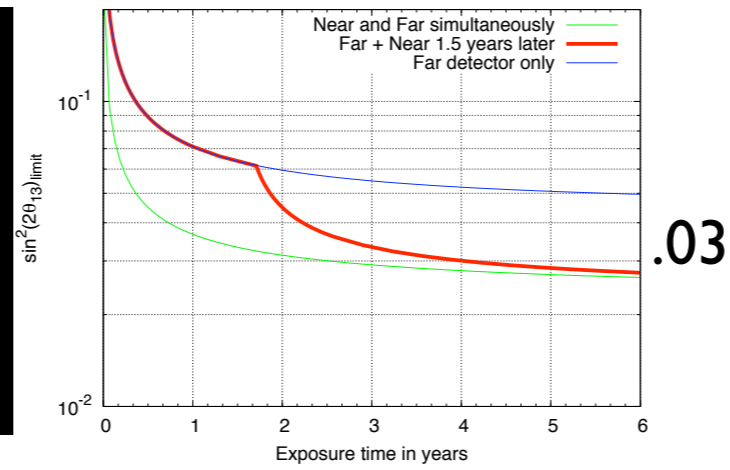
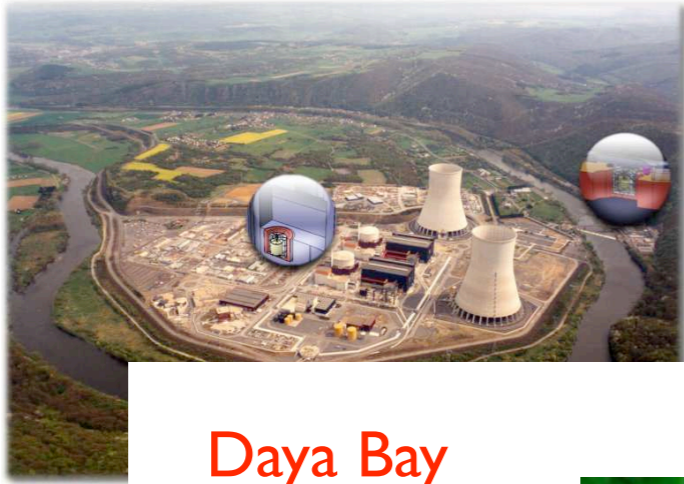


Figure 18: $\sin^2(2\theta_{13})$ sensitivity limit for the detectors installation scheduled scenario

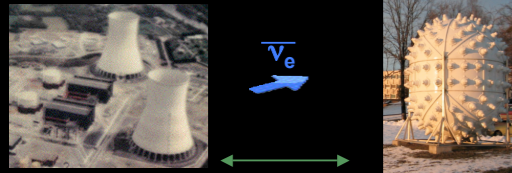
Double
Chooz:



Daya Bay

90% C.L.
Sensitivities:

One nuclear plant & two



Nuclear reactor

Near detector

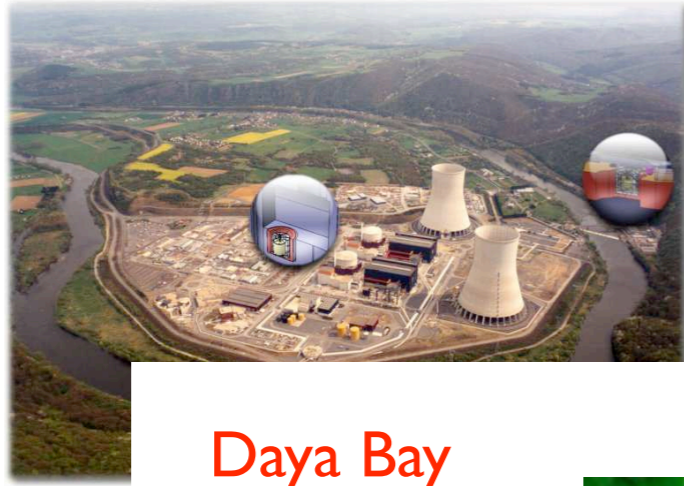
1,2 core(s) → ON/OFF : ok
≥ 4 cores → ON/OFF : no !

5-130 tons
> 50 mwe



push the limit on
 $\sin^2 2\theta_{13} < 0.01$

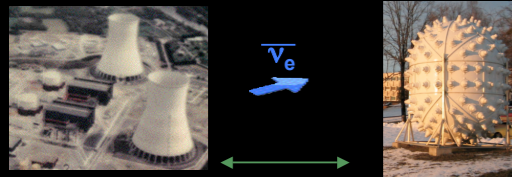
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Nuclear reactor

Near detector

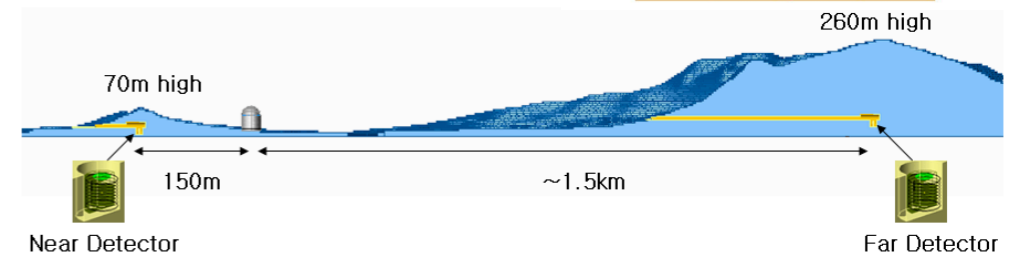
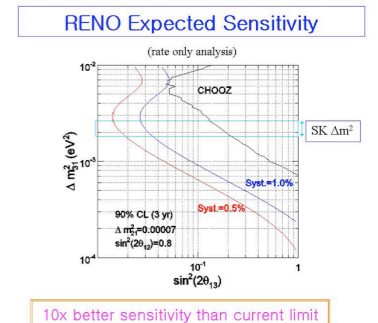
1,2 core(s) → ON/OFF : ok
 4 cores → ON/OFF : no !

5-130 tons
 > 50 mwe



(Reactor Experiment for Neutrino Oscillation)

push the limit
 $\sin^2 2\theta_{13} < 0$

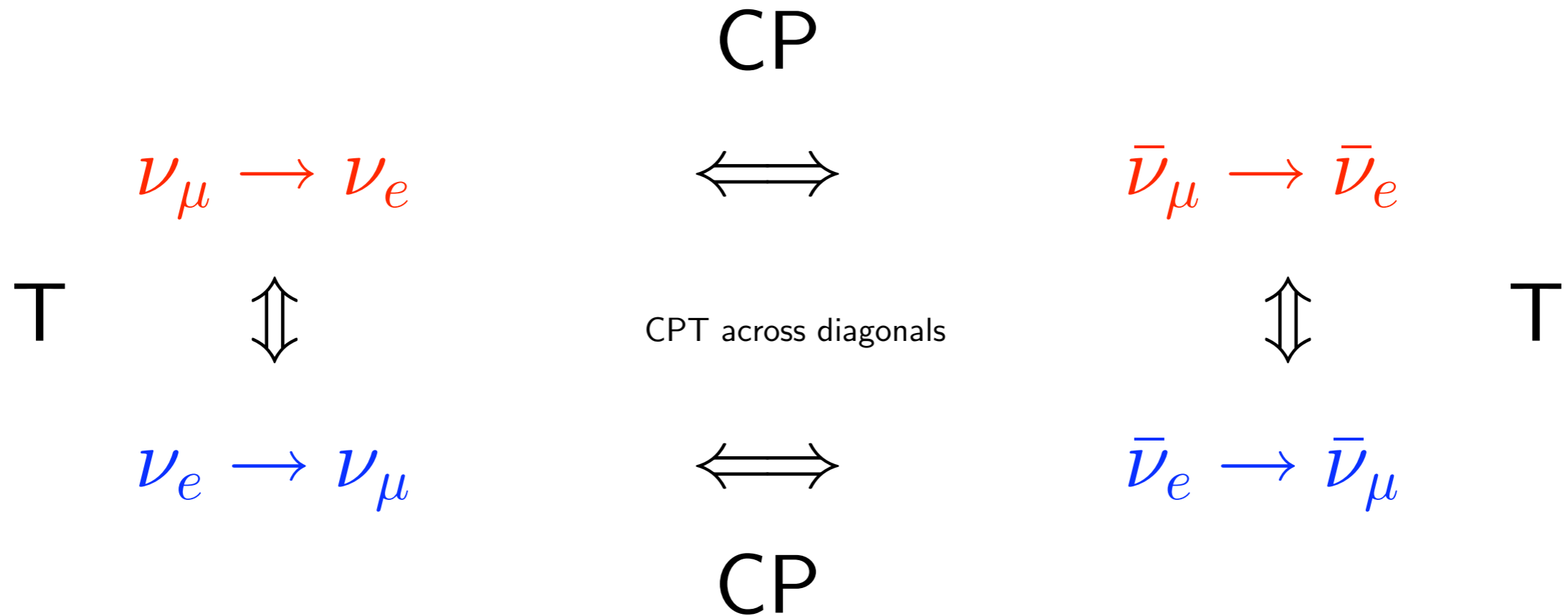


Sutherland Falls 580m #5



$$\nu_{\mu} \longrightarrow \nu_e$$

and related processes:



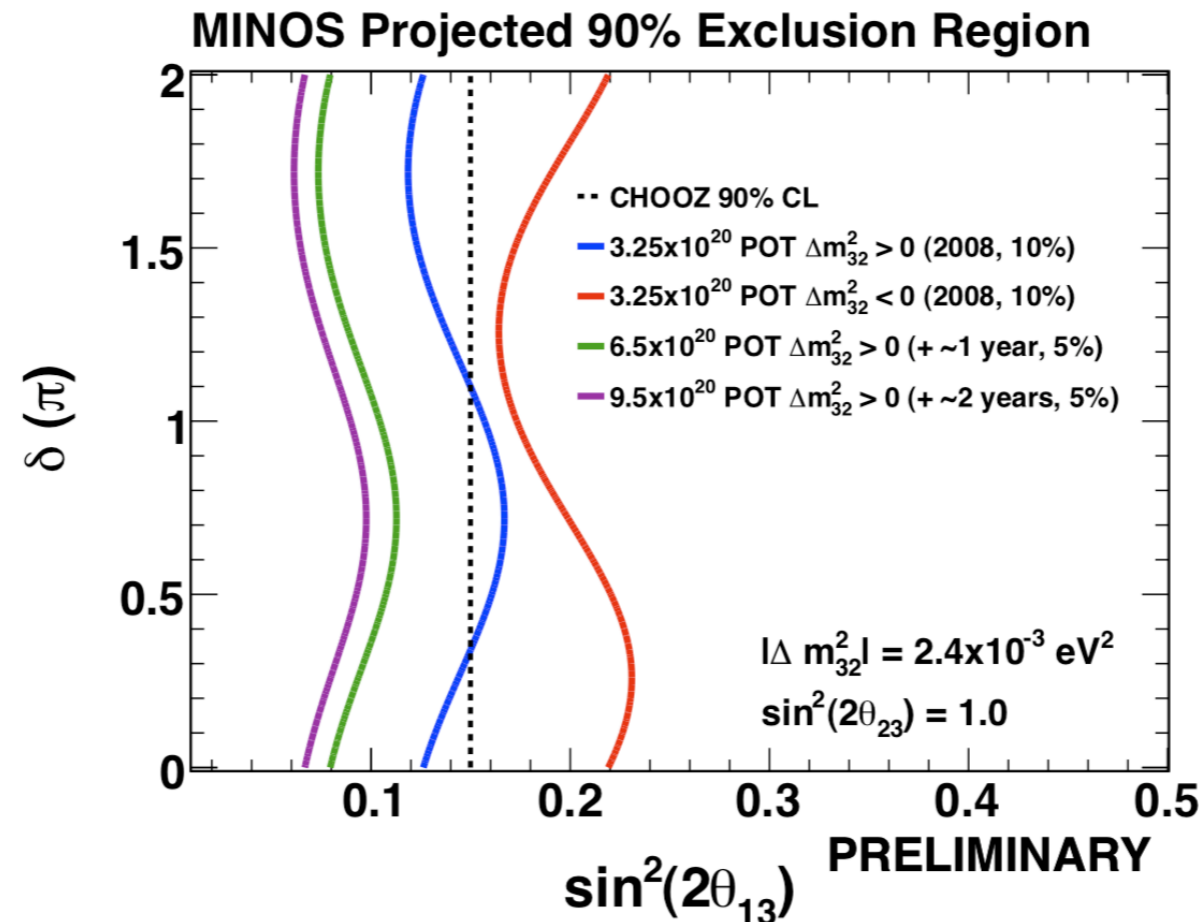
- First Row: Superbeams where ν_e contamination $\sim 1\%$
- Second Row: ν -Factory or β -Beams, no beam contamination

Even in matter, a vestige of CPT exists:
 Instead of **switch matter to anti-matter**, **switch neutrino hierarchy !!!**

MINOS:

ν_e Sensitivity

H. Gallagher
Tufts University
Neutrino 2008
May 27, 2008



At CHOOZ limit expect 12 ν_e signal events and 42 background events with 3.25×10^{20} protons.

Use sidebands to study predicted far detector backgrounds.

$$\nu_{\mu} \rightarrow \nu_e$$

$$\left| U_{\mu 3}^* e^{-im_3^2 L/2E} U_{e3} + U_{\mu 2}^* e^{-im_2^2 L/2E} U_{e2} + U_{\mu 1}^* e^{-im_1^2 L/2E} U_{e1} \right|^2$$

$$\nu_\mu \longrightarrow \nu_e$$

$$\left| U_{\mu 3}^* e^{-im_3^2 L/2E} U_{e3} + U_{\mu 2}^* e^{-im_2^2 L/2E} U_{e2} + U_{\mu 1}^* e^{-im_1^2 L/2E} U_{e1} \right|^2$$

use unitarity to eliminate $U_{\mu 1}^* U_{e1}$ term:

$$P(\nu_\mu \longrightarrow \nu_e) = \left| 2U_{\mu 3}^* U_{e3} \sin \Delta_{31} e^{-i\Delta_{32}} + 2U_{\mu 2}^* U_{e2} \sin \Delta_{21} \right|^2$$

$$\nu_\mu \longrightarrow \nu_e$$

$$\left| U_{\mu 3}^* e^{-im_3^2 L/2E} U_{e3} + U_{\mu 2}^* e^{-im_2^2 L/2E} U_{e2} + U_{\mu 1}^* e^{-im_1^2 L/2E} U_{e1} \right|^2$$

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Atmospheric δm^2

$$\sqrt{P_{atm}}$$

Solar δm^2

$$\sqrt{P_{sol}}$$

Vacuum LBL:

$$\nu_{\mu} \rightarrow \nu_e$$

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

$$\Delta_{ij} = \delta m_{ij}^2 L / 4E$$

CP violation !!!

where $\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \sin \Delta_{31}$

and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \sin \Delta_{21}$

Vacuum LBL:

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and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \sin \Delta_{21}$

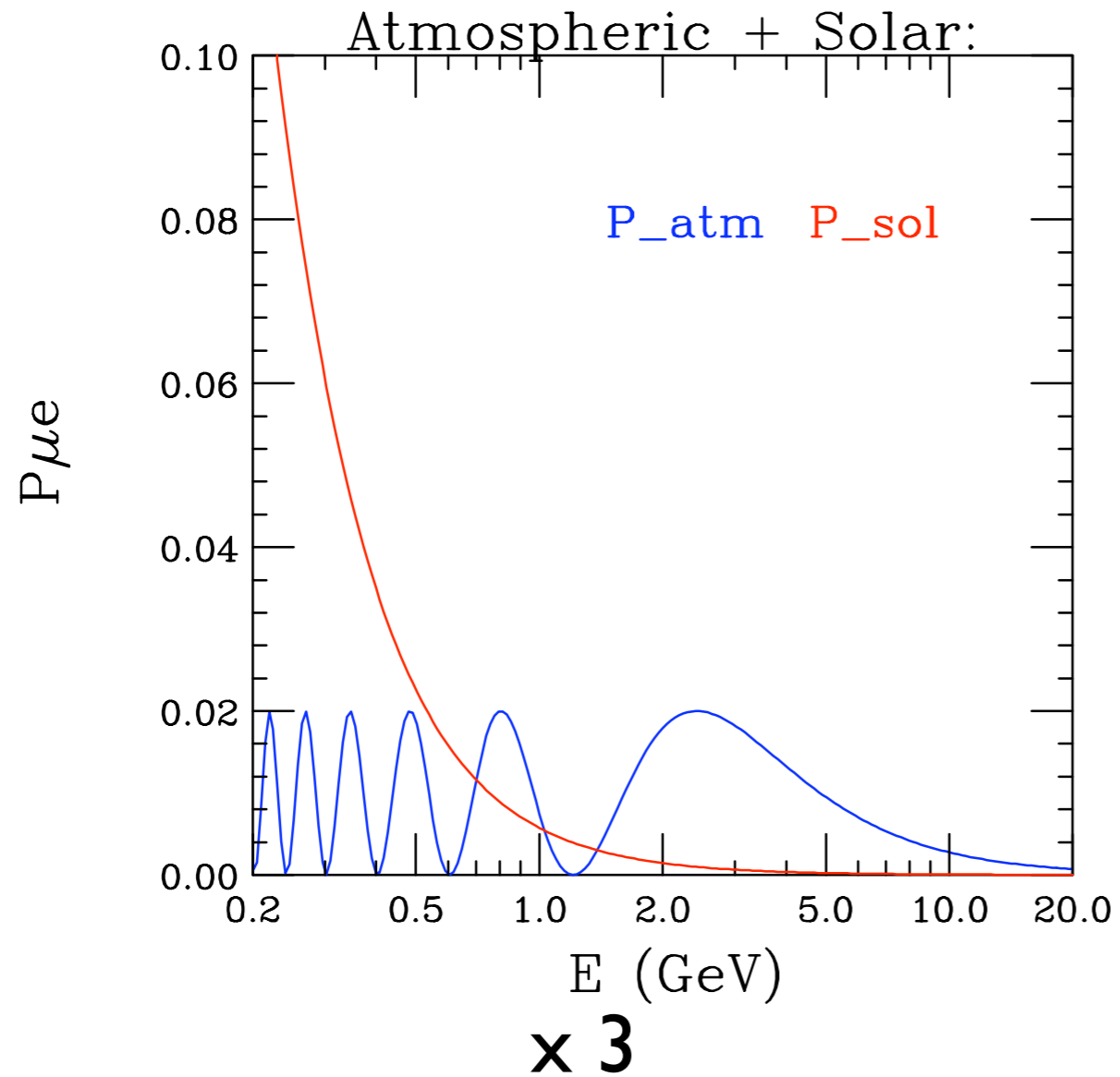
$$P_{\mu \rightarrow e} \approx P_{atm} + 2\sqrt{P_{atm}P_{sol}} \cos(\Delta_{32} \pm \delta) + P_{sol}$$

only CPV

$$\cos(\Delta_{32} \pm \delta) = \cos \Delta_{32} \cos \delta \mp \sin \Delta_{32} \sin \delta$$

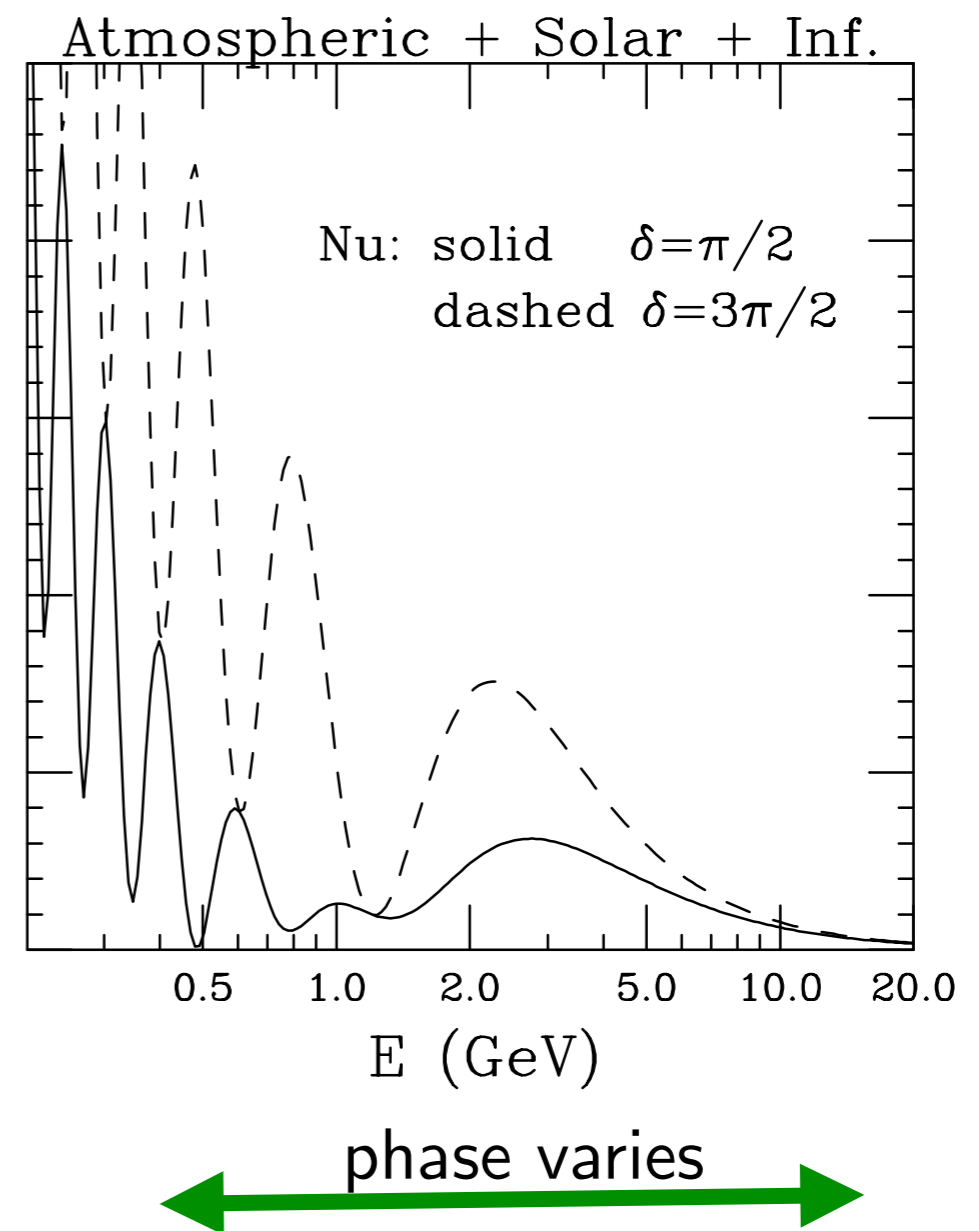
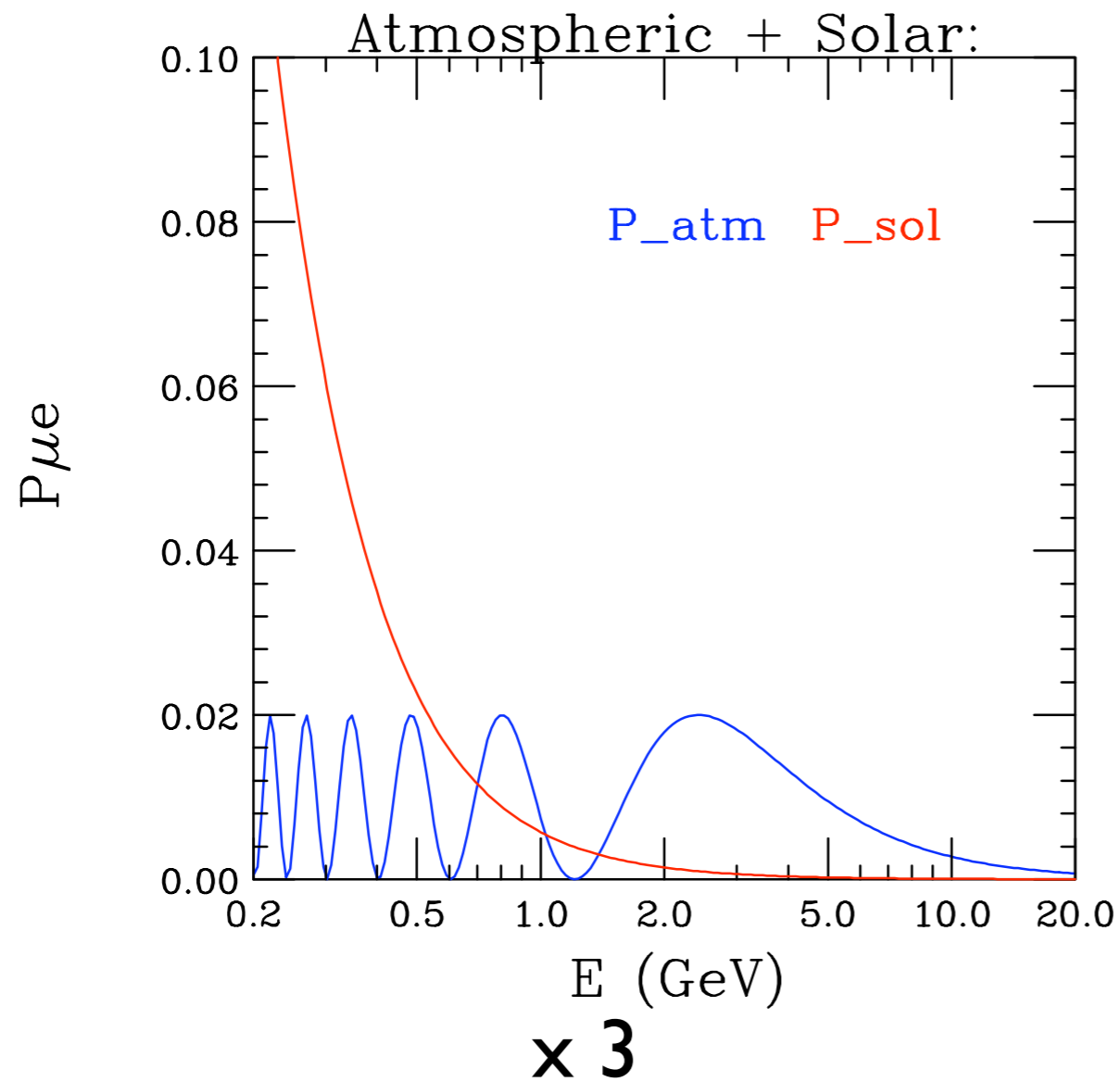
$$P(\nu_\mu \rightarrow \nu_e) \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} + \delta)} + \sqrt{P_{sol}} \right|^2$$

For $L = 1200 \text{ km}$
and $\sin^2 2\theta_{13} = 0.04$



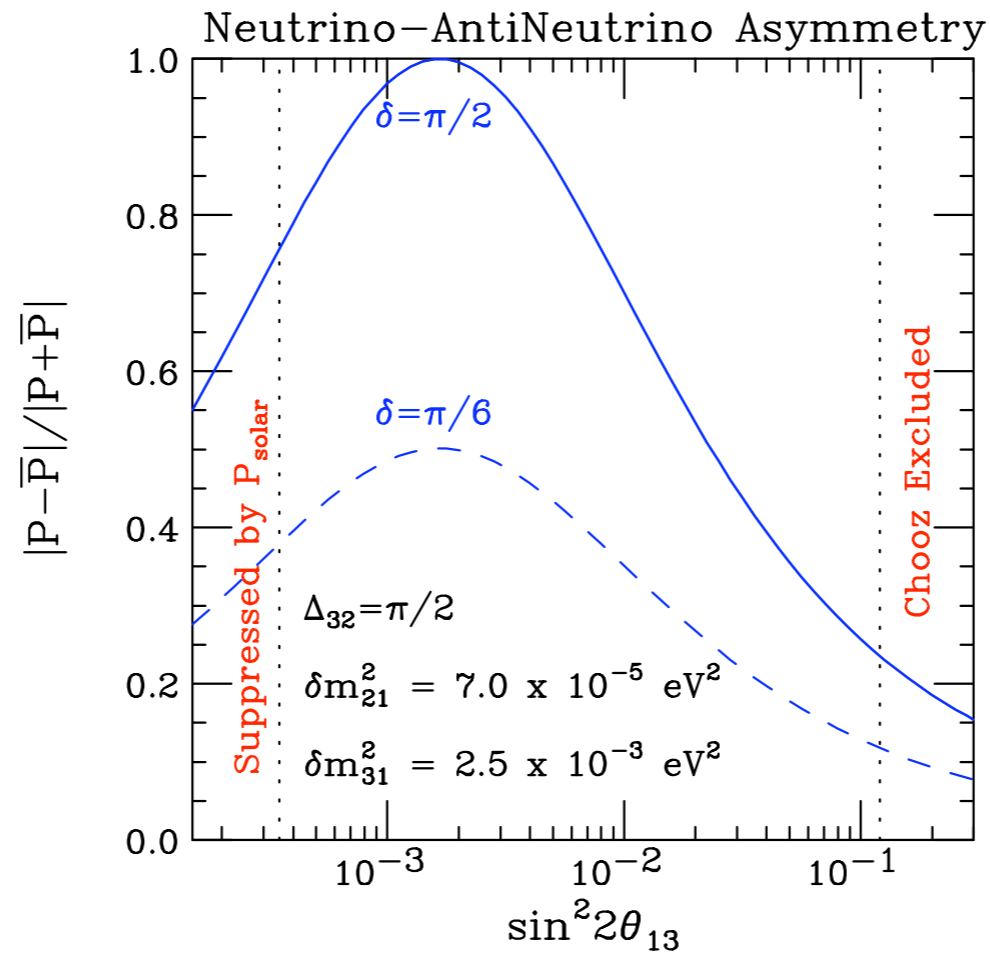
$$P(\nu_\mu \rightarrow \nu_e) \approx |\sqrt{P_{atm}} e^{-i(\Delta_{32} + \delta)} + \sqrt{P_{sol}}|^2$$

For $L = 1200 \text{ km}$
and $\sin^2 2\theta_{13} = 0.04$



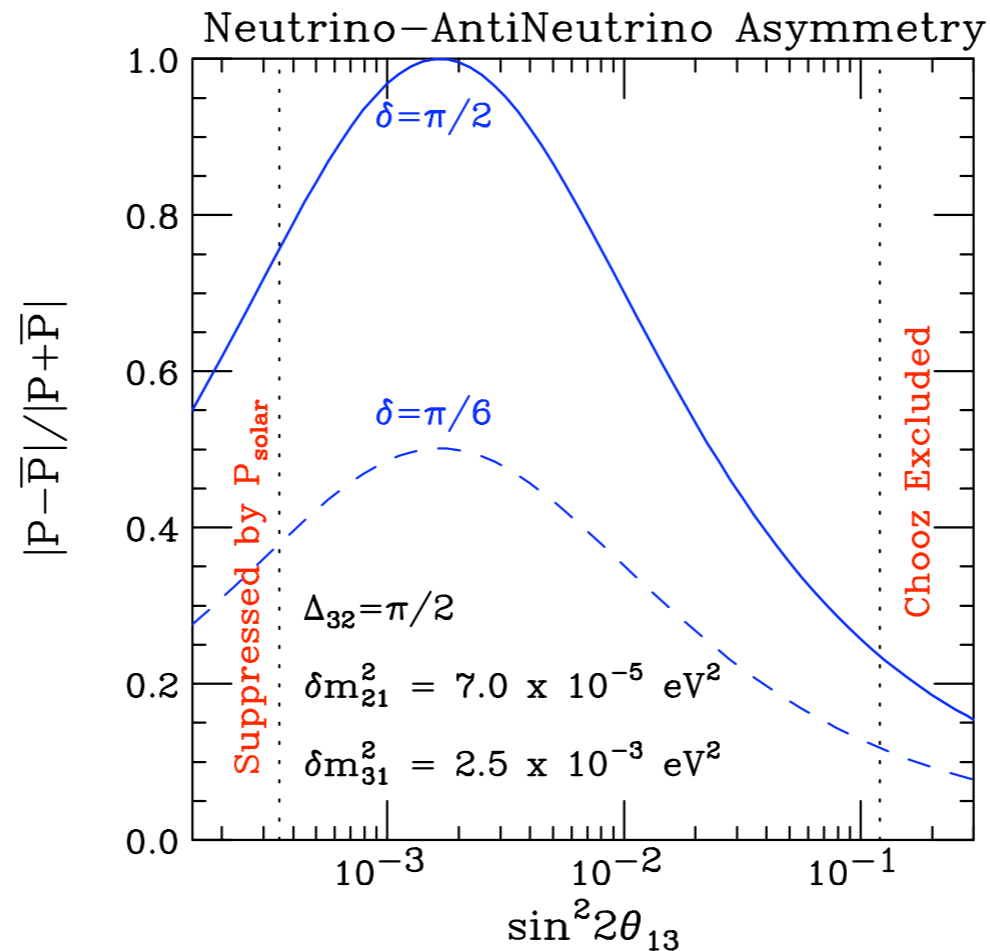
$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

Asymmetry Peaks:



$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

Asymmetry Peaks:



$$P_{atm} \leq P_{sol} \quad \text{when} \quad \sin^2 2\theta_{13} \leq \frac{\sin^2 2\theta_{12}}{\tan^2 \theta_{23}} \left(\frac{\delta m_{21}^2}{\delta m_{31}^2} \right)^2 \approx 0.001$$

In Matter:

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

where $\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31}$

and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{(aL)} \Delta_{21}$

$$a = G_F N_e / \sqrt{2} = (4000 \text{ km})^{-1},$$

In Matter:

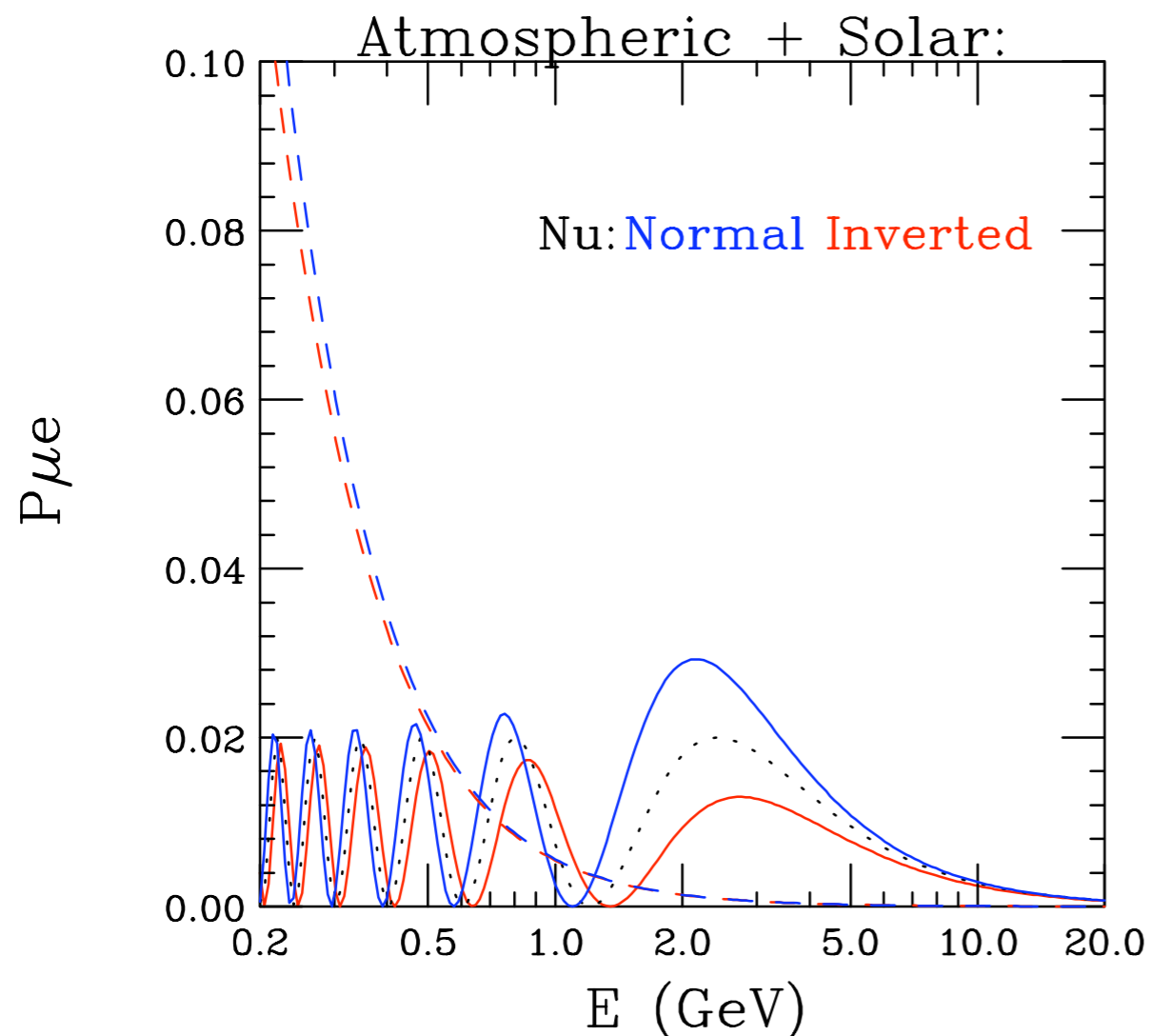
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In Matter:

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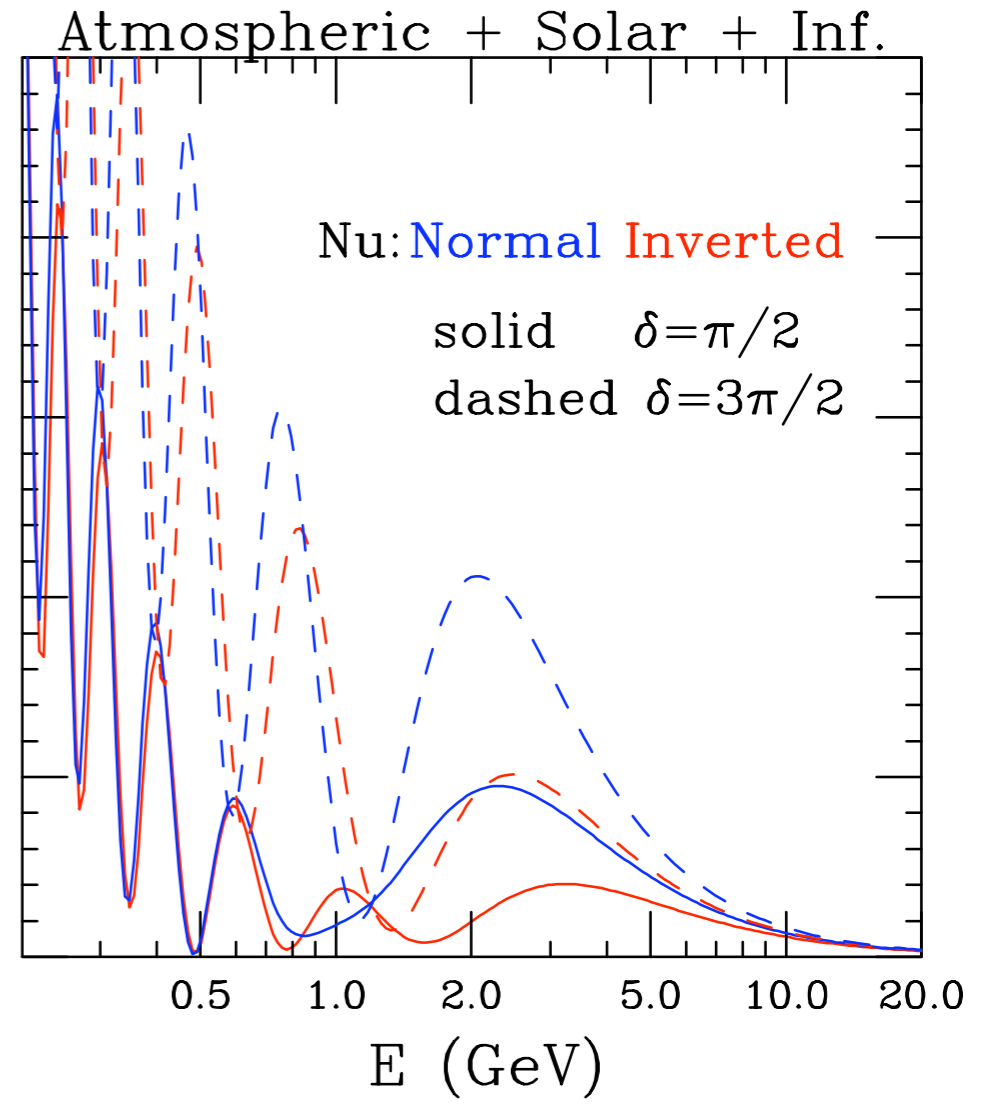
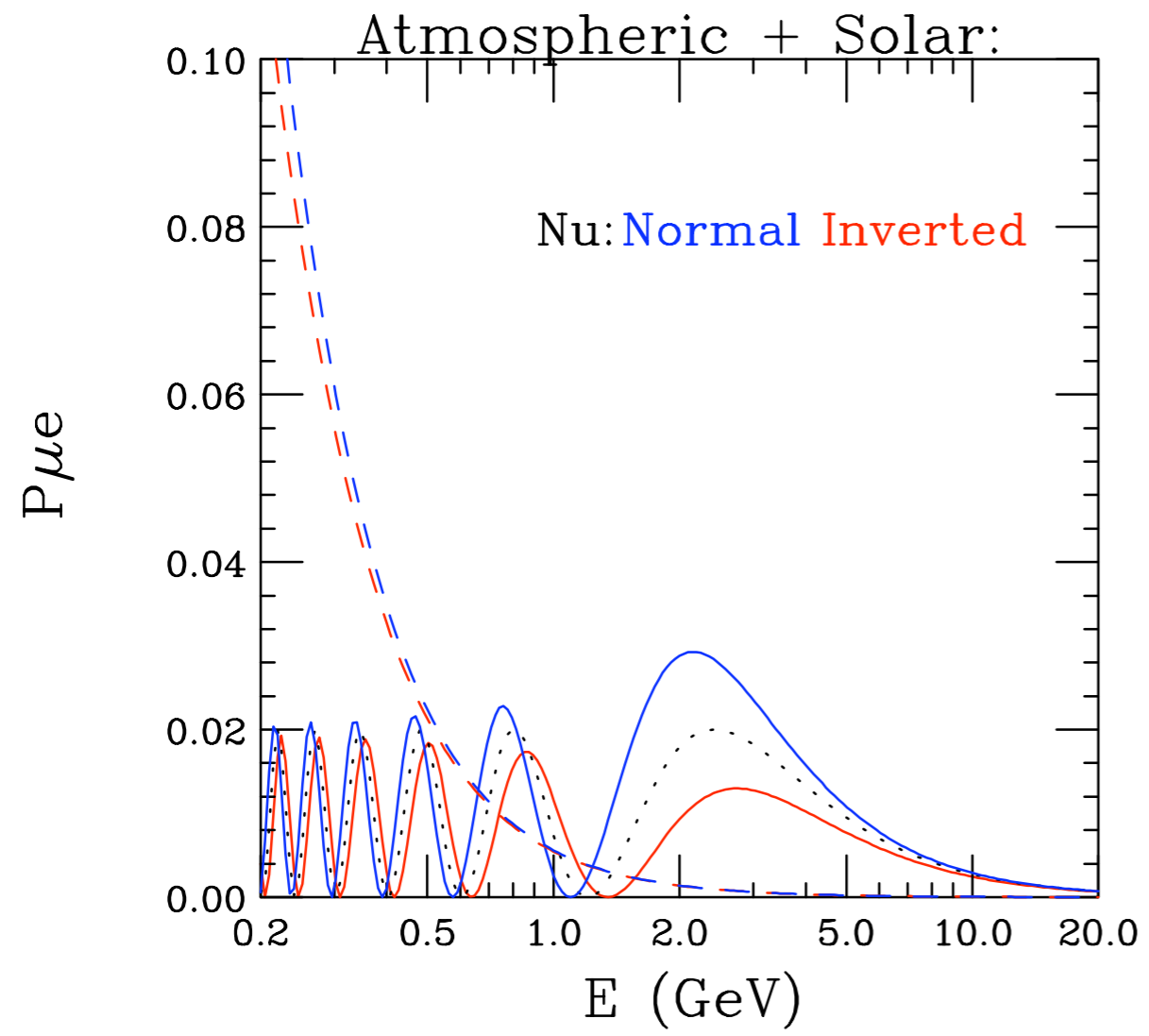
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For $L = 1200 \text{ km}$
and $\sin^2 2\theta_{13} = 0.04$

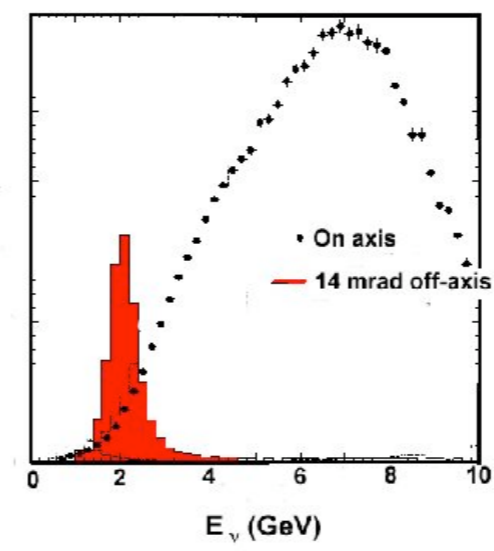
$$a = G_F N_e / \sqrt{2} = (4000 \text{ km})^{-1},$$

Anti-Nu: Normal Inverted
dashes $\delta = \pi/2$
solid $\delta = 3\pi/2$



Off-Axis Beams

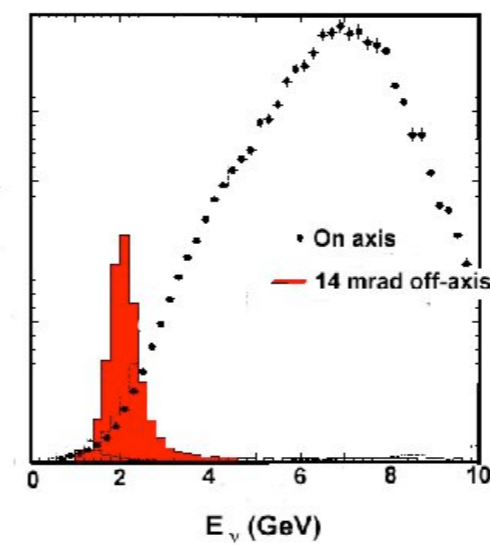
BNL 1994



π^0 suppression

Off-Axis Beams

BNL 1994



π^0 suppression

T2K

JHF \rightarrow Super-Kamiokande

- 295 km baseline
- Super-Kamiokande:
 - 22.5 kton fiducial
 - Excellent e/μ ID
 - Additional π^0/e ID
- Hyper-Kamiokande
 - 20 \times fiducial mass of SuperK
- Matter effects small
- Study using fully simulated and reconstructed data



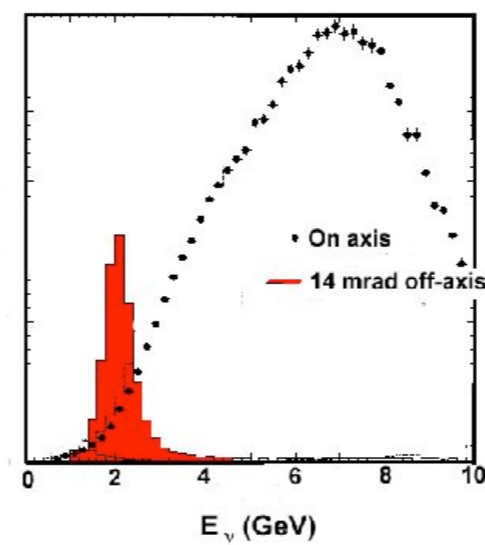
$L=295$ km and
Energy at Vac. Osc. Max. (vom)

$$E_{vom} = 0.6 \text{ GeV} \left\{ \frac{\delta m_{32}^2}{2.5 \times 10^{-3} \text{ eV}^2} \right\}$$

0.75 upgrade to 4 MW

Off-Axis Beams

BNL 1994



π^0 suppression

T2K

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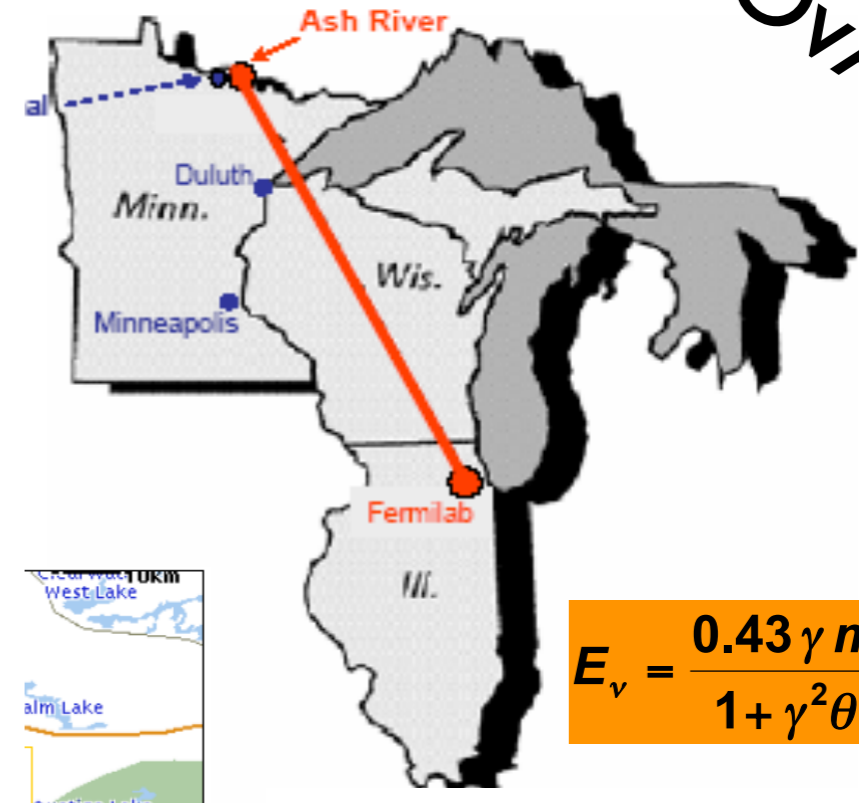


L=295 km and
Energy at Vac. Osc. Max. (vom)

$$E_{vom} = 0.6 \text{ GeV} \left\{ \frac{\delta m_{32}^2}{2.5 \times 10^{-3} \text{ eV}^2} \right\}$$

0.75 upgrade to 4 MW

NOVA



$$E_\nu = \frac{0.43 \gamma m_\pi}{1 + \gamma^2 \theta^2}$$

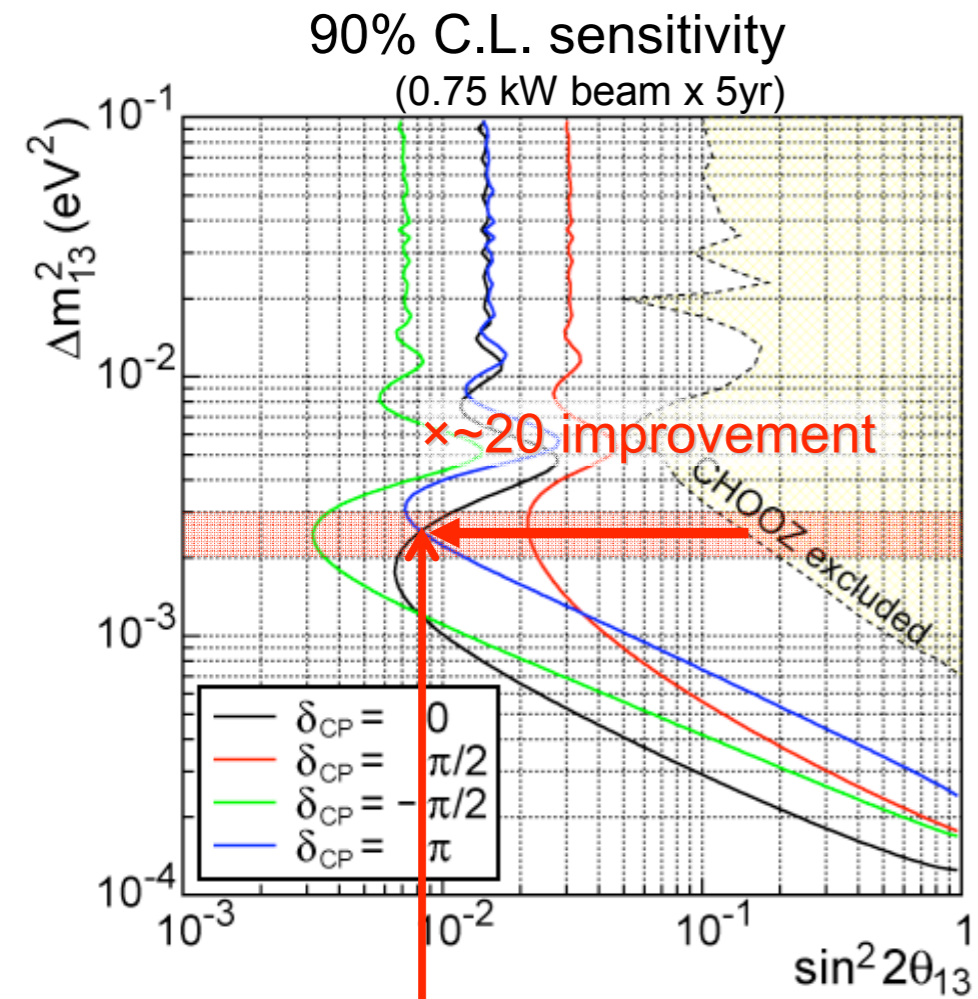
L=700 - 1000 km and
Energy near 2 GeV

$$E_{vom} = 1.8 \text{ GeV} \left\{ \frac{\delta m_{32}^2}{2.5 \times 10^{-3} \text{ eV}^2} \right\} \times \left\{ \frac{L}{820 \text{ km}} \right\}$$

0.4 upgrade to 2 MW

T2K

Search for ν_e appearance



$\sin^2 2\theta_{13} \sim 0.008$ ($\delta_{\text{CP}} = 0, \pi$)

T2K

For LARGE δm_{31}^2

$$\langle P(\nu_\mu \rightarrow \nu_e) \rangle = \frac{1}{2} \sin^2 \theta_{23} \sin^2 2\theta_{13} - \frac{1}{2} J \cdot \Delta_{21} + \cos^2 \theta_{23} \sin^2 2\theta_{12} \Delta_{12}^2$$

$$J = \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13} \sin \delta$$

$$\Delta_{21} = \delta m_{21}^2 L / 4E$$

At $\delta = 0$ or π

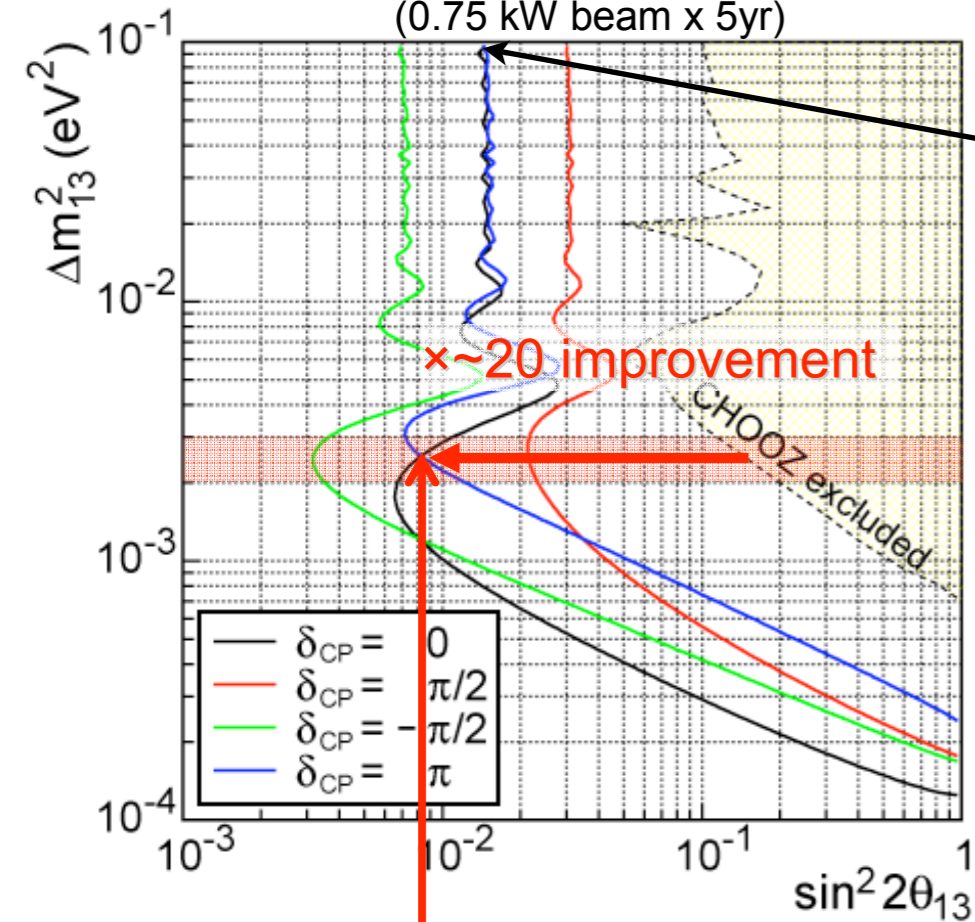
$$\langle P(\nu_\mu \rightarrow \nu_e) \rangle = \frac{1}{2} \sin^2 \theta_{23} \sin^2 2\theta_{13} + \cos^2 \theta_{23} \sin^2 2\theta_{12} \Delta_{12}^2 \approx 0.5\%$$

$$\langle P(\nu_\mu \rightarrow \nu_e) \rangle_{T2K} \approx 0.5\%$$

0.5% ν_e in beam

Search for ν_e appearance

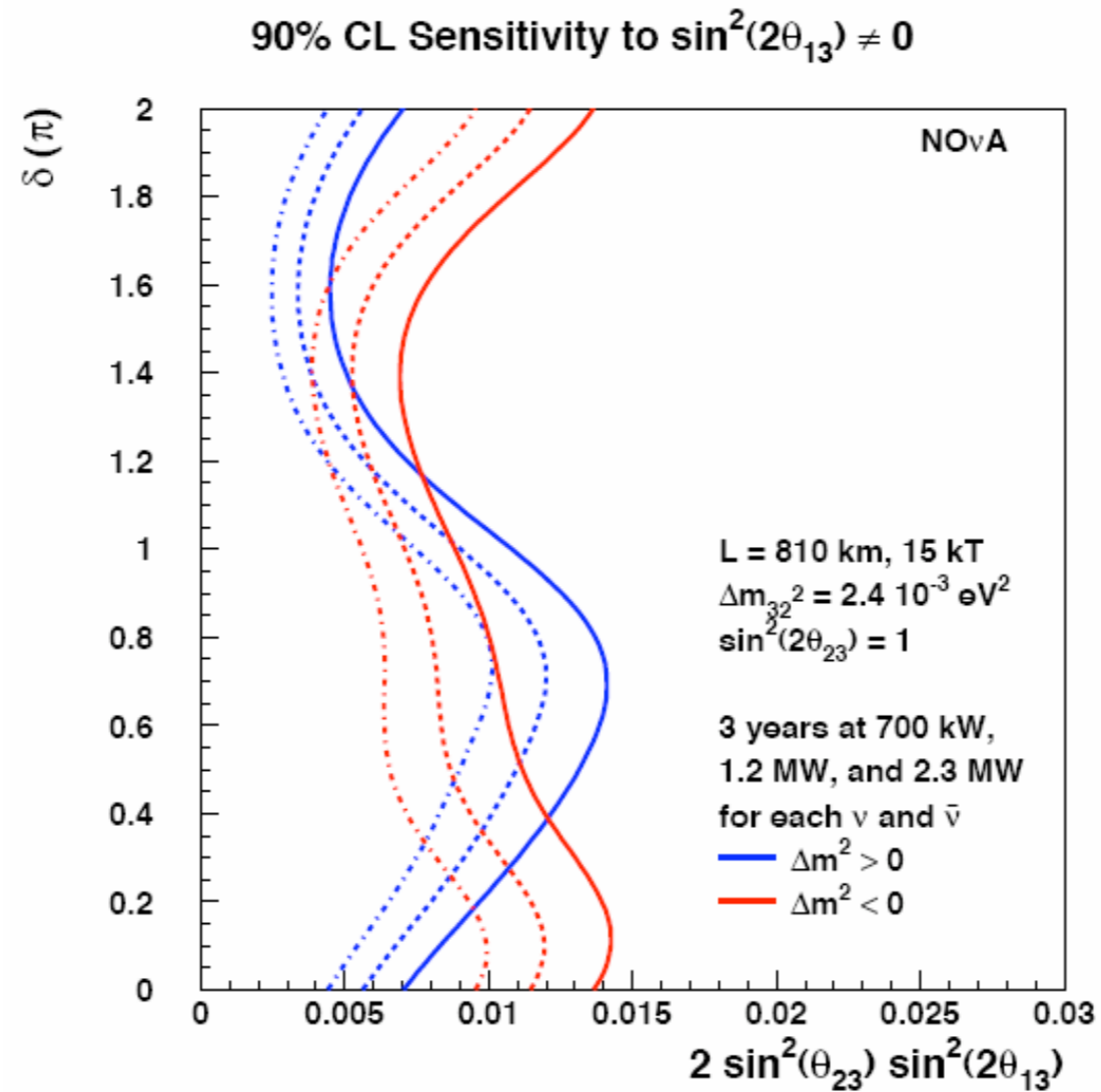
90% C.L. sensitivity
(0.75 kW beam x 5yr)



$$\sin^2 2\theta_{13} \sim 0.008 \quad (\delta_{CP} = 0, \pi)$$

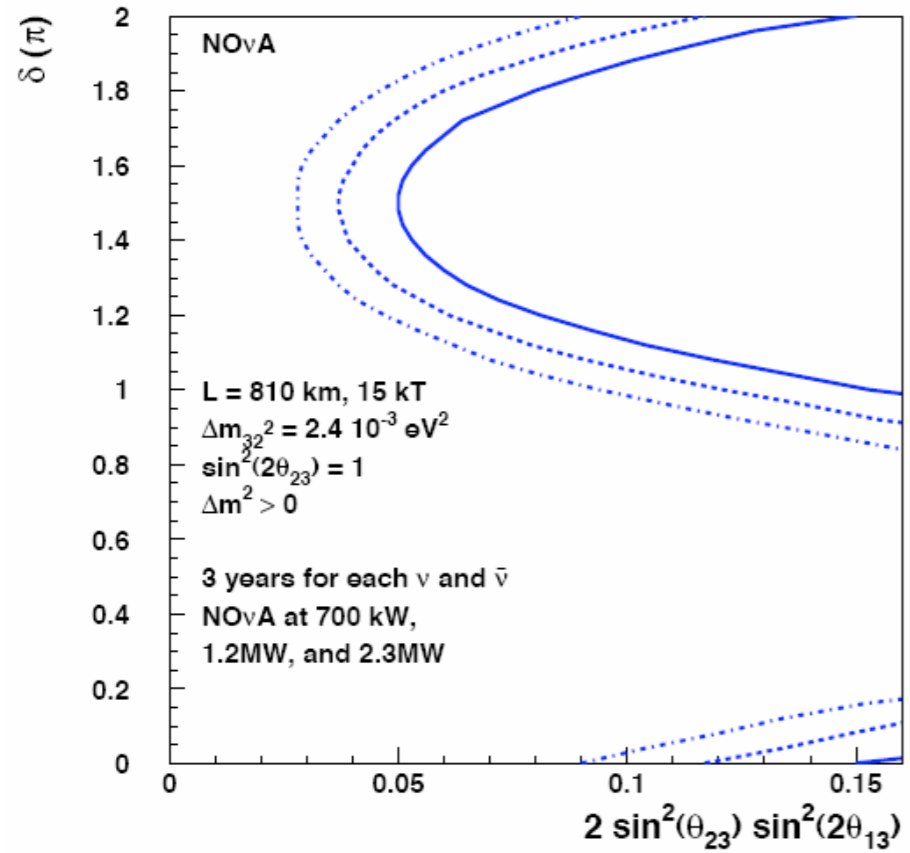


Sensitivity to $\sin^2(2\theta_{13}) \neq 0$





95% CL Resolution of the Mass Ordering
NOvA Alone

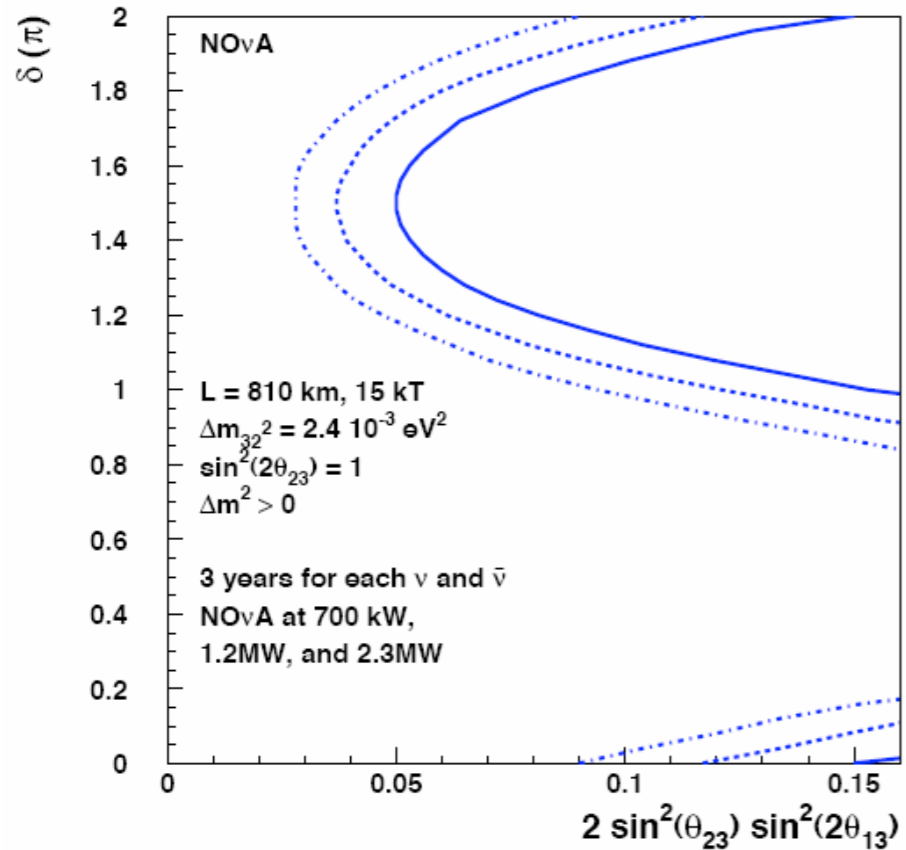


Normal Ordering

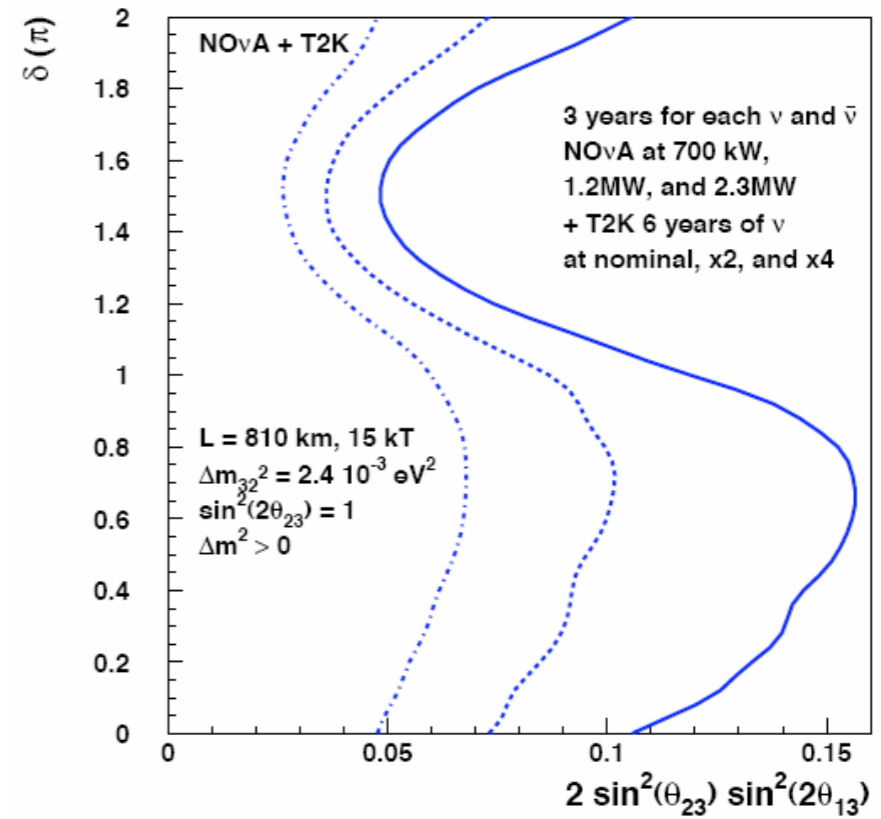


95% CL Resolution of the Mass Ordering
NOvA Alone

95% CL Resolution of the Mass Ordering
NOvA Plus T2K



Normal Ordering

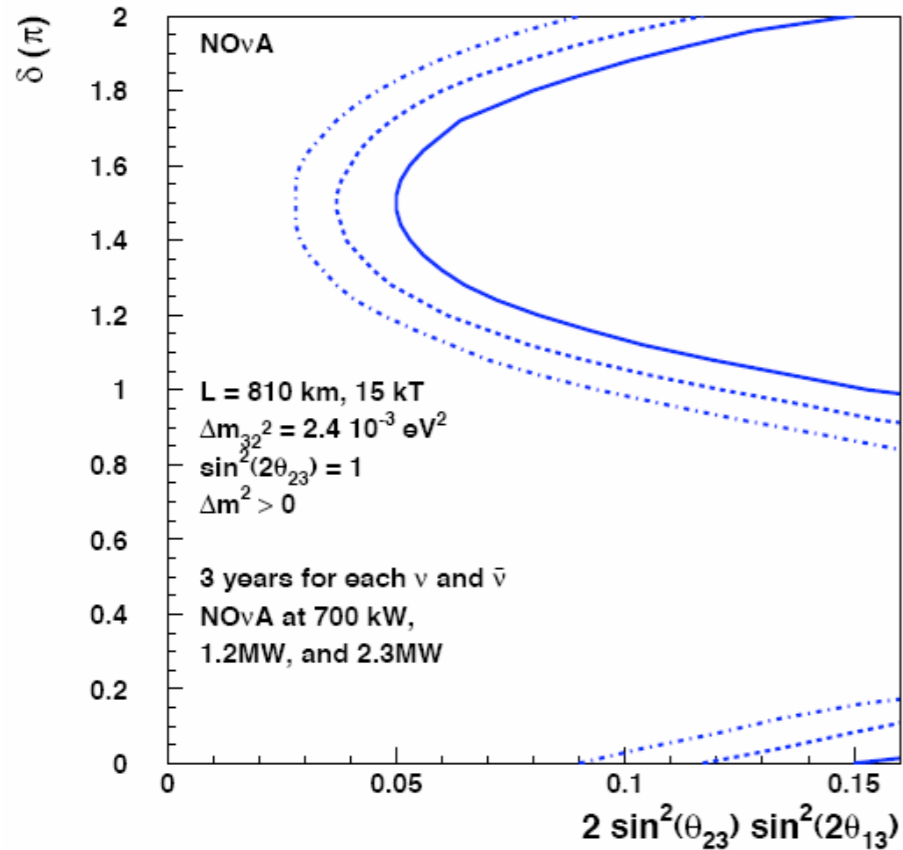


Normal Ordering

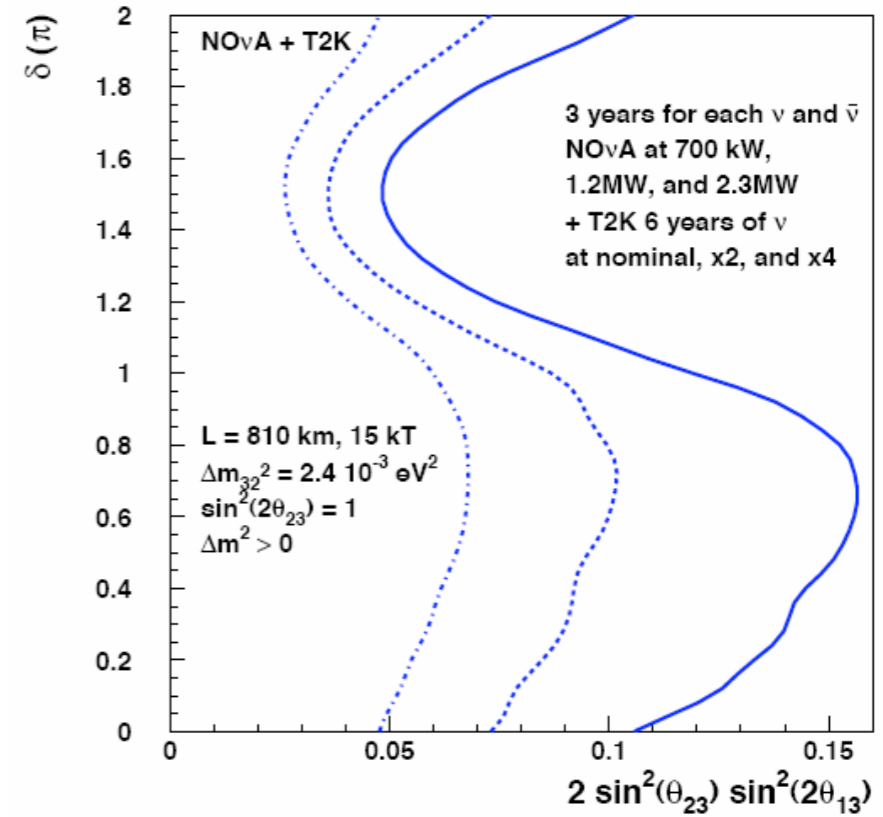


95% CL Resolution of the Mass Ordering
NOvA Alone

95% CL Resolution of the Mass Ordering
NOvA Plus T2K



Normal Ordering



Normal Ordering

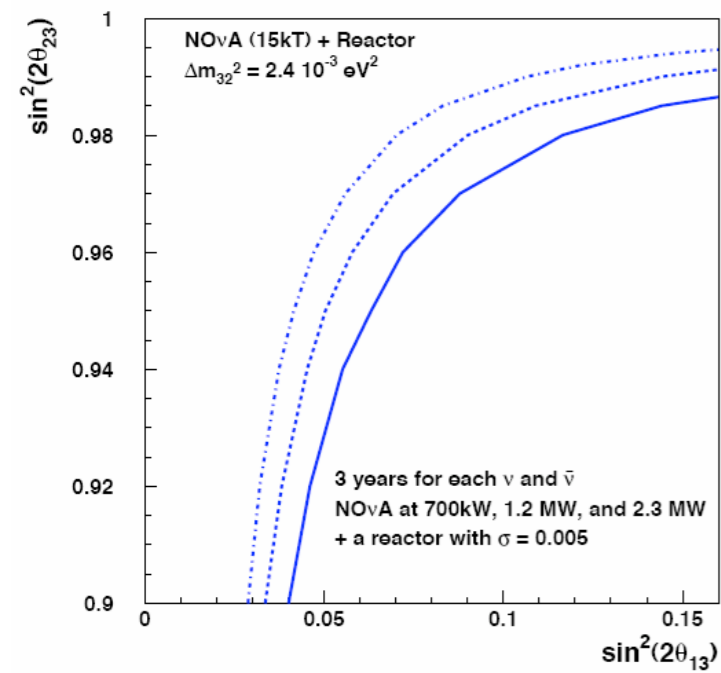
for Inverted Hierarchy $\delta \rightarrow \pi - \delta$

95% CL Resolution of the θ_{23} Ambiguity

at Vac. Osc. Max.

$$P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \approx 2 \sin^2 \theta_{23} \sin^2 2\theta_{13} + 2P_\odot$$

in $P + \bar{P}$ the matter effects approx. cancel
and CP effects approx. cancel.

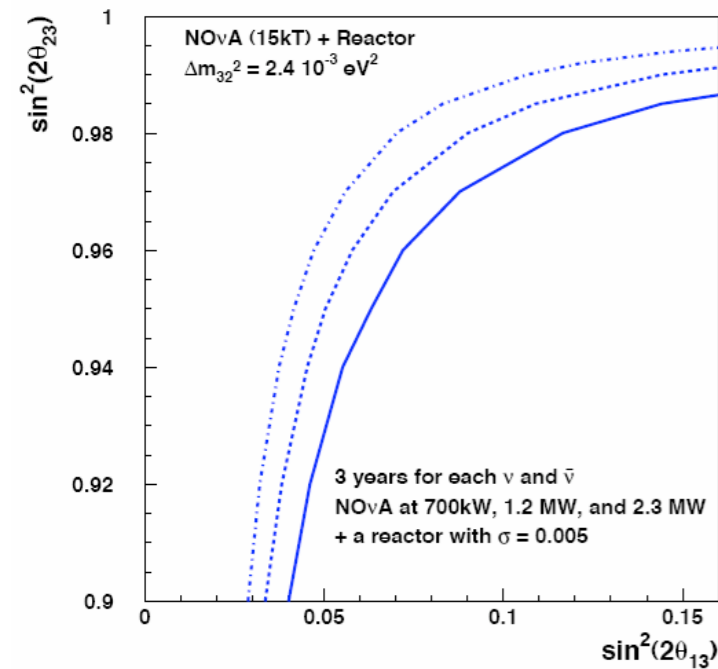


95% CL Resolution of the θ_{23} Ambiguity

at Vac. Osc. Max.

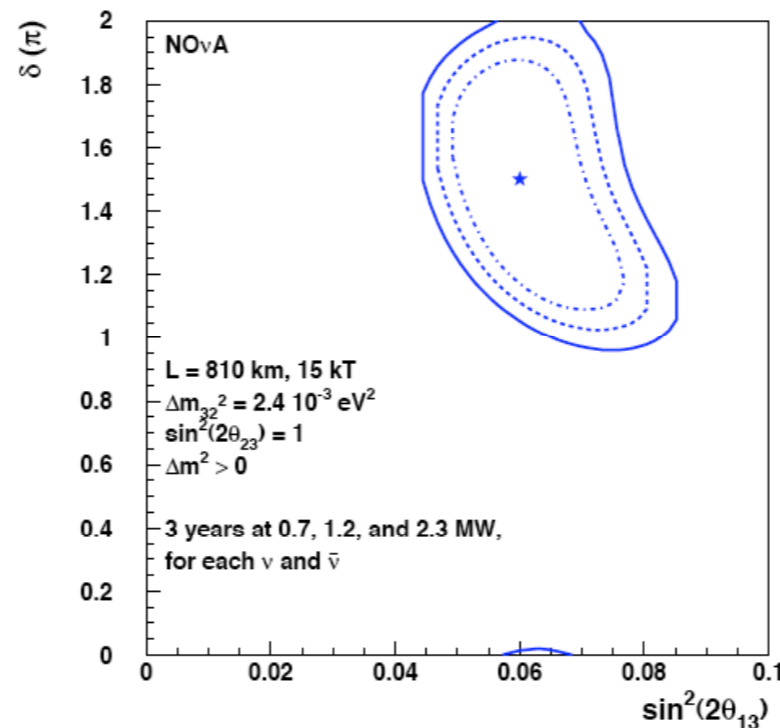
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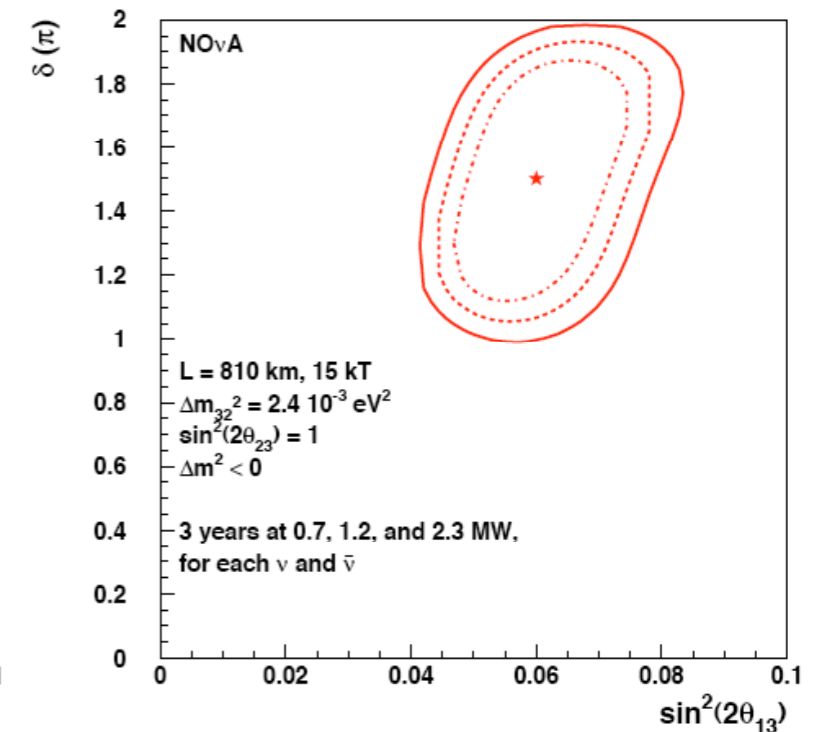


CP Violation: δ vs. θ_{13} Contours

1 σ Contours for Starred Point



1 σ Contours for Starred Point

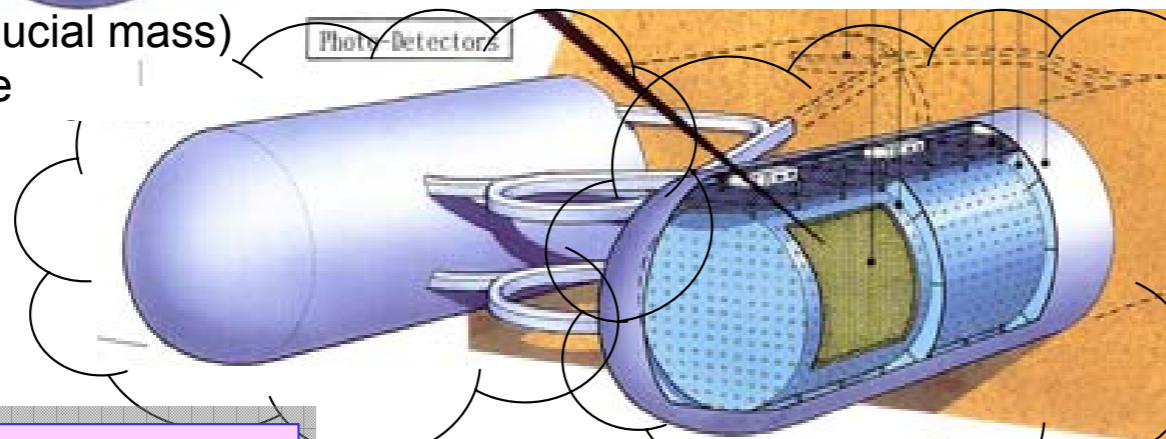


Off Axis:



Some recent progress: detector in Korea

1Mton (0.54Mton fiducial mass)
Hyper-Kamiokande

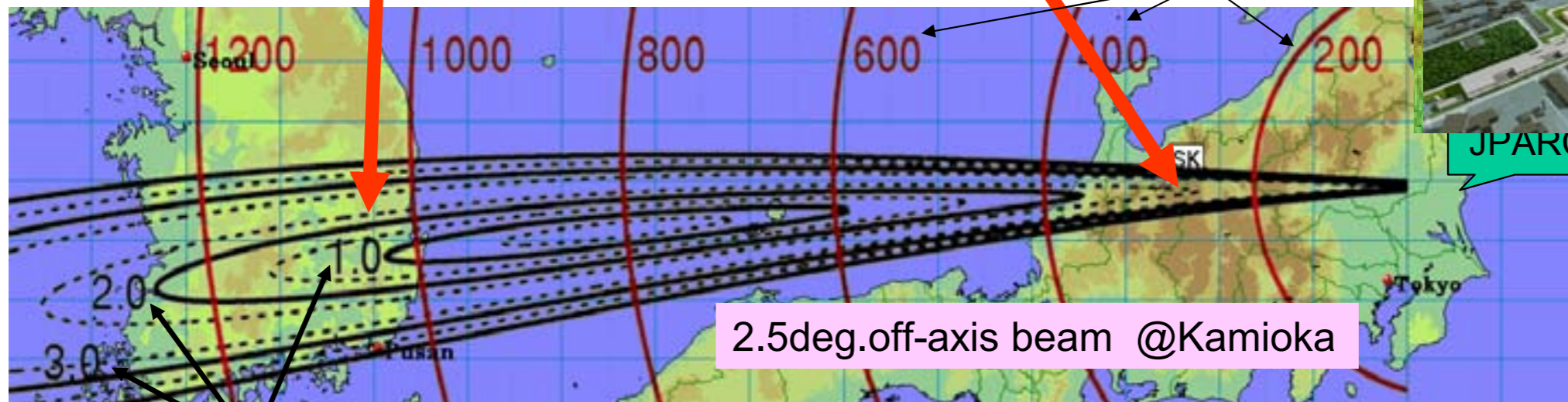


Total cost must be similar to the baseline design.

2.5 deg. off axis

2.5 deg. off axis

Distance from the target (km)

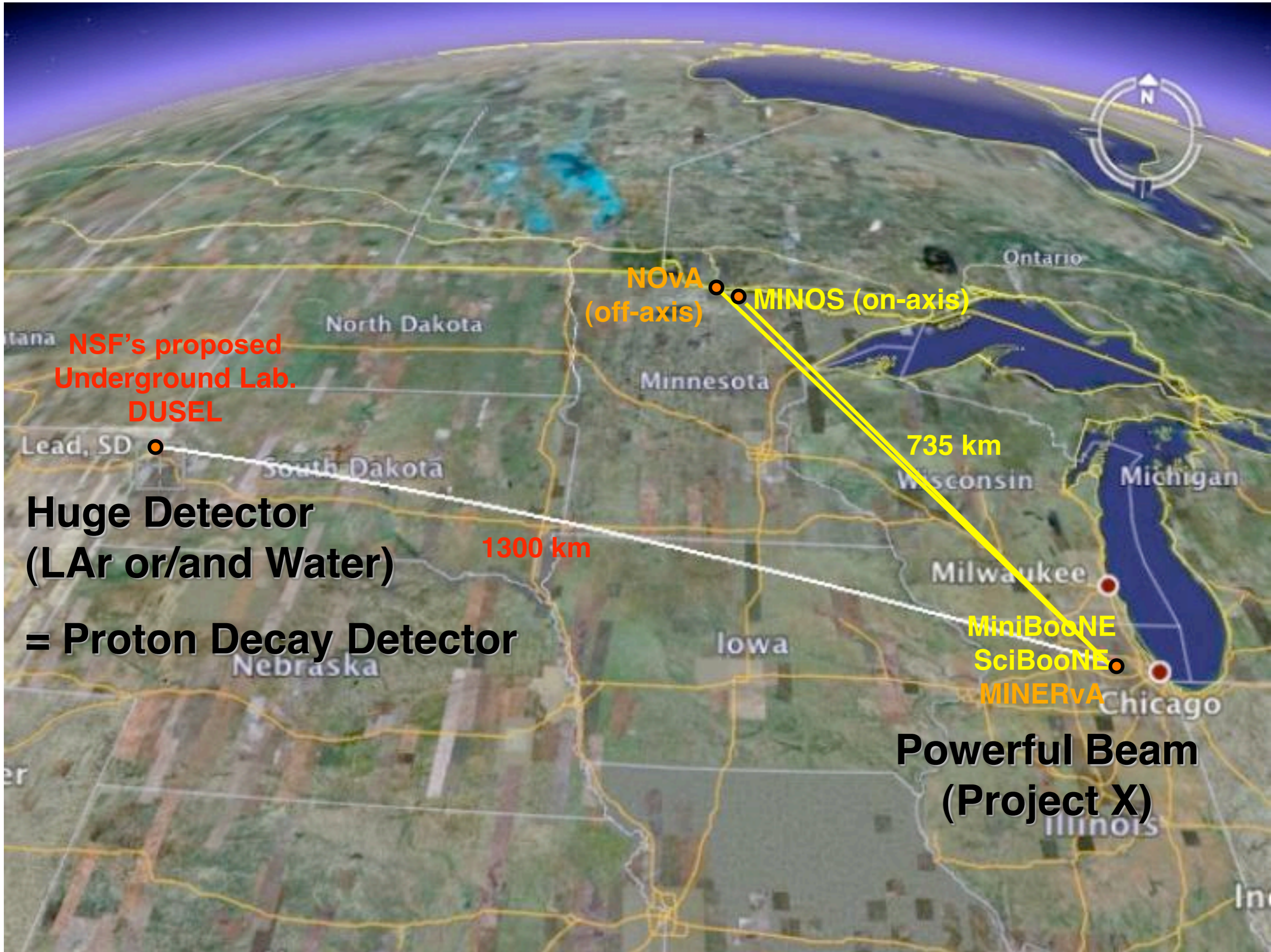


JPARC

2.5deg.off-axis beam @Kamioaka

Off-axis angle

see Kajita talk:



NSF's proposed
Underground Lab.
DUSEL

**Huge Detector
(LAr or/and Water)
= Proton Decay Detector**

NOvA
(off-axis)

MINOS (on-axis)

735 km

1300 km

MiniBooNE
SciBooNE
MINERvA

**Powerful Beam
(Project X)**

Narrow Band Beam: Same E, Longer L T2KK

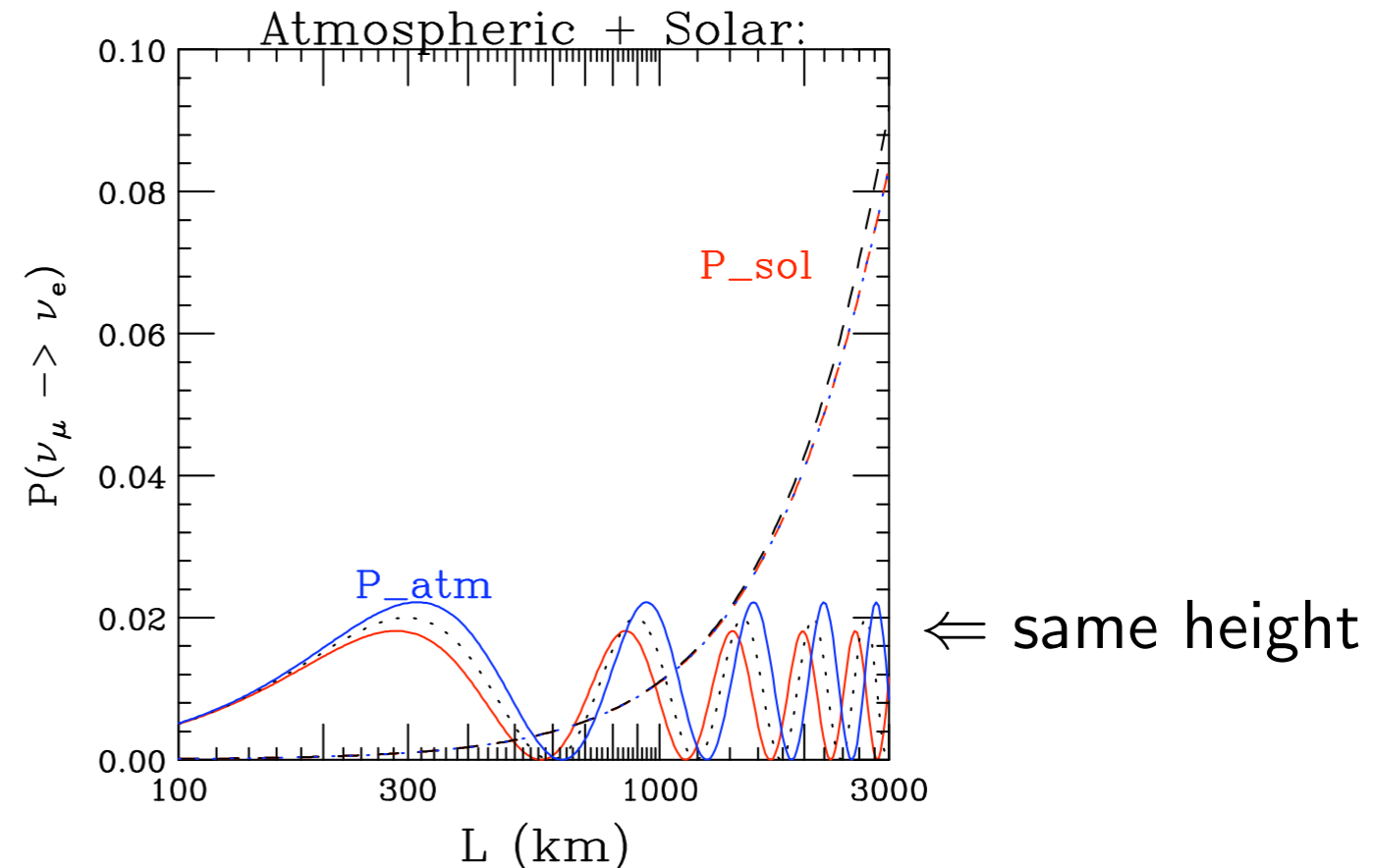
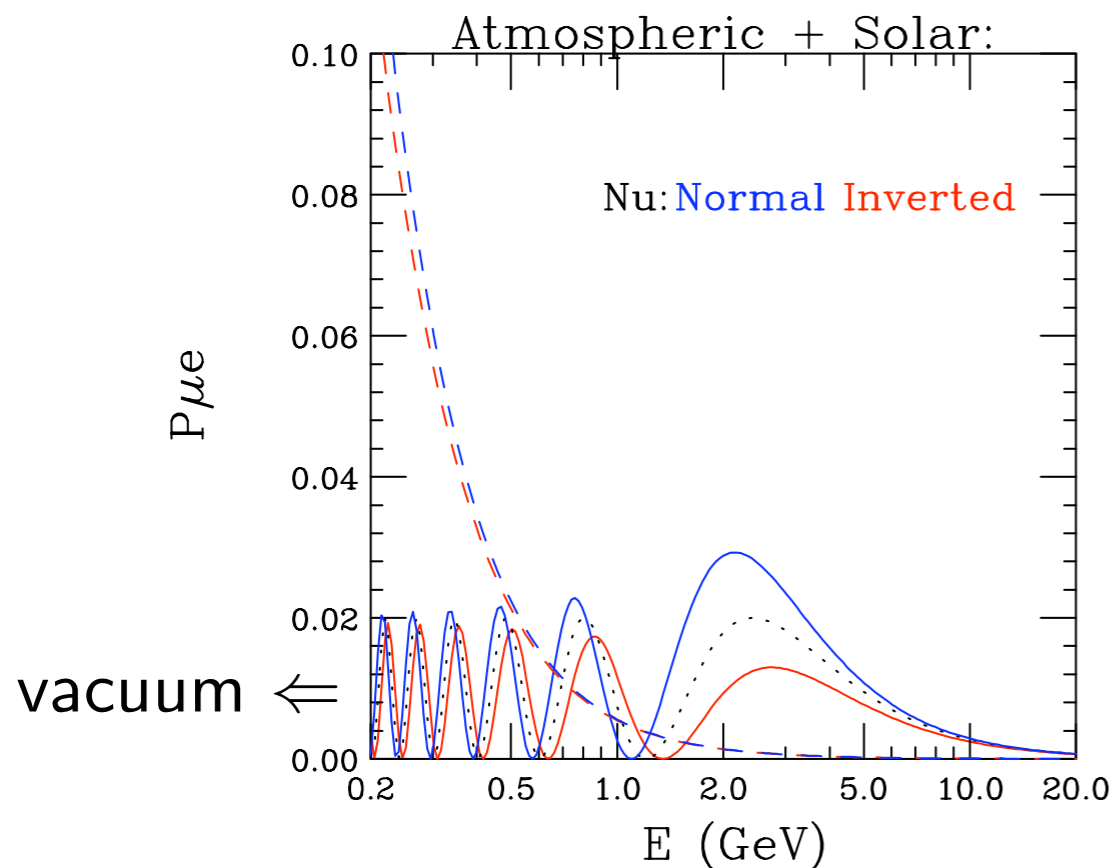
Broadband Beam: Same L, Lower E Fermilab to DUSEL

In VACUUM the SAME but NOT in MATTER

$$\sin^2 2\theta_{13} = 0.04$$

L=1200km

E=0.6 GeV

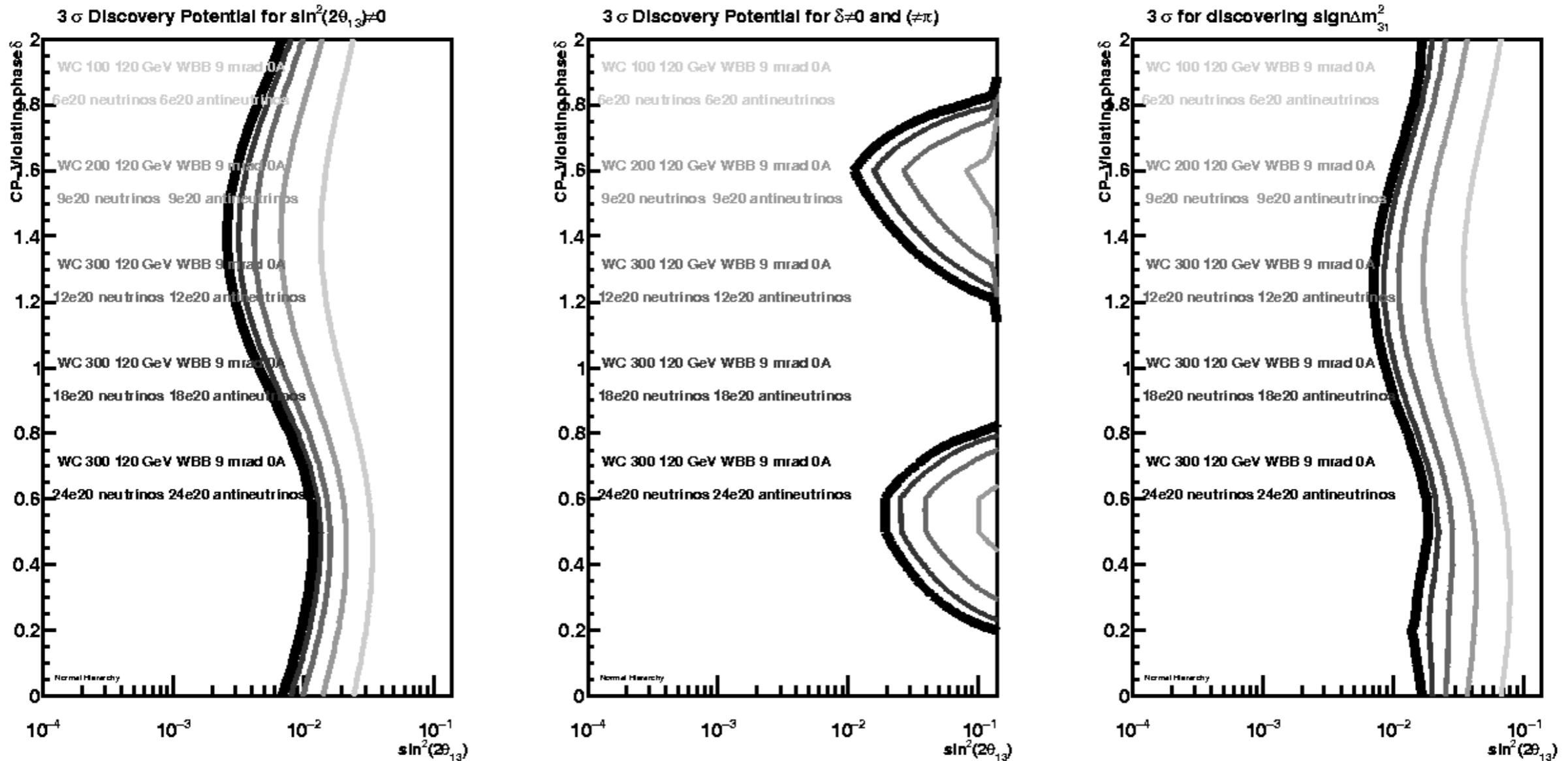


$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

Sensitivities for Water Cherenkov Detector :



Scenario 1 “Detector Module” = 100 KT

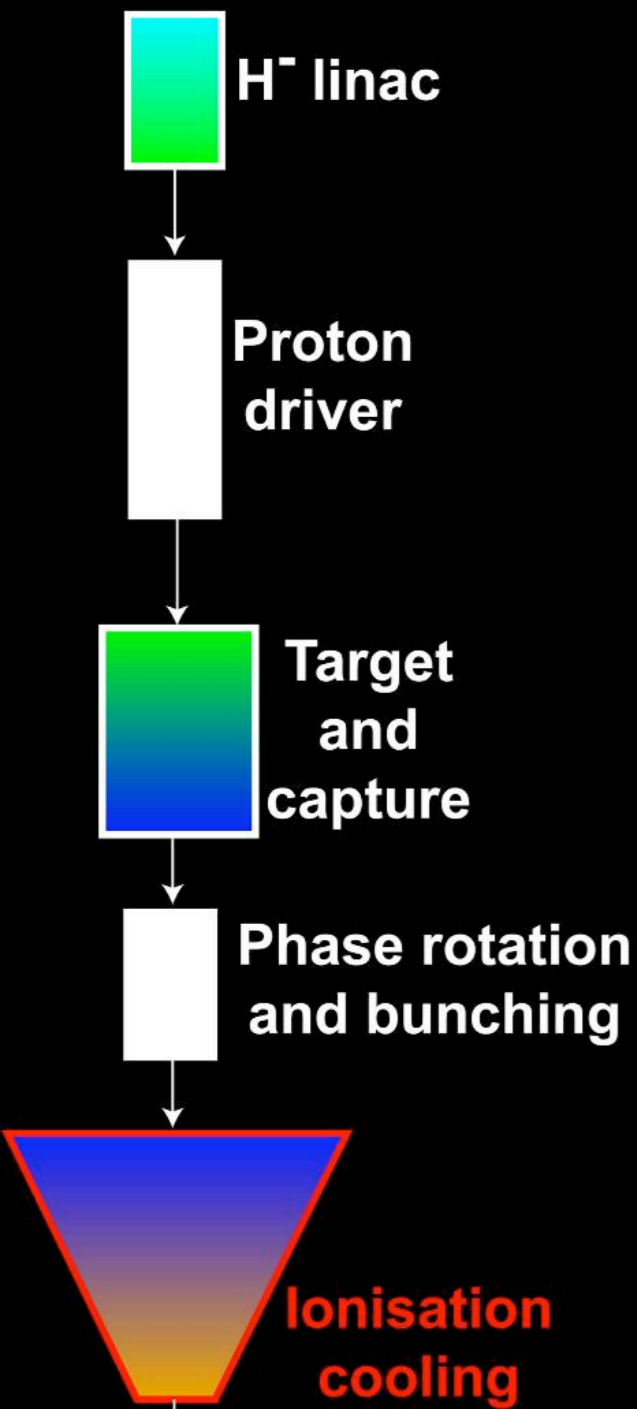


The 5 Curves correspond to the five 2 year intervals for the total 10 year period.

First curve is discovery potential after 2 years of running, second curve is discovery potential after 4 years of running etc.

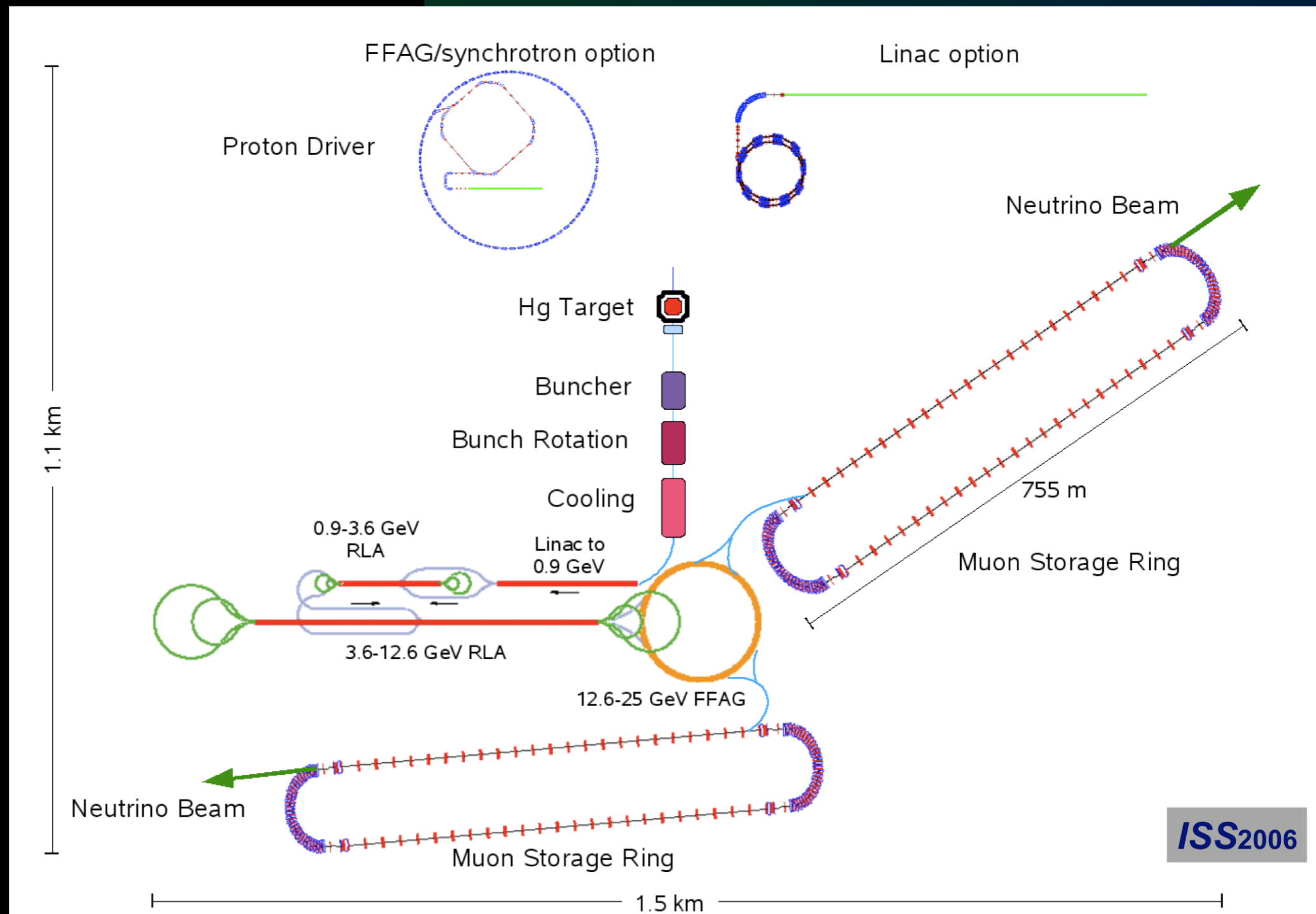
Neutrino Factory:

IDS-NF baseline: accelerator



Muon acceleration

Muon storage



where $\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31}$

and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{(aL)} \Delta_{21}$

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

depends on θ_{13}
amplification or suppression
by matter (E)

Suppression \geq Enhancement

independent of θ_{13}
 \approx independent of
matter effect

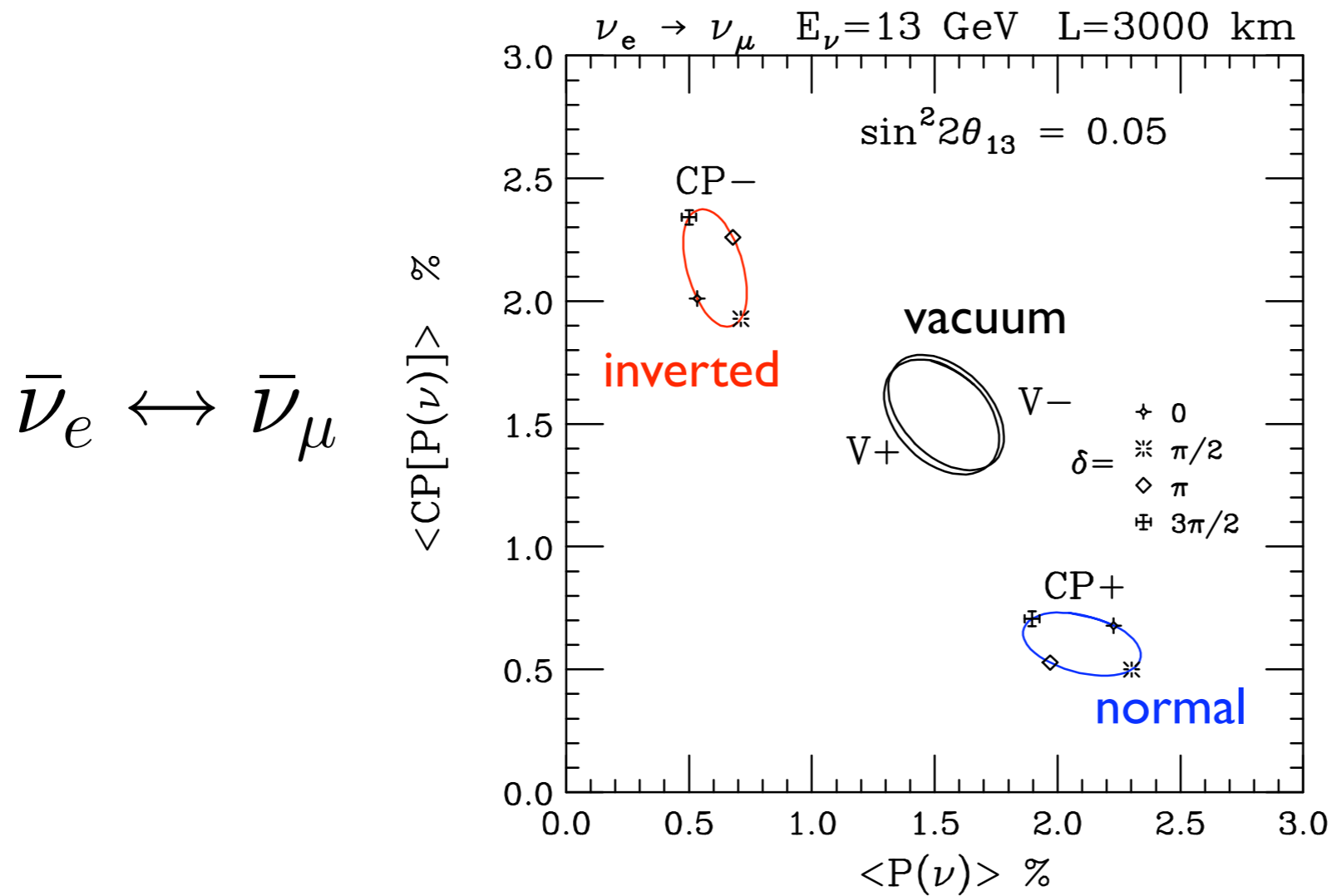
when $\sin(aL)/(aL) \approx 1$

$L/E \geq$ significant fraction of 500 km/GeV

Event rate: $E(E/L)^2$

Mass Hierarchy: – sign of δm_{31}^2

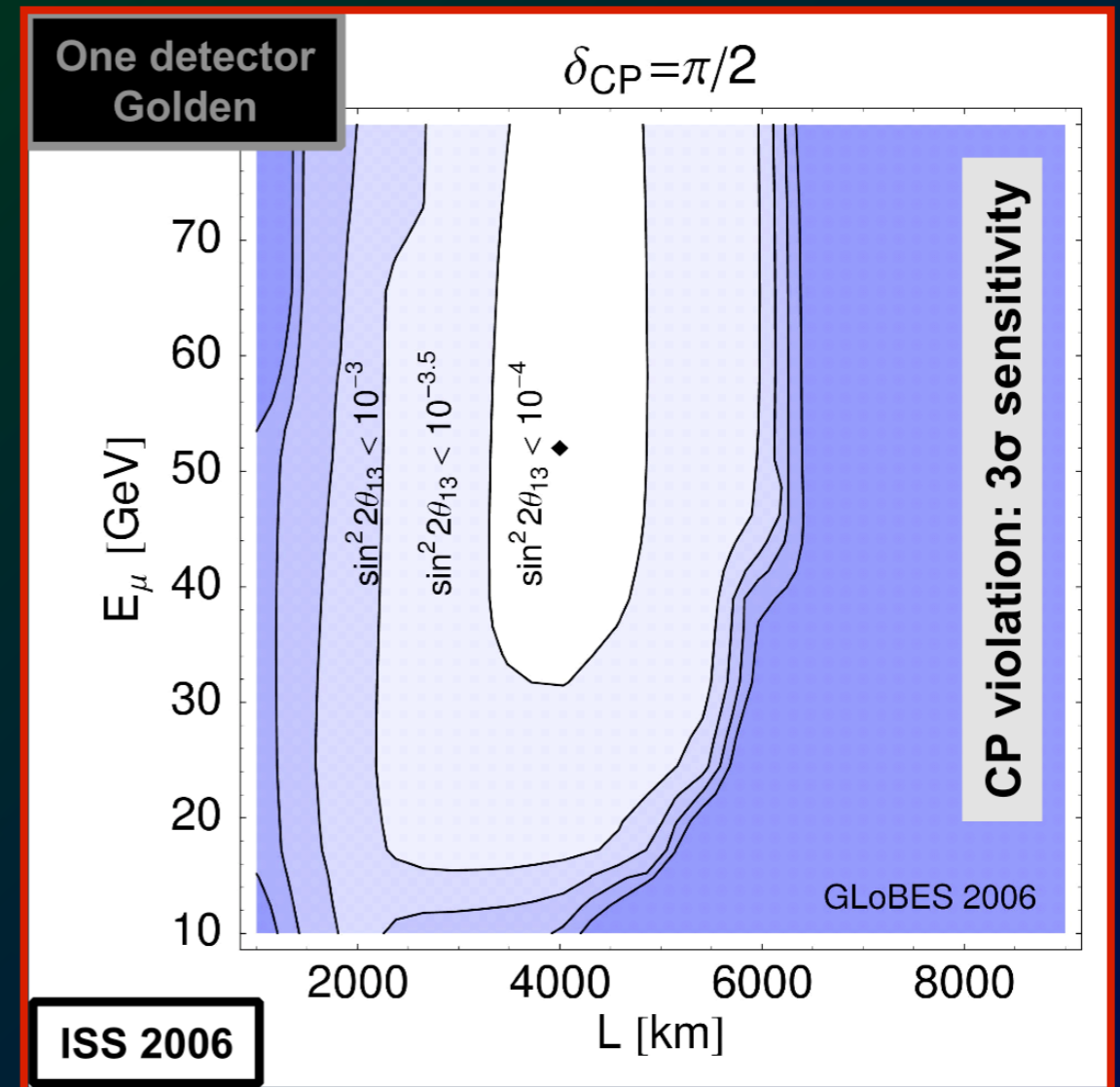
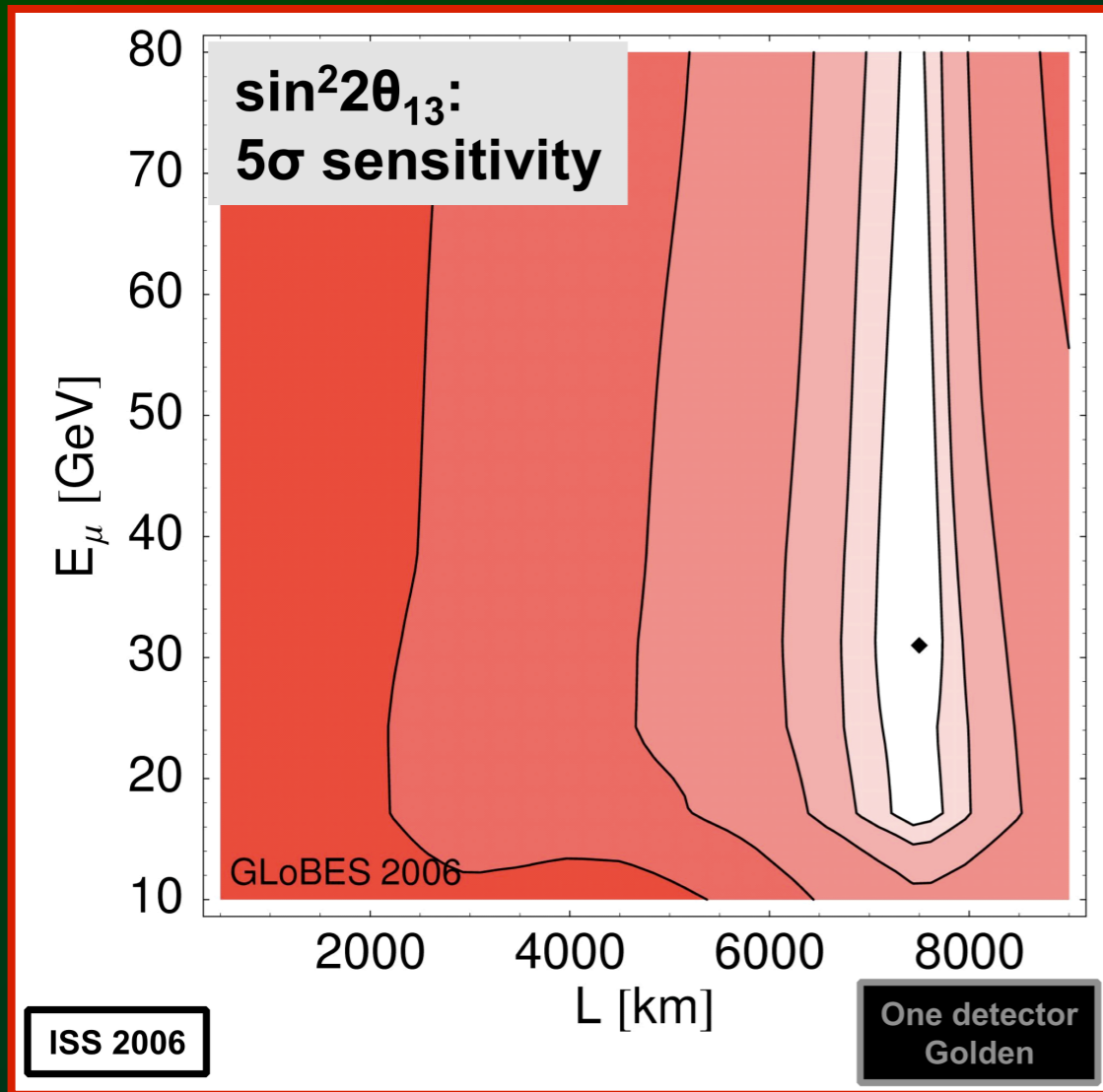
Matter Effects



$$\nu_e \leftrightarrow \nu_\mu$$

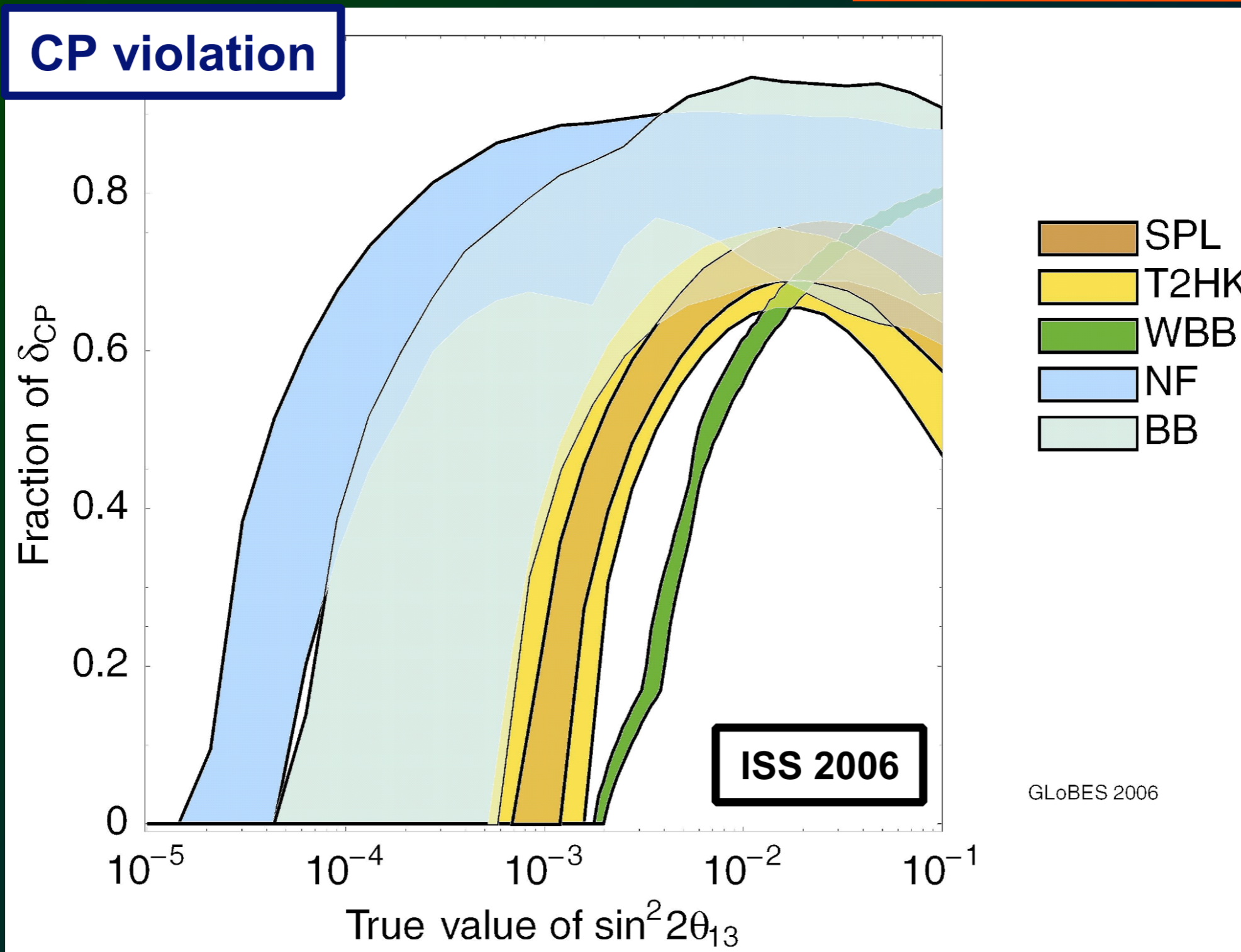
Suppression \geq Enhancement

NF: $\nu_e \rightarrow \nu_\mu$ (golden channel) optimisation



- Magic baseline (7500 km) good degeneracy solver
- Best sensitivity to CP requires baseline ~ 4000 km
- Stored muon energy: 25 GeV

Comparison:



Neutrino Factory

Golden, 4000,
 $E_\mu = 50$ GeV
 Golden* (4000 km), Golden* (7500 km)
 $E_\mu = 20$ GeV

Beta beam

$\gamma = 100$
 500 kT H₂O ζ (130 km)

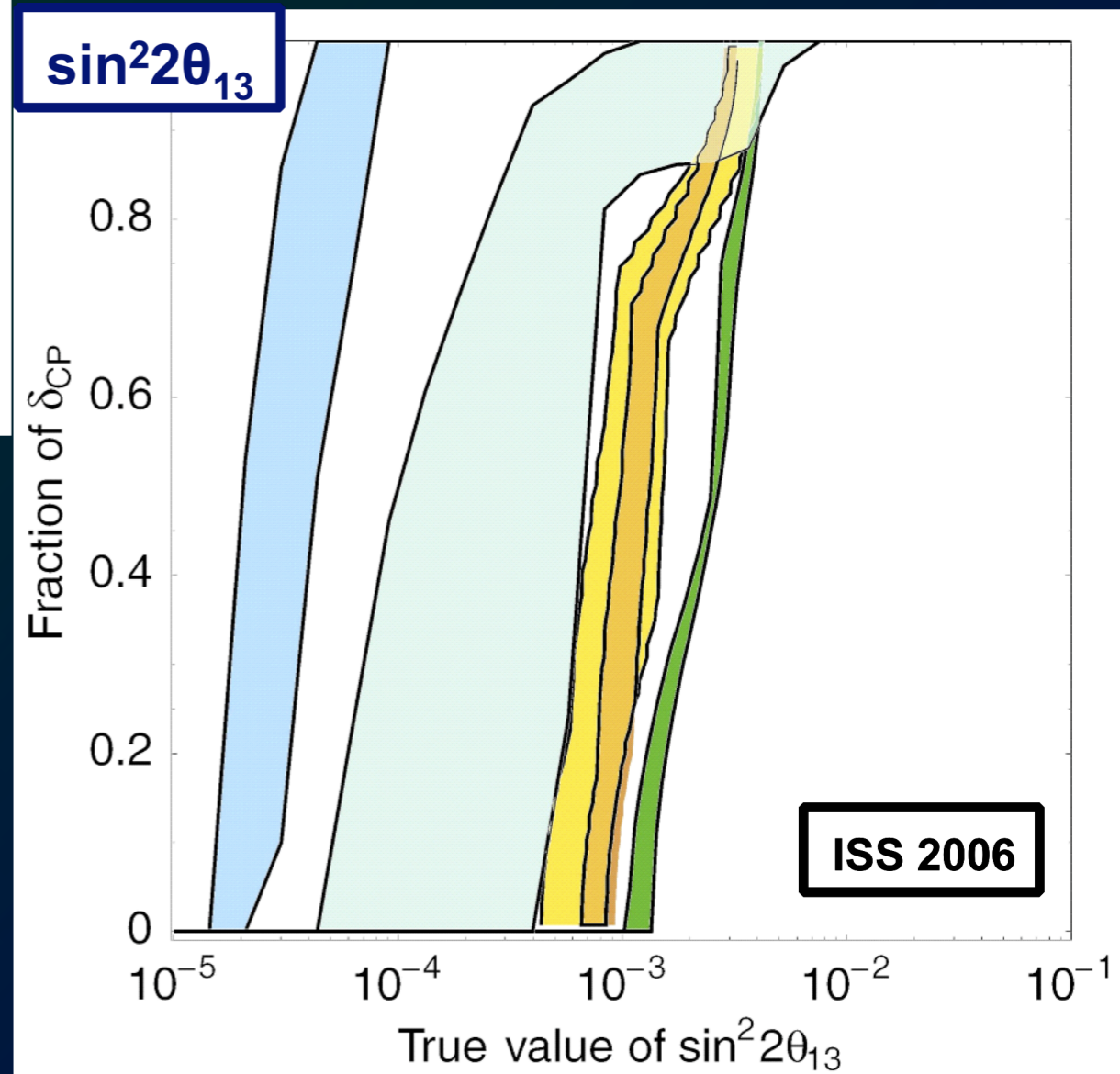
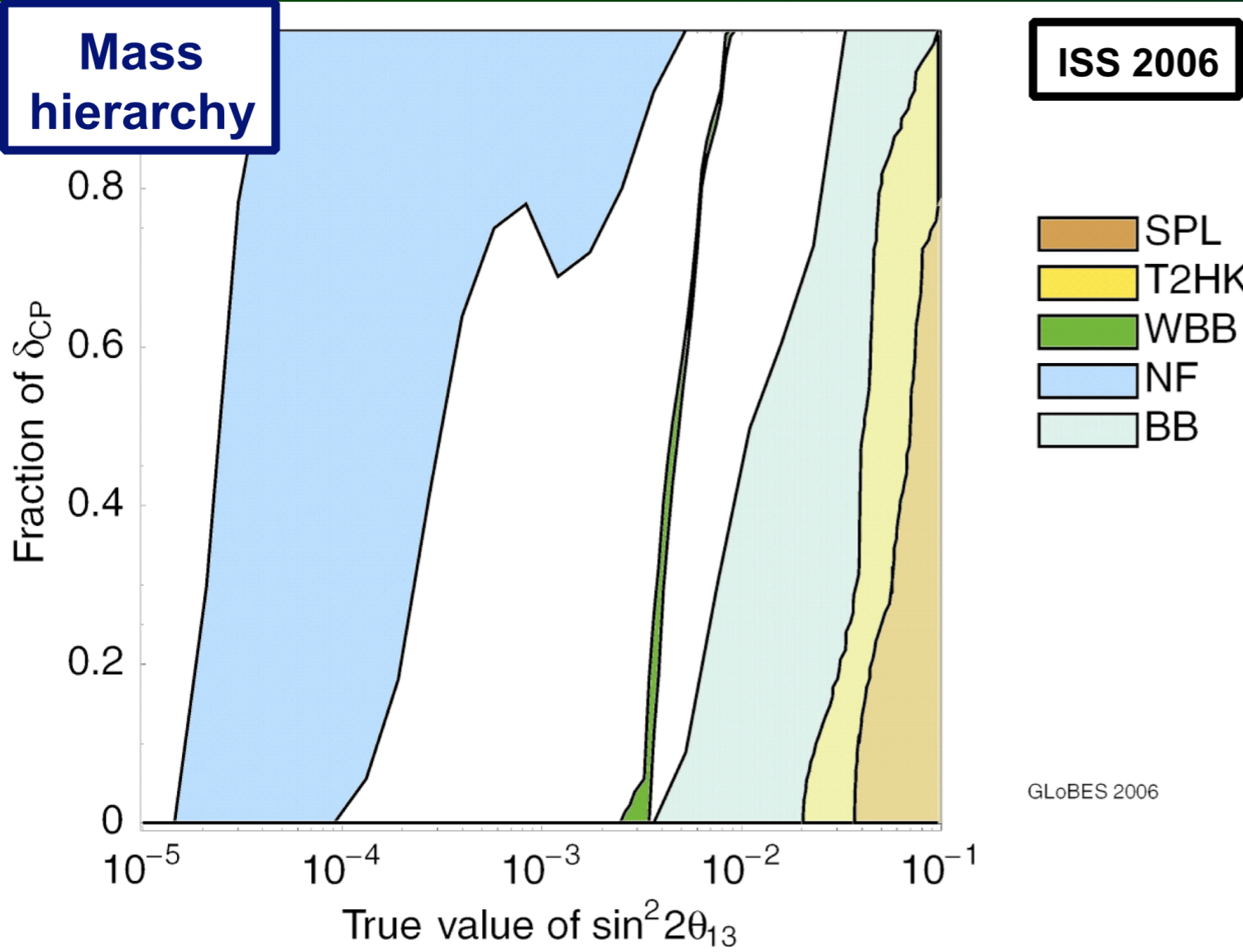
$\gamma = 350$
 500 kT H₂O ζ (730 km)

WBB
 Systematics from
 proposal

T2HK
 Systematics: 2% – 5%

SPL
 Systematics: 2% – 5%

Comparison:



Neutrino Factory

Golden, 4000,
 $E_\mu = 50$ GeV
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WBB
 Systematics from proposal

T2HK
 Systematics: 2% – 5%

SPL
 Systematics: 2% – 5%

Neutrino Factory & Non-Standard Interactions: (NSI)

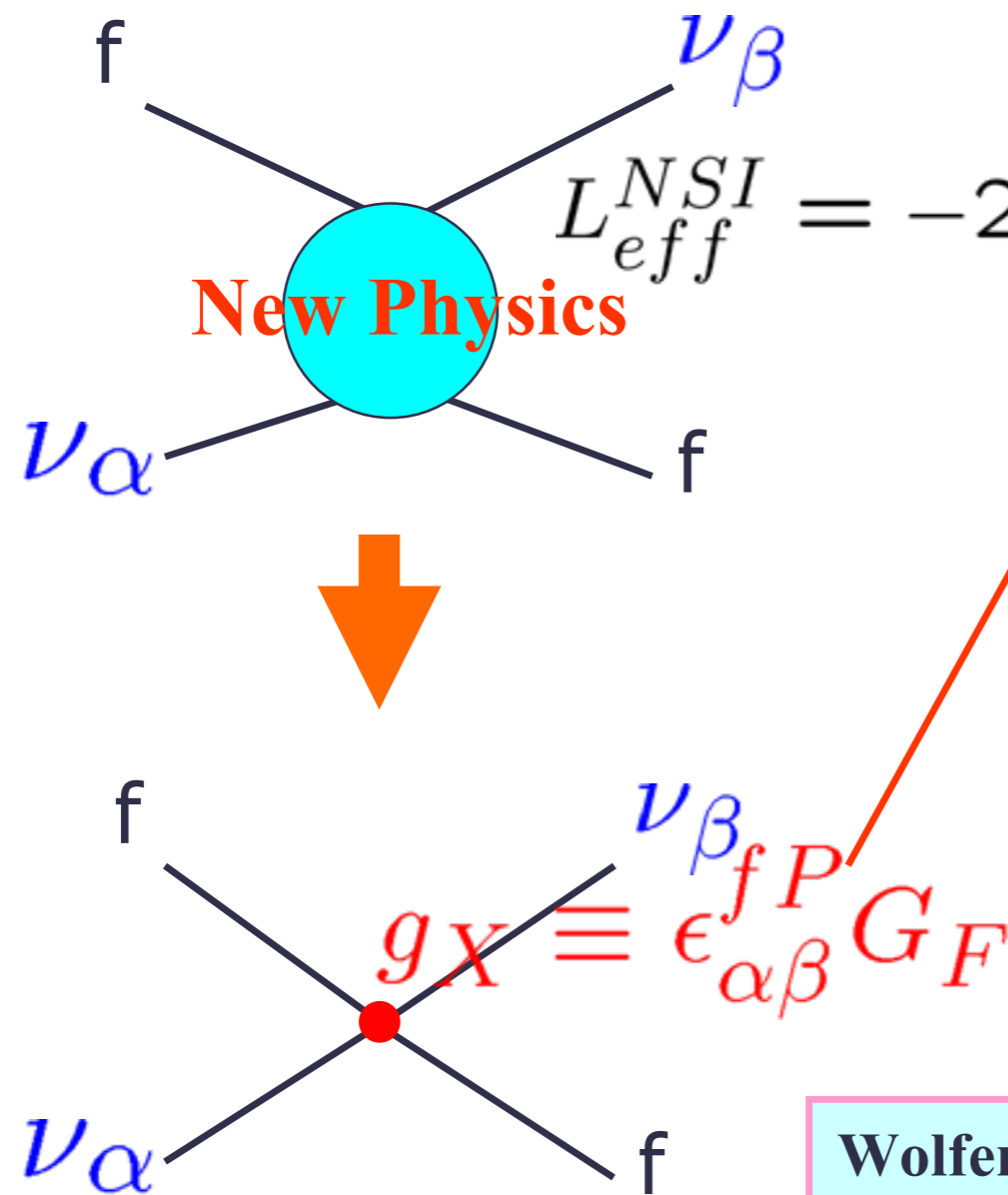
eg Z' at LHC

Non-Standard neutrino Interaction

Shoichi Uchinami

Nei Cipriano Ribeiro^{*1}, Hisakazu Minakata^{*2}, Hiroshi Nunokawa,^{*1} and Renata Zukanovich Funchal^{*3}

Non-Standard Interaction



$$L_{eff}^{NSI} = -2\sqrt{2}\epsilon_{\alpha\beta}^{fP} G_F (\bar{\nu}_\alpha \gamma_\mu P_L \nu_\beta) (\bar{f} \gamma^\mu P f)$$

$$P_L \equiv \frac{1}{2}(1 - \gamma_5), P_R \equiv \frac{1}{2}(1 + \gamma_5)$$

Naive Estimation $\epsilon_{\alpha\beta} \propto \frac{m_W^2}{m_X^2}$

If new physics scale $\sim 1(10)$ TeV

$$\epsilon_{\alpha\beta} \sim 0.01(0.0001)$$

Wolfenstein '78, Grossman '95, Berezhiani-Rossi '02 and many people...

We concentrated on effects of NSI in ν propagation in matter

$$H = \frac{1}{2E} \left[U \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix} U^{-1} + \begin{pmatrix} 2\sqrt{2}G_F n_e E & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} \right]$$

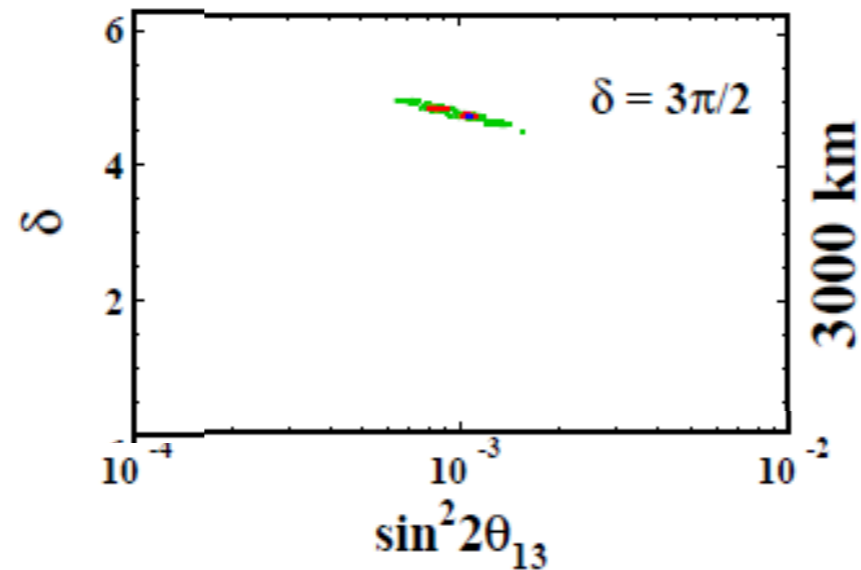


$$H = \frac{1}{2E} \left[U \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix} U^{-1} + 2Ea \begin{pmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{e\tau}^* & \epsilon_{\mu\tau}^* & \epsilon_{\tau\tau} \end{pmatrix} \right]$$

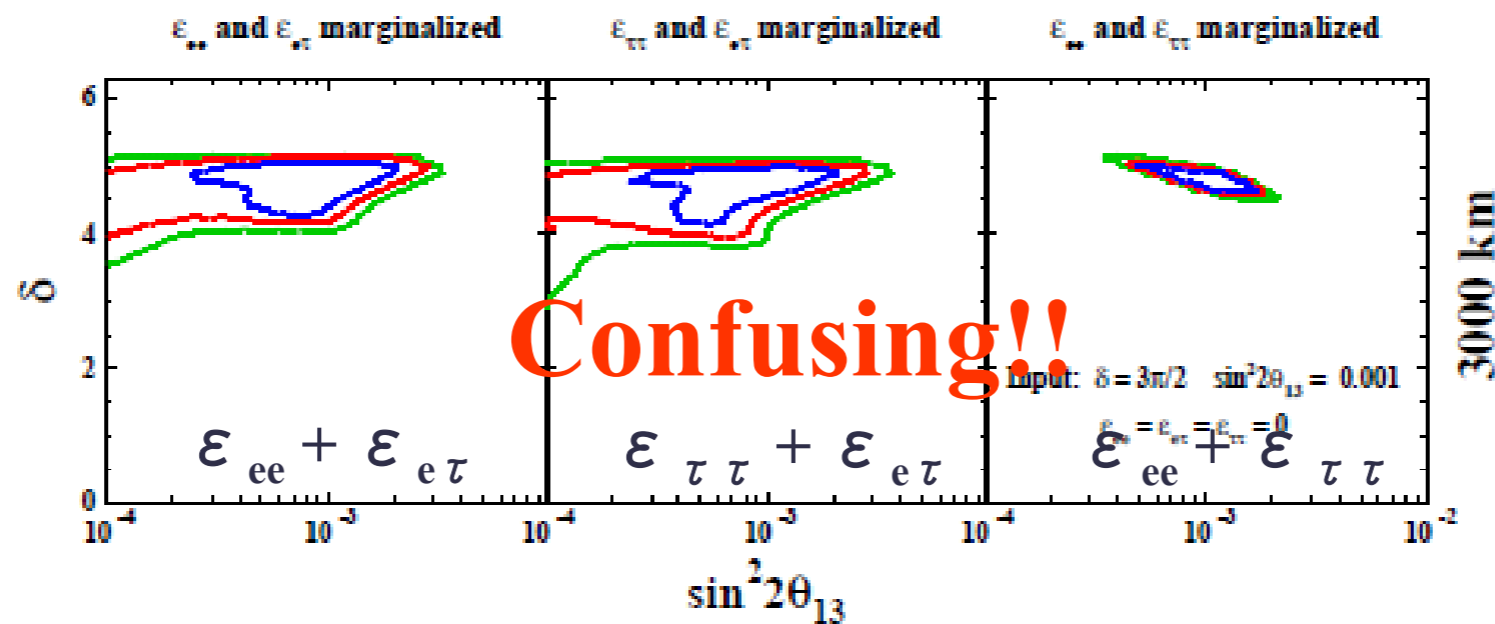
$$a \equiv \sqrt{2}G_F n_e$$

Valle, Gago-Guzzo-Nunokawa-Teves-Zukanovich Funchal

standard case



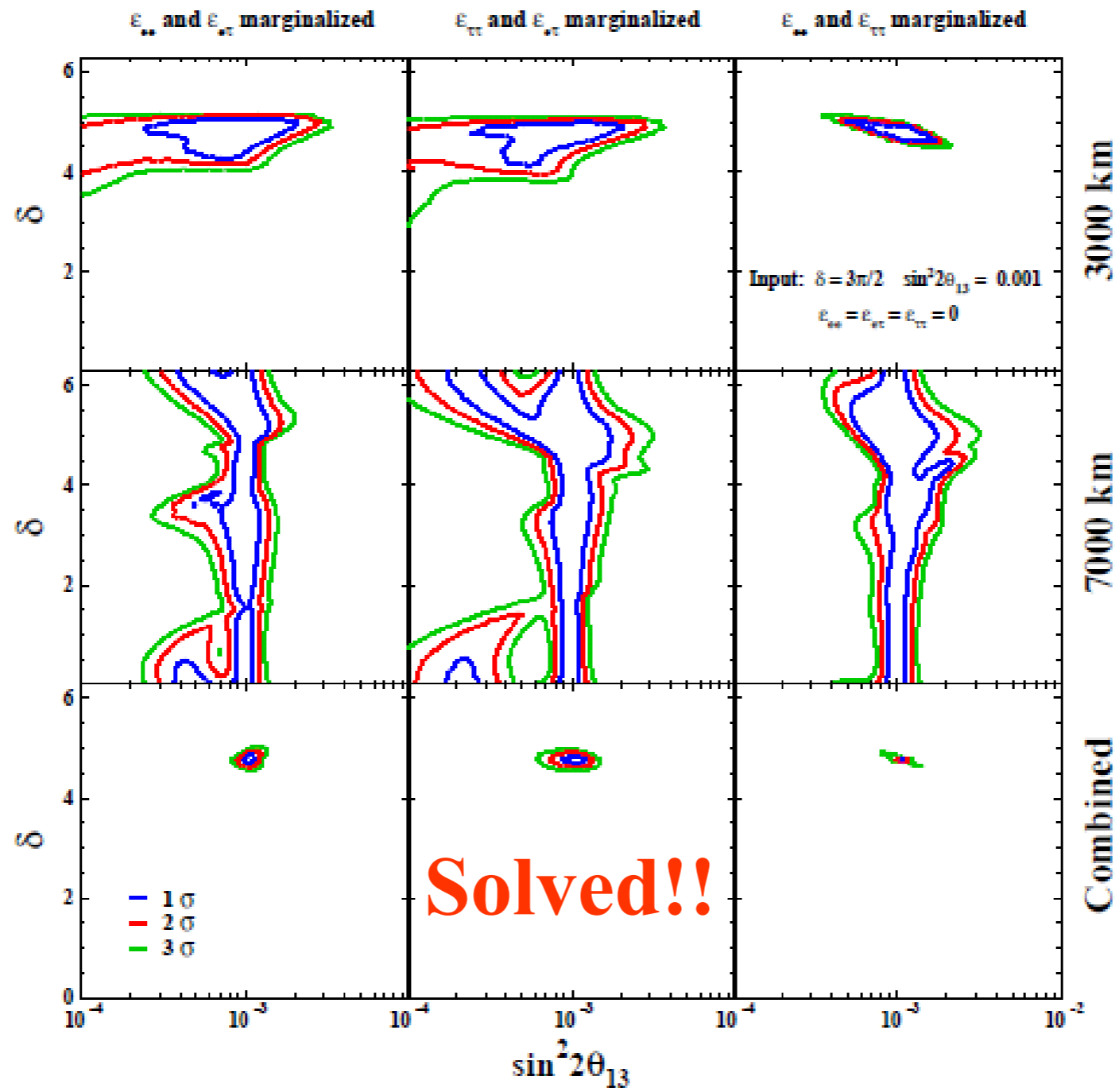
If we take account of the possible existence of NSI



all $\epsilon = 0$

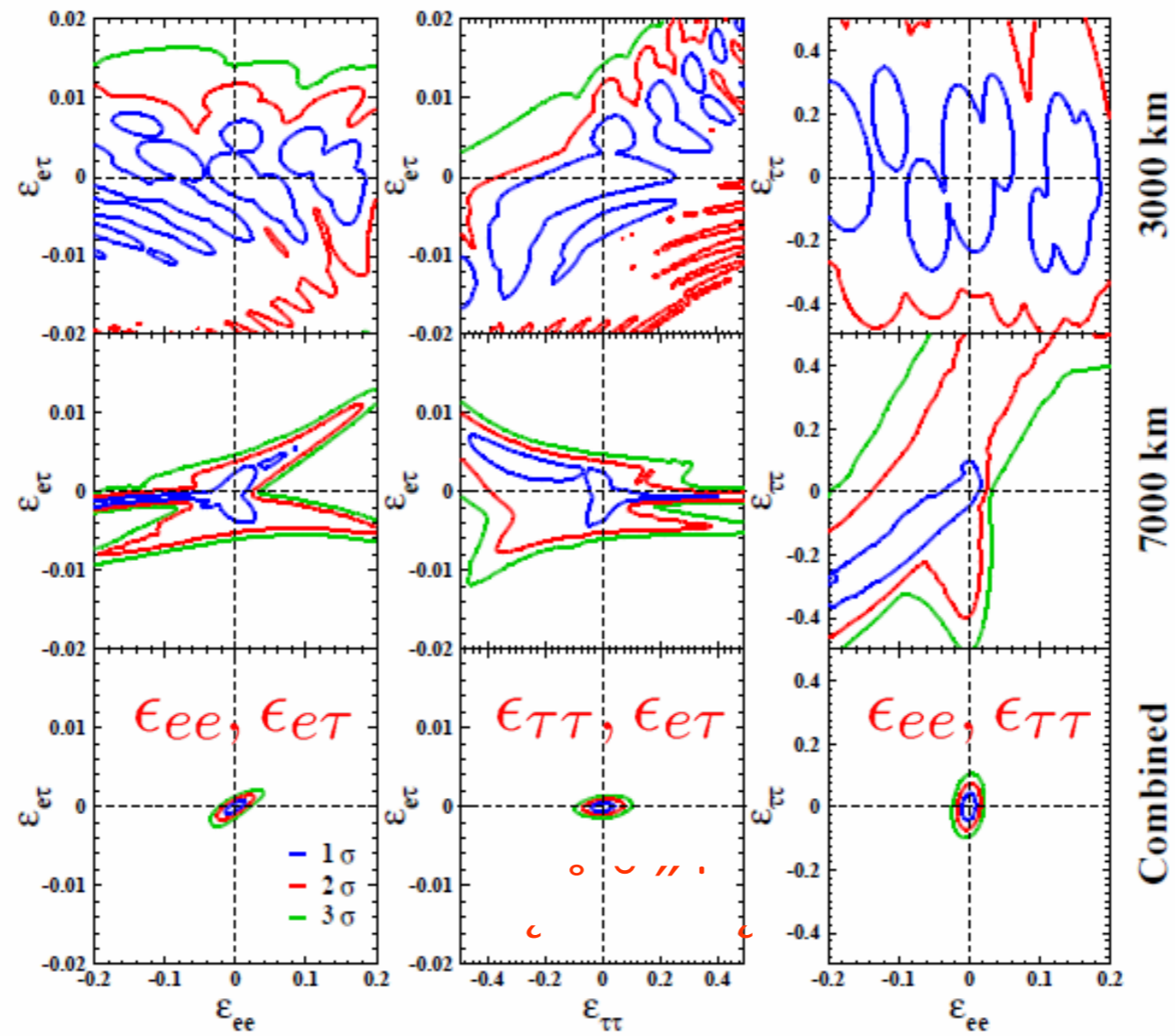
our setting

2 detector setting



$\sin^2 2\theta_{13}$ and δ marginalized

Input: $\delta = \pi/4$ $\sin^2 2\theta_{13} = 0.001$
 $\epsilon_{ee} = \epsilon_{e\tau} = \epsilon_{\tau\tau} = 0$



Neutrino factory with two detectors at $L=3000\text{km}$ and 7000km

1. solves the problem that the presence of NSI confuses the precision measurement of θ_{13} and δ
2. is powerful enough to probe into extremely small values of the NSI parameters.

$$\epsilon_{e\tau} \sim O(10^{-3})$$

$$\epsilon_{e\mu} \sim O(10^{-4})$$

(2 σ)

Conclusions:

- Neutrino Physics desperately needs to go beyond Megawatt traditional neutrino beams and Megaton water Cerenkov detectors: Neutrino Factory is an excellent possibility.
- For large $\sin^2 2\theta_{13}$ ($\geq 0.003-0.01$ say) the low energy option could provide precision measurements of the mixings to give meaningful tests to various sum rules coming from models and also explore the possibility of new physics as sub-leading effects.
- For smaller values of $\sin^2 2\theta_{13}$ the higher energy option provides unprecedented sensitivity to small values $\sin^2 2\theta_{13}$ and has the capability to untangle neutrino mixings from other new physics. (eg NSI Z’).